

NO. 165 ARIZONA-NASA ATLAS OF THE INFRARED SOLAR SPECTRUM,  
REPORT IX

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ABSTRACT

In this paper we give the solar spectrum  $\lambda\lambda$  25583–30920 Å as obtained from the NASA CV-990 Jet. A laboratory spectrum of the 2.7  $\mu$  H<sub>2</sub>O bands is included; Courtoy's laboratory spectrum of the 2.7  $\mu$  CO<sub>2</sub> bands is given.

In the summer of 1968 two LPL spectrometers recorded the solar spectrum from the NASA CV 990 Jet at high altitude. This paper concludes the preliminary reports on the solar spectrum as obtained with the LPL 4-m spectrometer. Previous reports were published as *Comm. LPL* Nos. 123, 124, 160, 161, 163, 164. The present report gives the solar spectrum  $\lambda\lambda$  25583–30920 Å; in a great part of the interval duplicate runs were available. The observing data are listed in Table I, pp. 142–43.

The wavelength scale is based on the wavelengths of water vapor lines as given by Plyler and Tidwell (1957) and by Gates *et al.* (1964); the wavelengths of CO<sub>2</sub> lines as given by Courtoy (1959); the wavelengths of N<sub>2</sub>O as given by Tidwell, Plyler and Benedict (1960). For the conversion of wavenumbers to wavelengths, Coleman's *Table of Wavenumbers* (NBS 1960) was used. Inaccuracy in the wavelength scale is caused by a periodic and a small random error in the dispersion, and by small incon-

sistencies in the above mentioned sources of wavelengths. The wavelength scale of Figs. 18–20 (Charts 67–69) had to be interpolated between the few water vapor absorptions, of which the wavelengths were taken from Gates *et al.* (1964).<sup>\*</sup> The positions of solar CO lines were calculated from the constants given by Goldberg and Müller (1953). The CO lines are indicated with an asterisk above the spectral traces. Atmospheric absorptions by H<sub>2</sub>O, N<sub>2</sub>O, and CO<sub>2</sub> are indicated by a dot, the symbol  $\phi$ , and a vertical line, respectively, all above the spectral trace. We tried to separate the several CO<sub>2</sub> bands in this region by placing the vertical lines at different levels and giving them a different appearance. In doing this we consulted Courtoy's spectrum of CO<sub>2</sub>, herein reproduced in Figs. 29 and 30.

Several runs of the water vapor spectrum at 2.7  $\mu$  were made, with different amounts of gas. In Figs. 21–28 we reproduce a spectrum with medium-strong absorptions. It shows the weak lines while the strong lines are saturated and their fine structure is not visible, contrary to the solar spectra. Some unidentified lines in the solar spectra may still be telluric.

<sup>\*</sup>Identification of solar lines on these charts were obtained from Migeotte, *et al.* (1956).

The solar spectra were obtained in the NASA CV-990 by Messrs. Kuiper and Cruikshank. The derivation of the wavelength scale and the identifications were all performed by Mr. Bijl, who also obtained the laboratory spectra of the  $2.7 \mu$   $H_2O$  bands and prepared the charts for publication.

*Acknowledgments* — We wish to thank Messrs. J. Percy, B. McClendon, A. Thomson and Rev. G. Sill of LPL and Mr. D. Olsen of NASA-Ames for their assistance during the flights. Mr. D. C. Benner constructed the wavelength scales for the laboratory spectra and assisted in the calculation of the CO line positions. Mrs. A. P. Agnieray and Mr. S. M. Larson assisted in the preparation of the figures. This research was supported by NASA through Grant NsG 161-61 and the University of Arizona Institutional Grant NGR-03-002-091.

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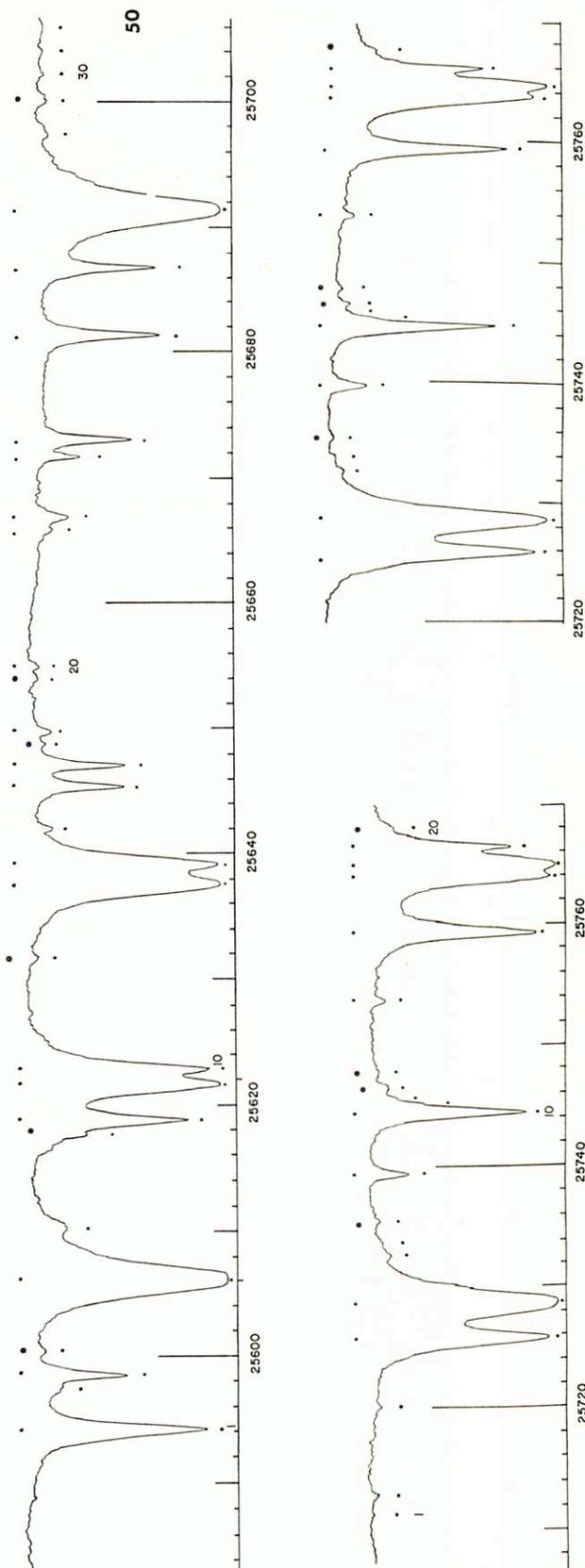


Fig. 1 Solar Spectrum  $\lambda$  25583–25770, in two strips (cf. Table 1).

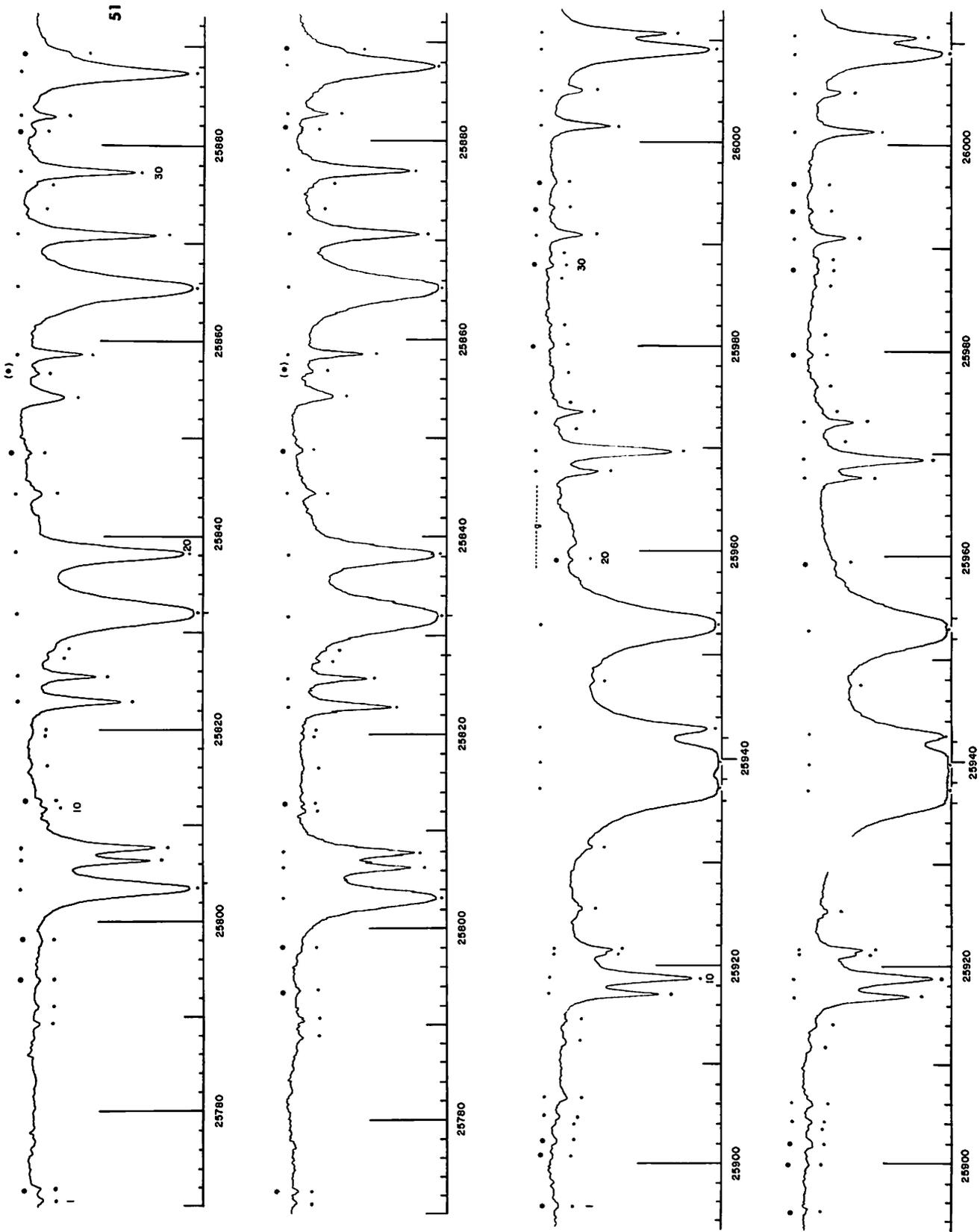


Fig. 2 Solar Spectrum  $\lambda\lambda$  25770-26013, in four strips (cf. Table 1).

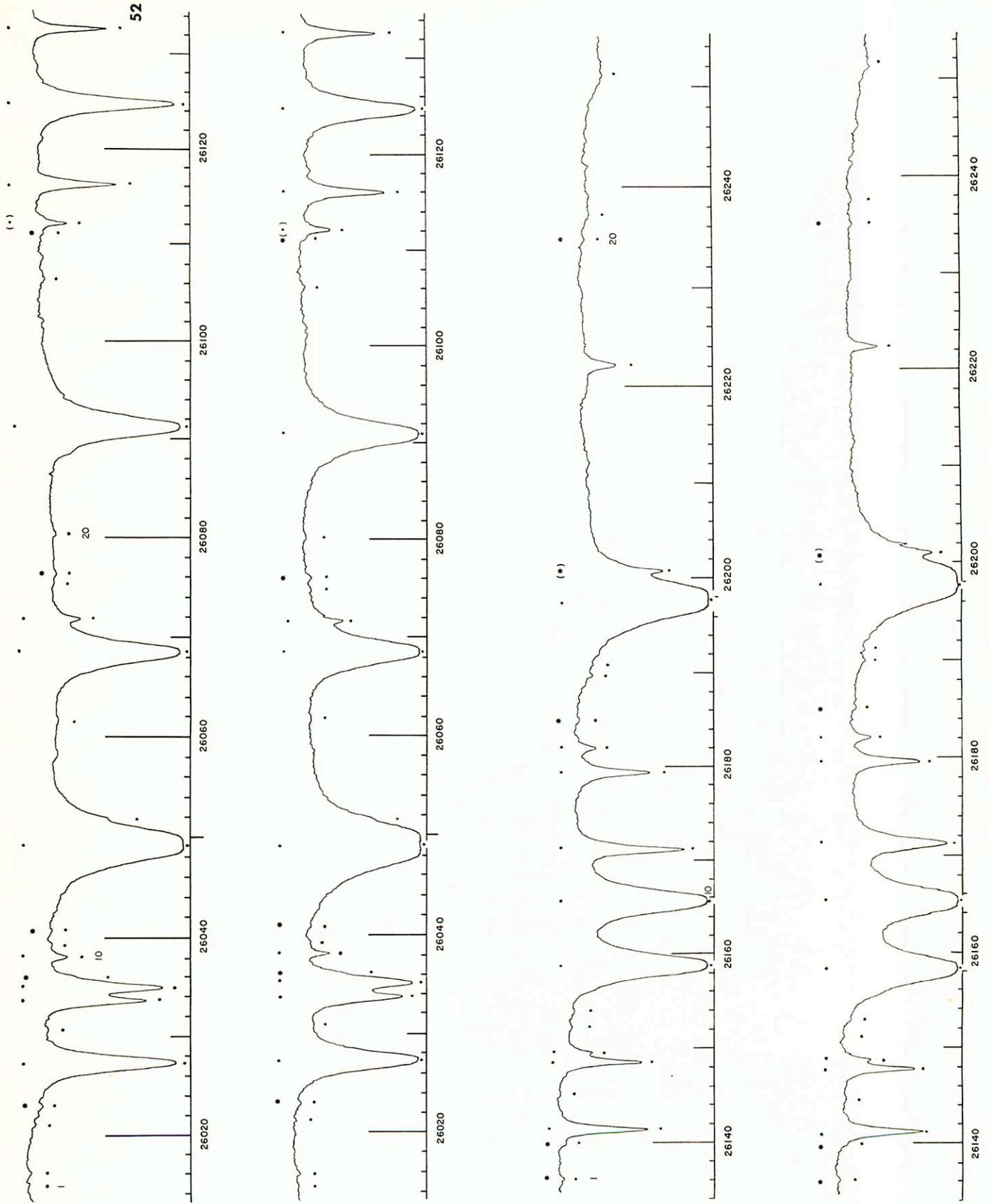


Fig. 3 Solar Spectrum  $\lambda\lambda$  26013-26255, in four strips (cf. Table 1).

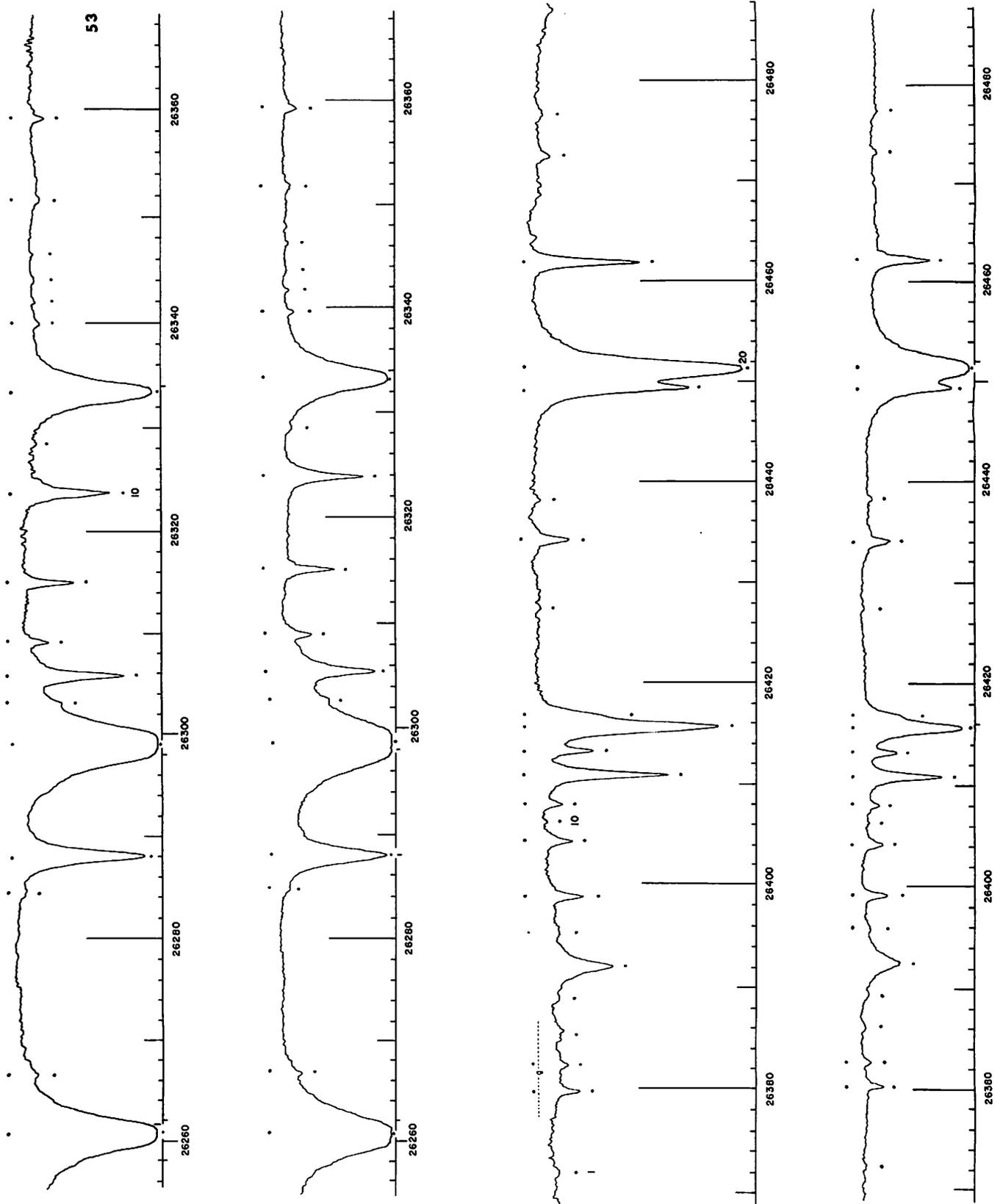


Fig. 4 Solar Spectrum  $\lambda$  26255-26488, in four strips (cf. Table 1).

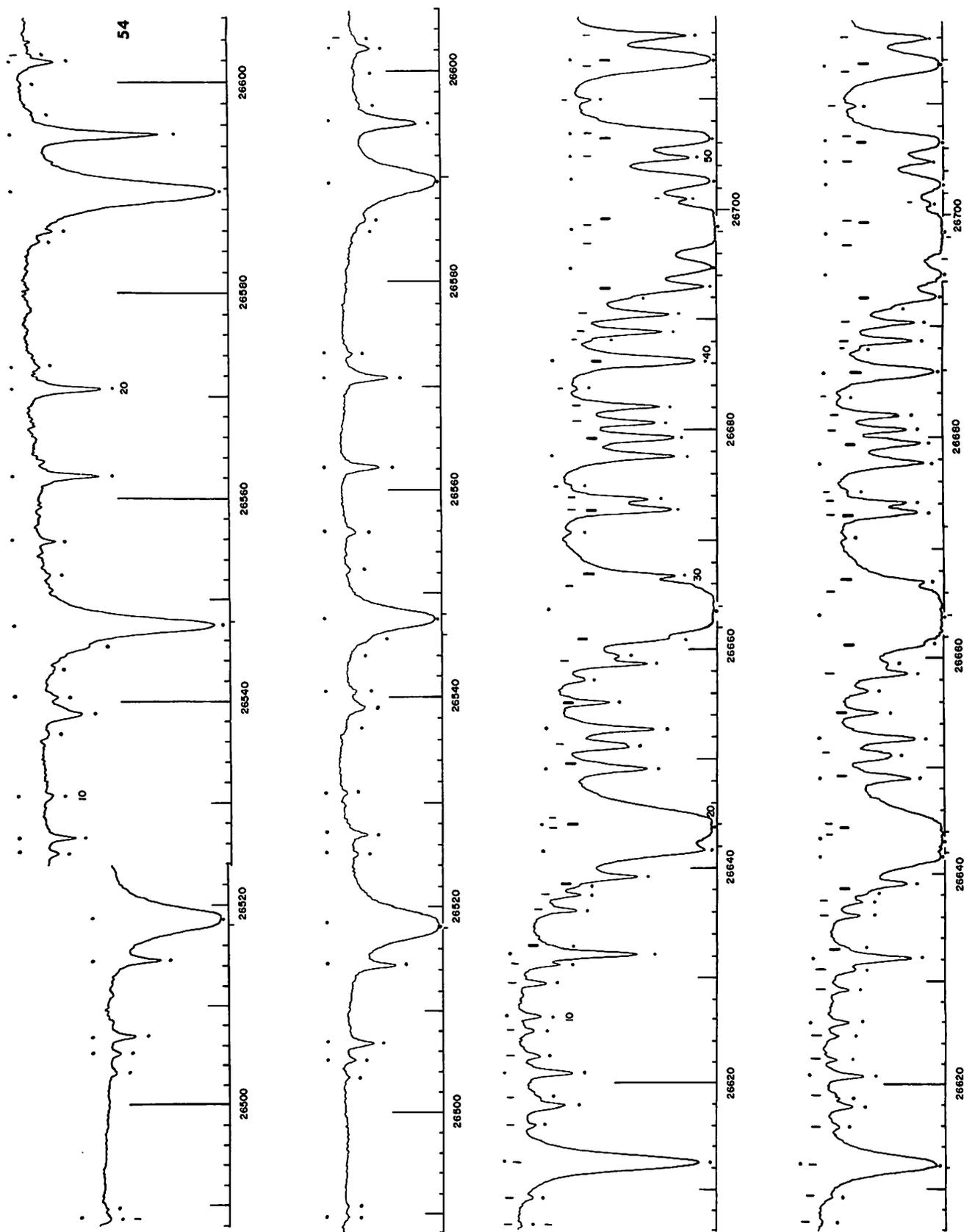


Fig. 5 Solar Spectrum  $\lambda$  26488–26718, in four strips (cf. Table 1).

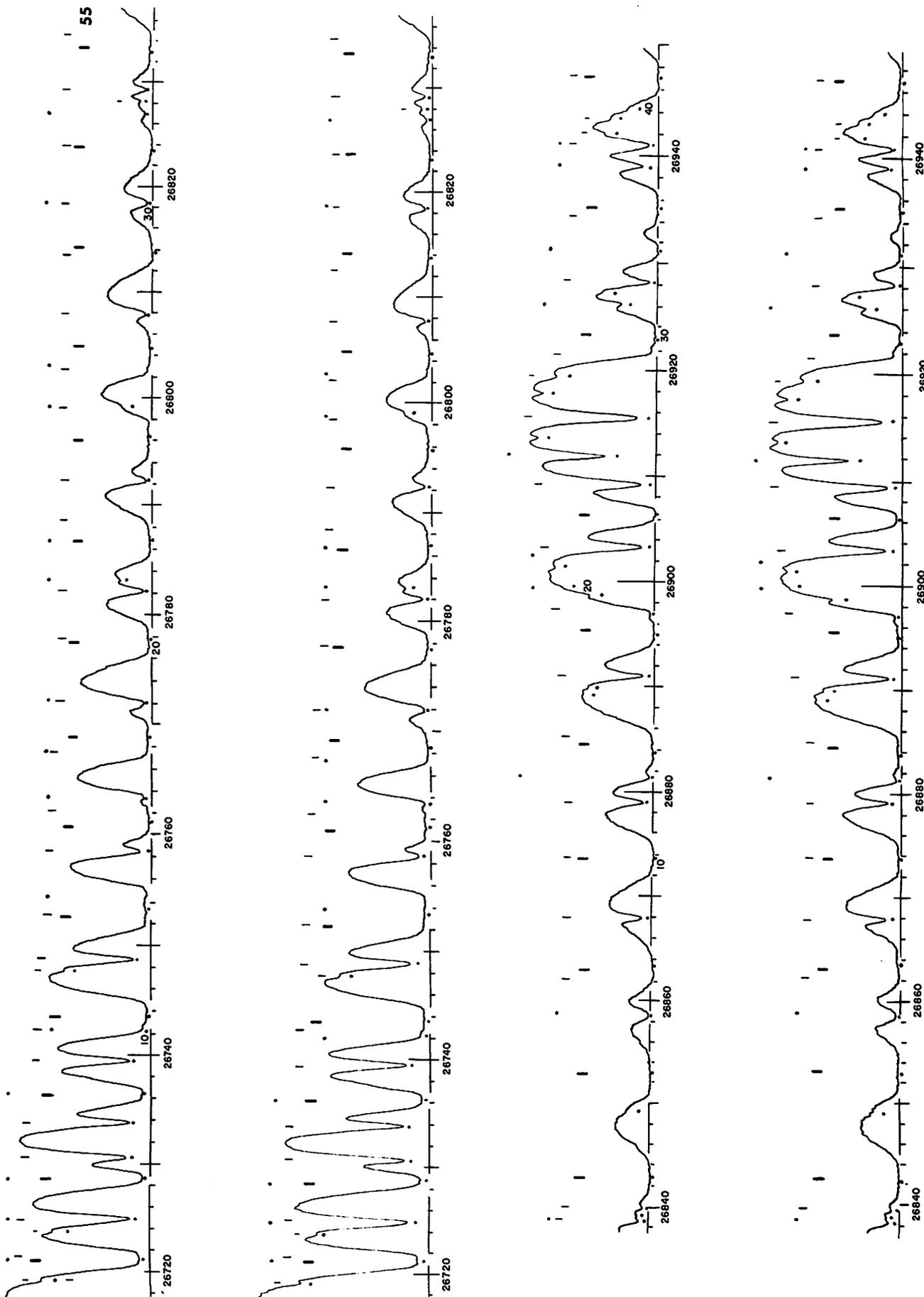


Fig. 6 Solar Spectrum  $\lambda\lambda$  26718-26950, in four strips (cf. Table 1).

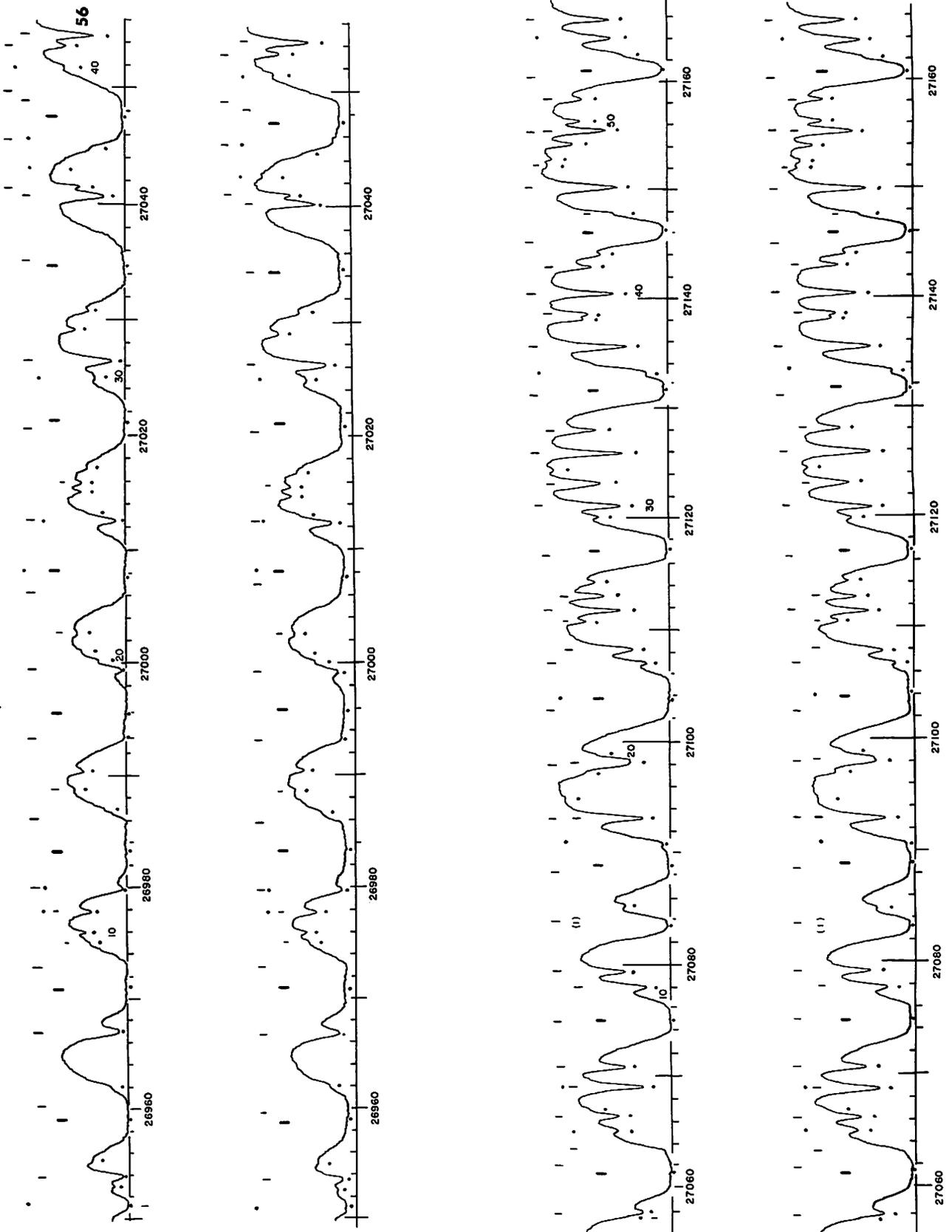


Fig. 7 Solar Spectrum  $\lambda\lambda$  26950–27167, in four strips (cf. Table I).

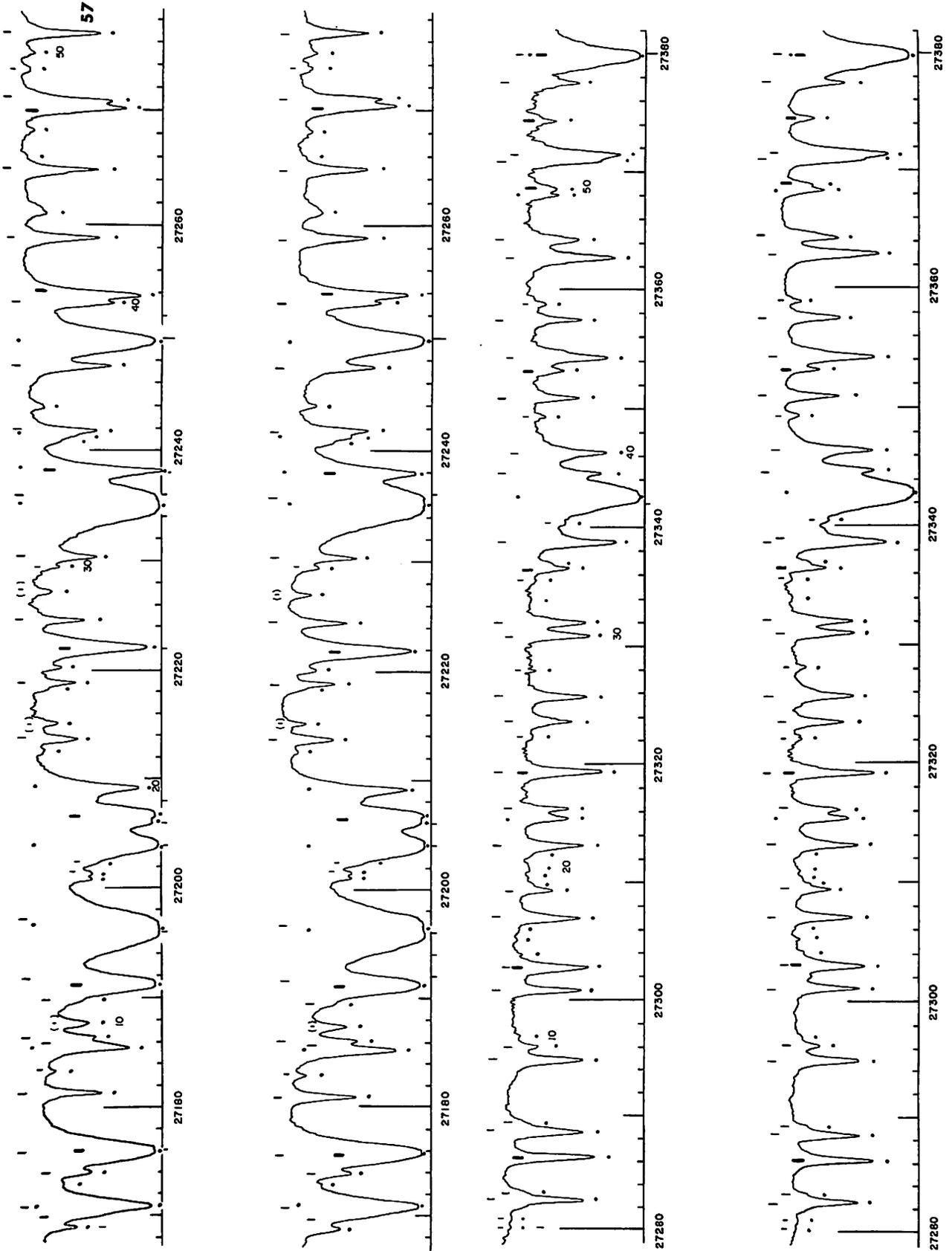


Fig. 8 Solar Spectrum  $\lambda\lambda$  27167-27382, in four strips (cf. Table 1).

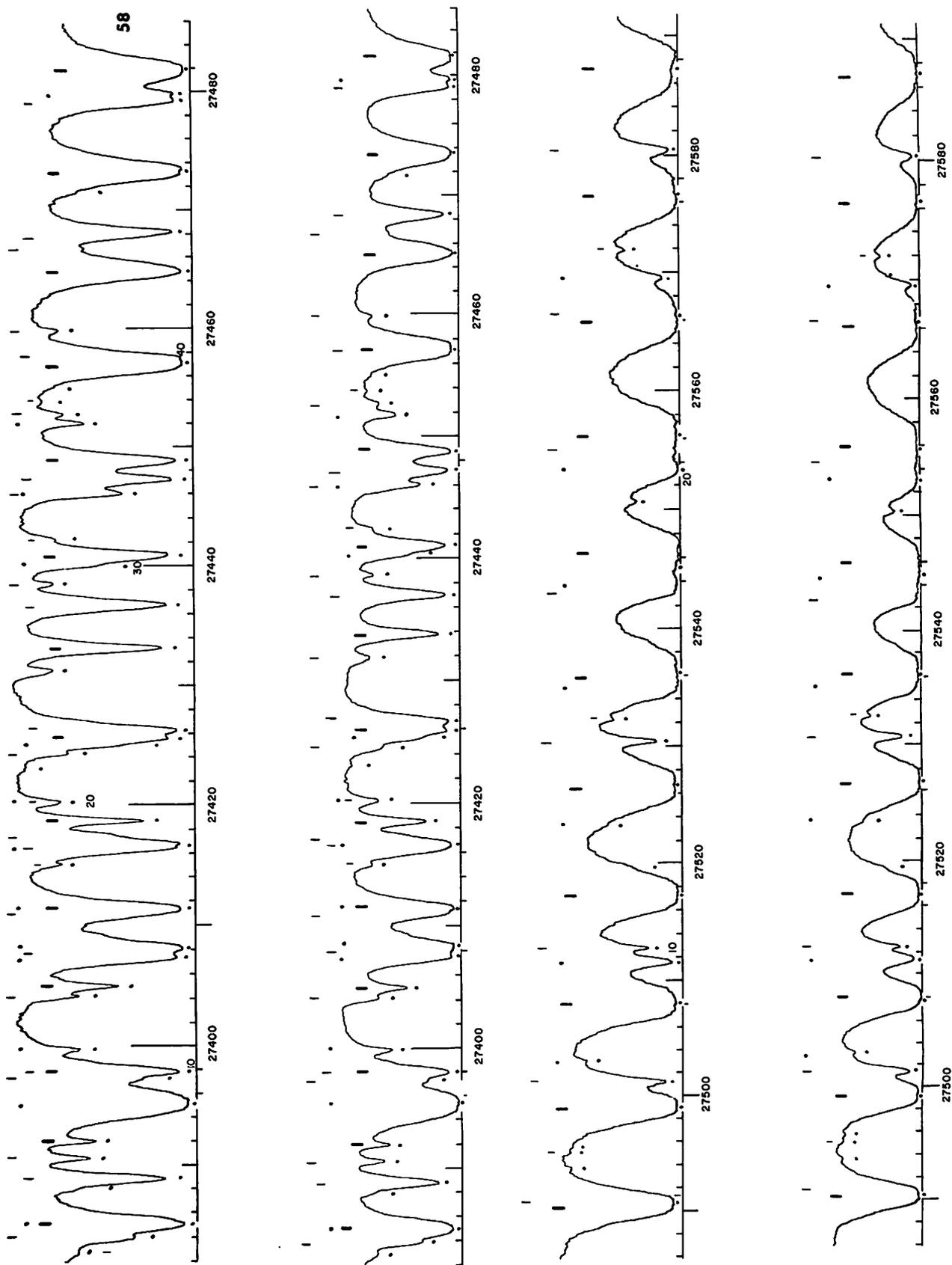


Fig. 9 Solar Spectrum  $\lambda\lambda$  27382-27592, in four strips (cf. Table 1).

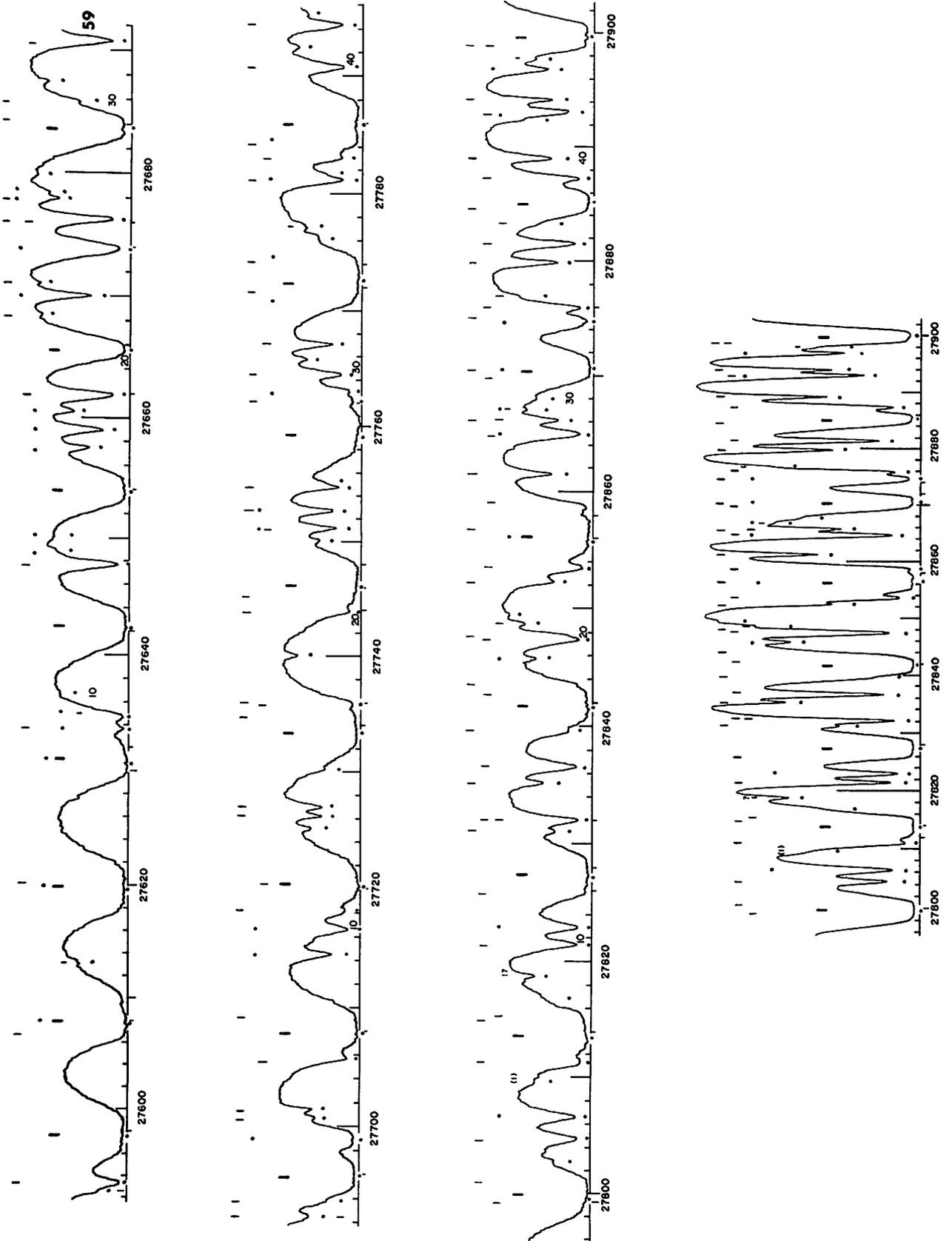


Fig. 10 Solar Spectrum  $\lambda$  27592-27903, in four strips (cf. Table 1).

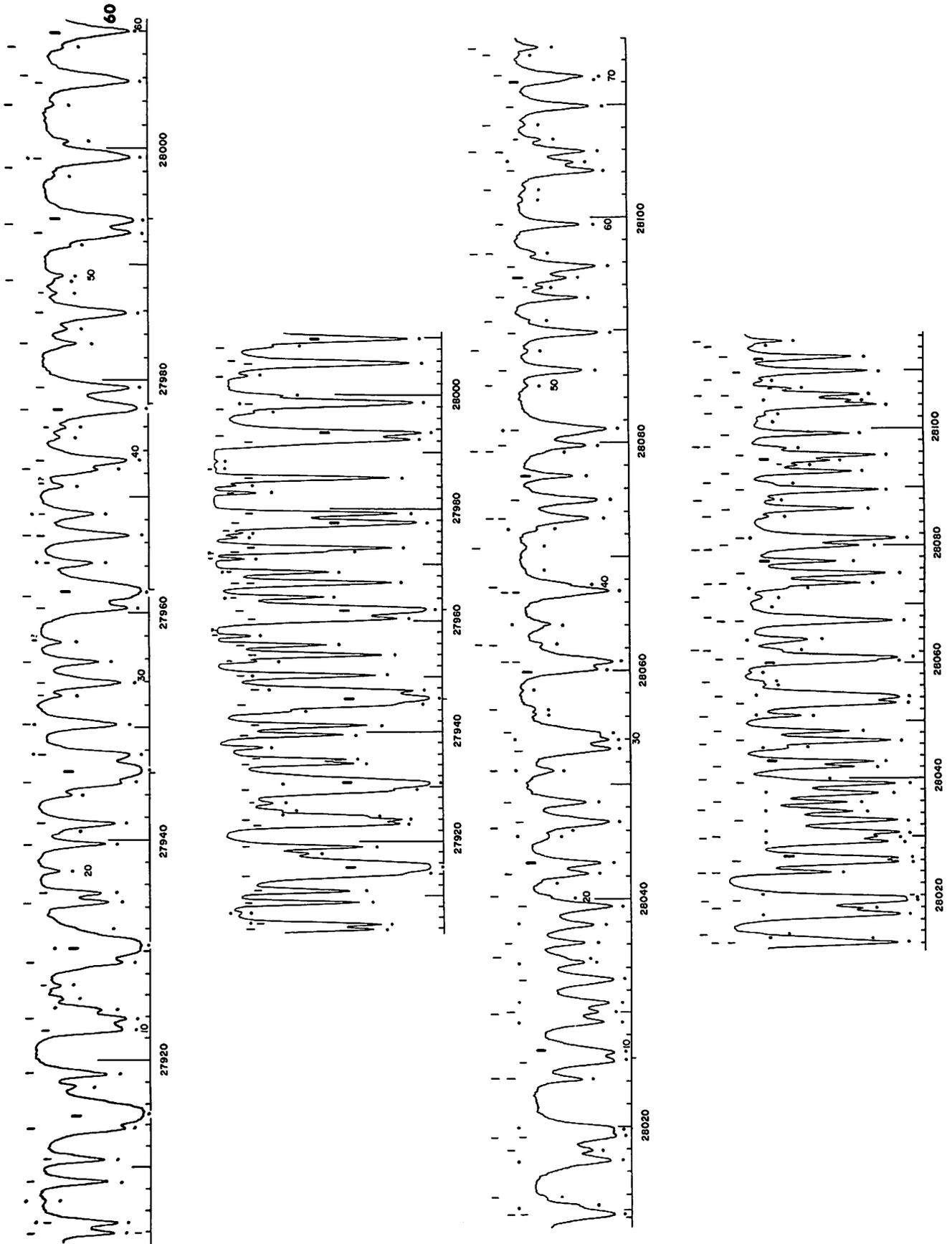


Fig. 11 Solar Spectrum  $\lambda\lambda$  27903-28116, in four strips (cf. Table 1).

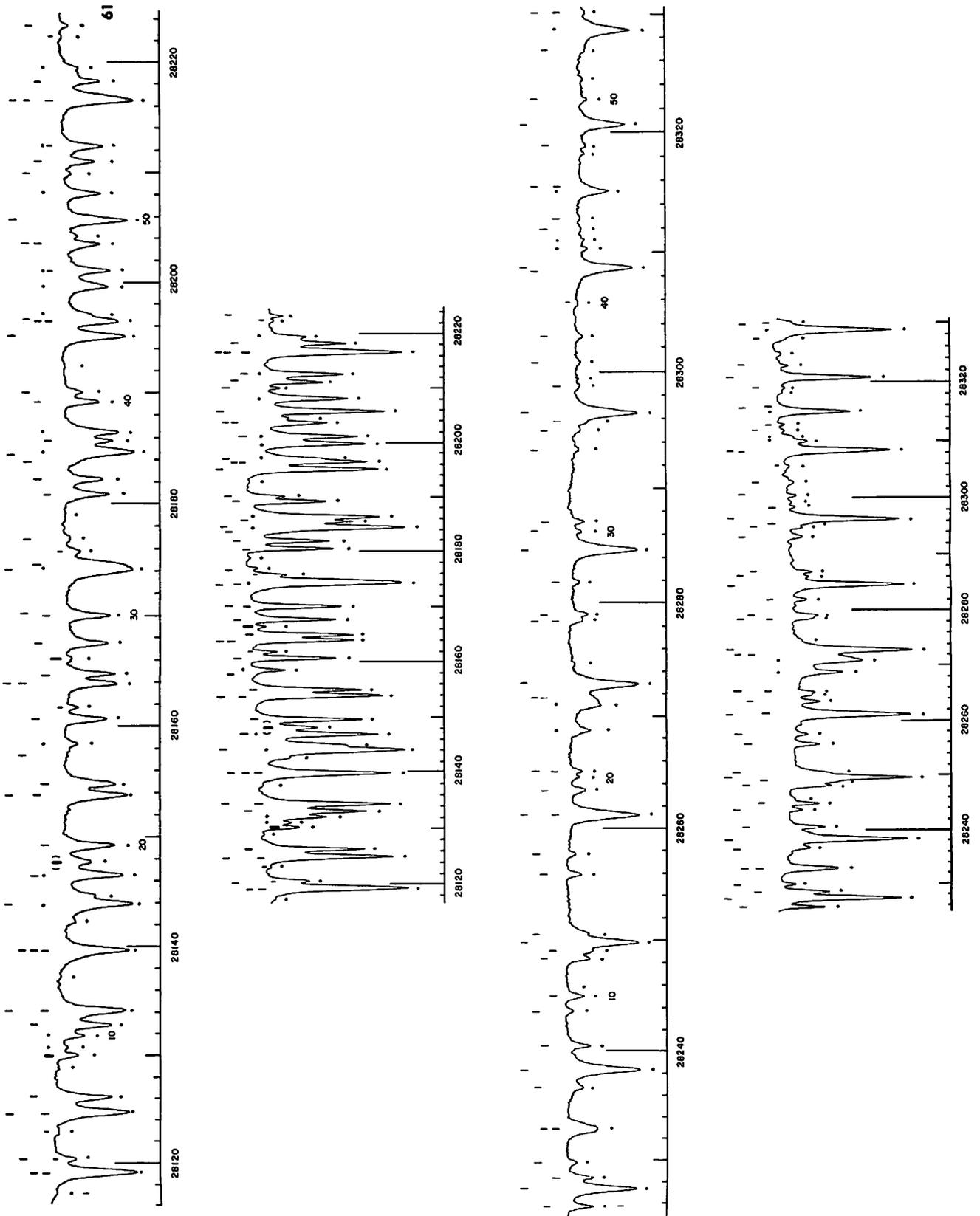


Fig. 12 Solar Spectrum  $\lambda$  28116–28331, in four strips (cf. Table 1).

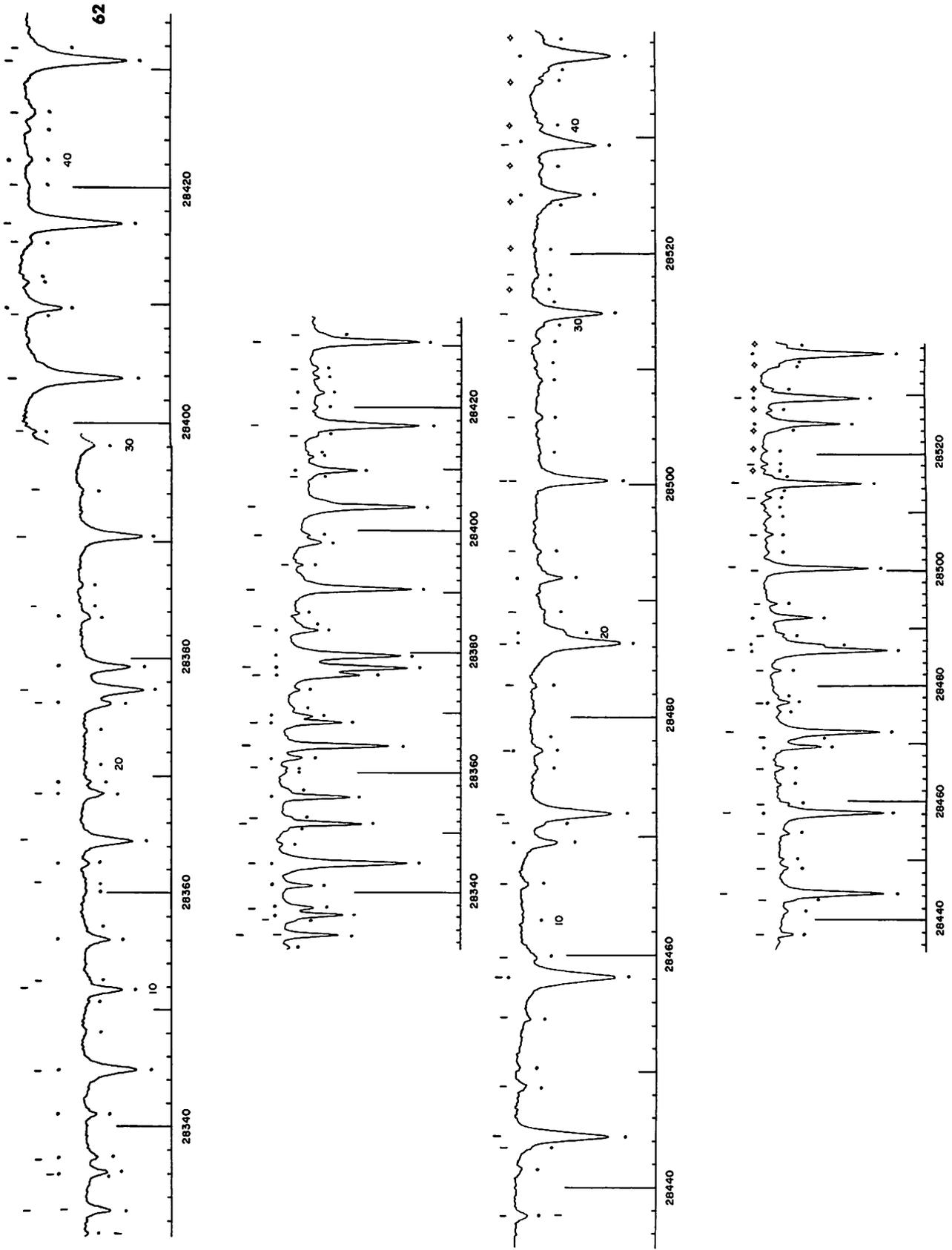


Fig. 13 Solar Spectrum  $\lambda$  28331-28539, in four strips (cf. Table 1).

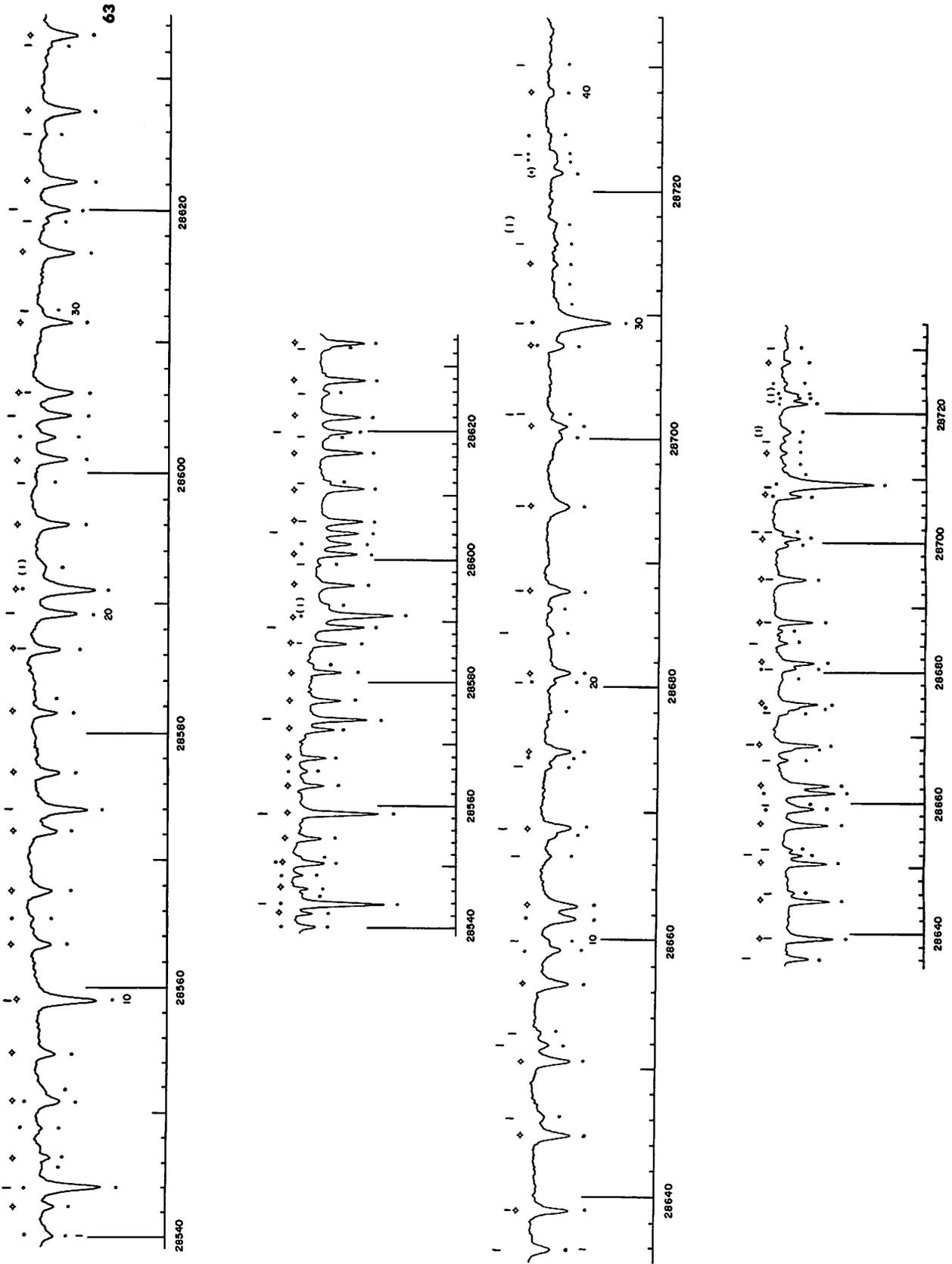


Fig. 14 Solar Spectrum  $\lambda\lambda$  28539-28734, in four strips (cf. Table 1).

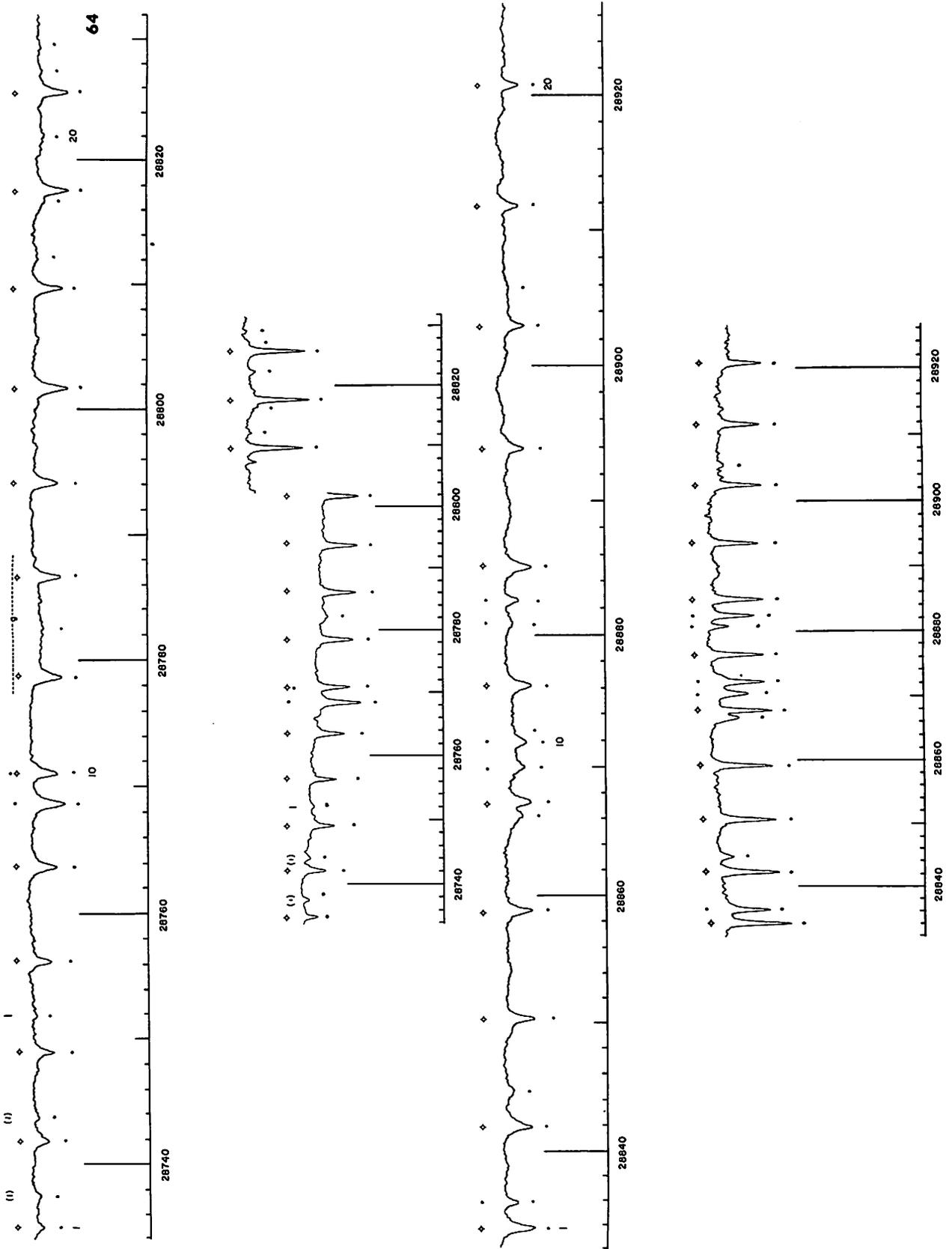


Fig. 15 Solar Spectrum  $\lambda\lambda$  28734–28927, in four strips (cf. Table 1).

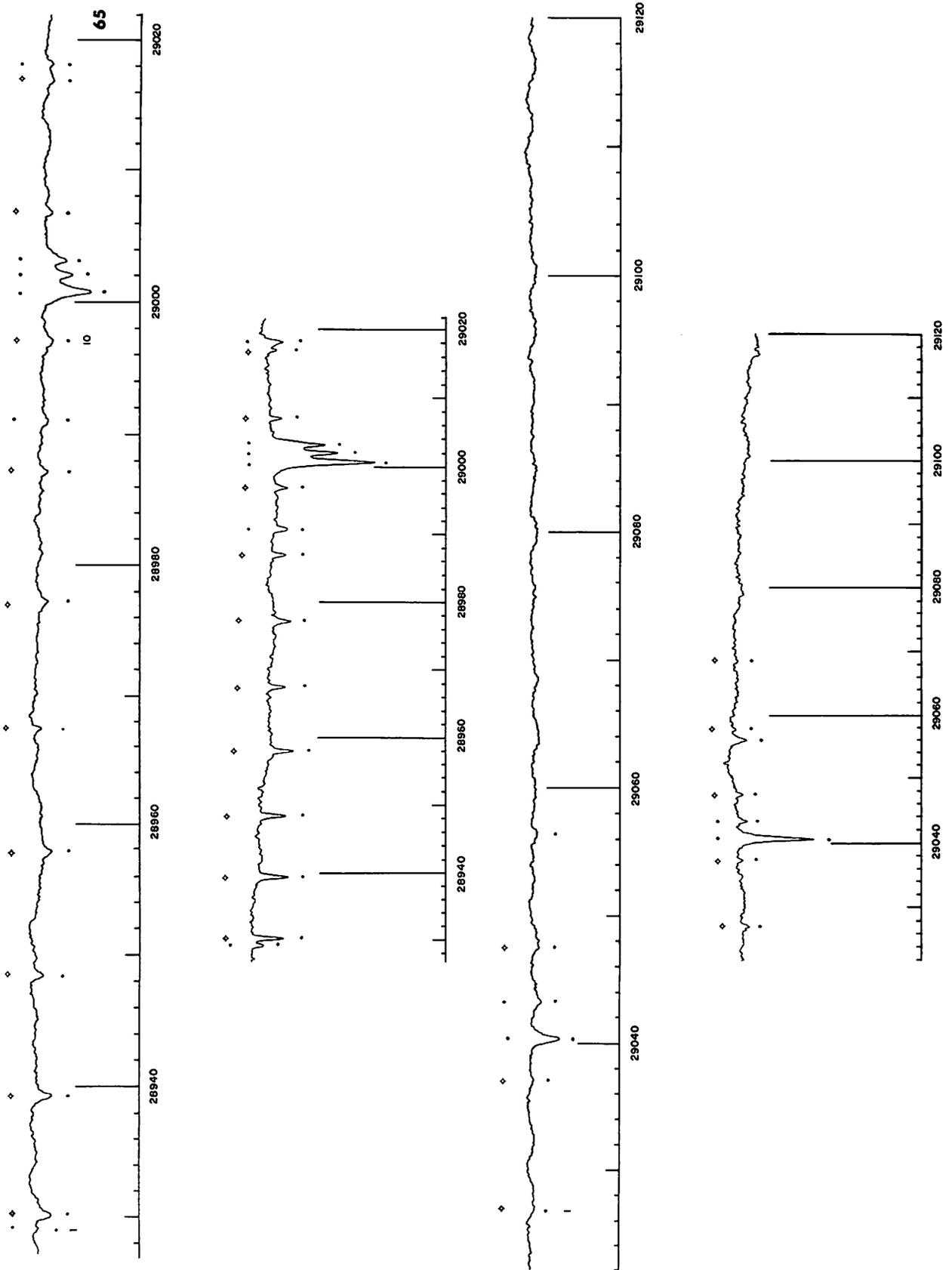


Fig. 16 Solar Spectrum  $\lambda$  28927-29120, in four strips (cf. Table 1).

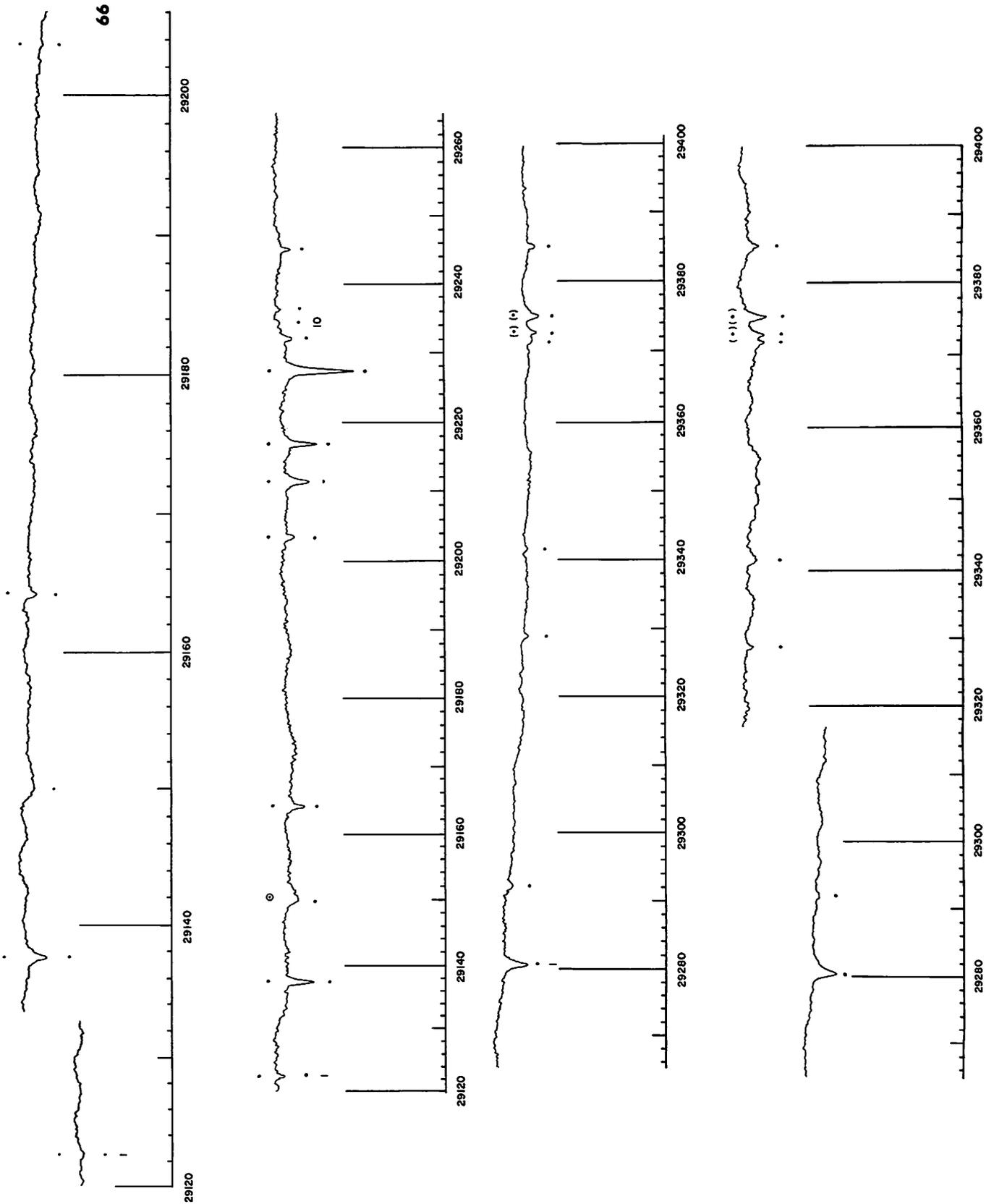


Fig. 17 Solar Spectrum  $\lambda\lambda$  29120-29400, in four strips (cf. Table 1).

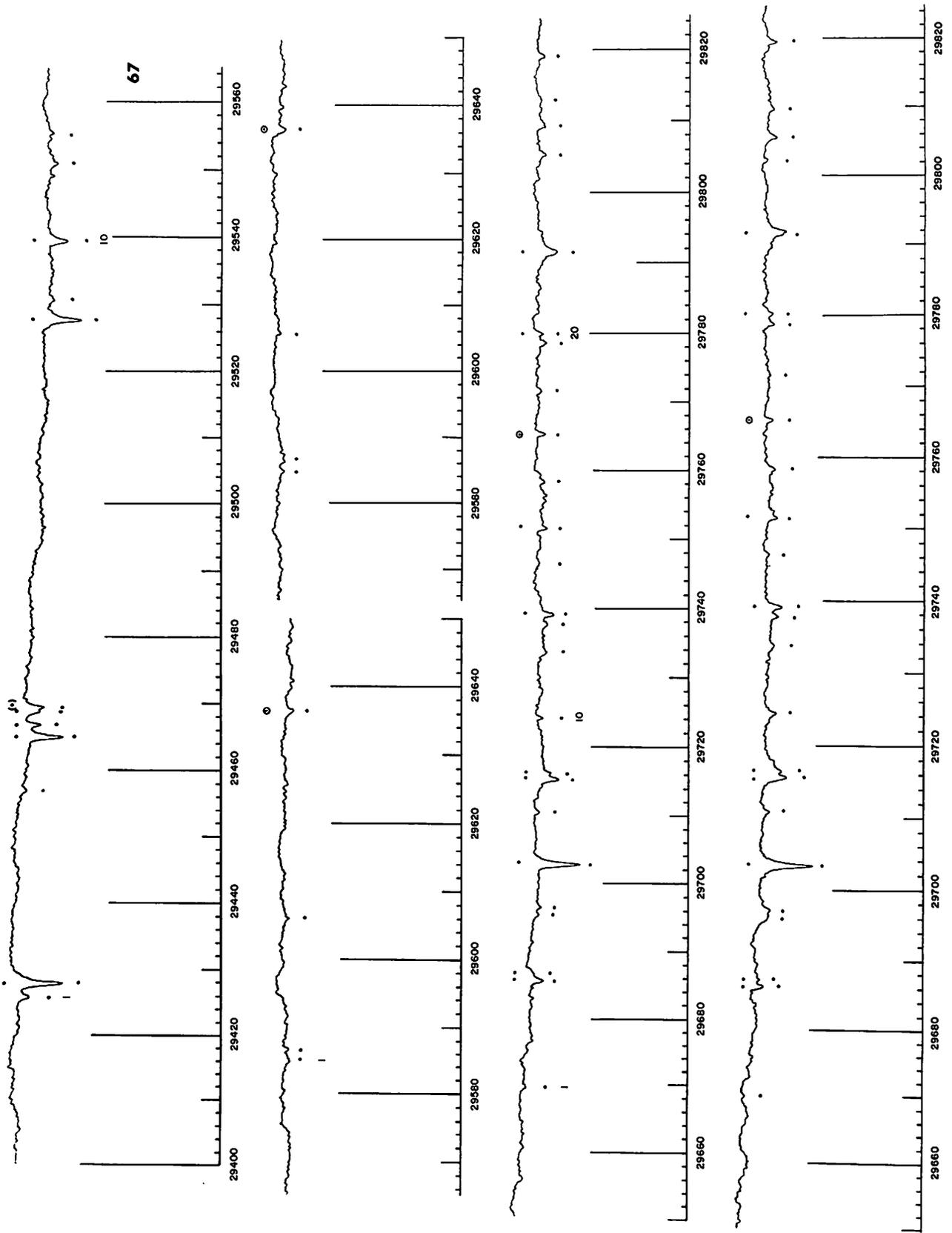


Fig. 18 Solar Spectrum  $\lambda$  29400-29825, in four strips (cf. Table 1).

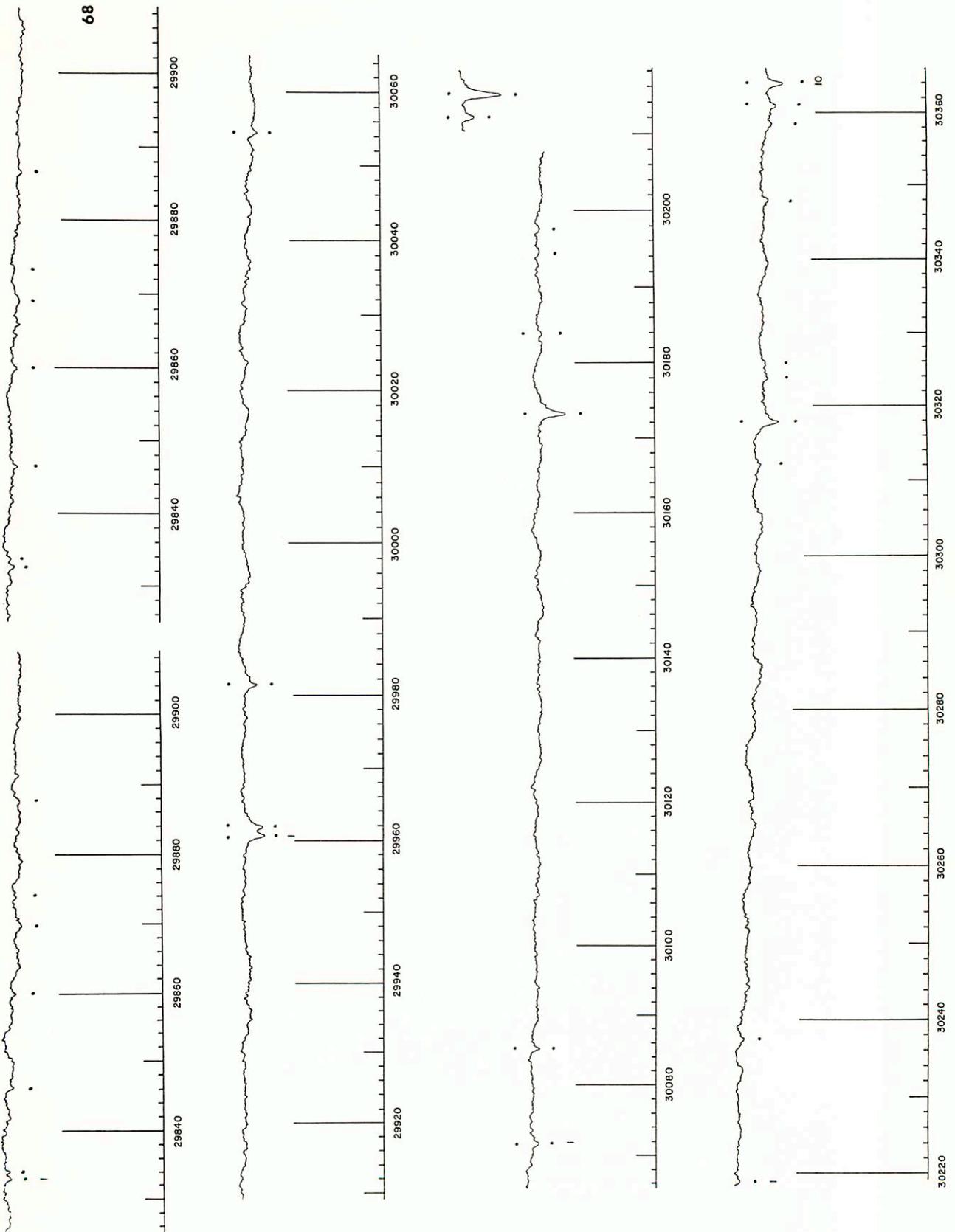


Fig. 19 Solar Spectrum  $\lambda\lambda$  29825-30366, in four strips (cf. Table 1).

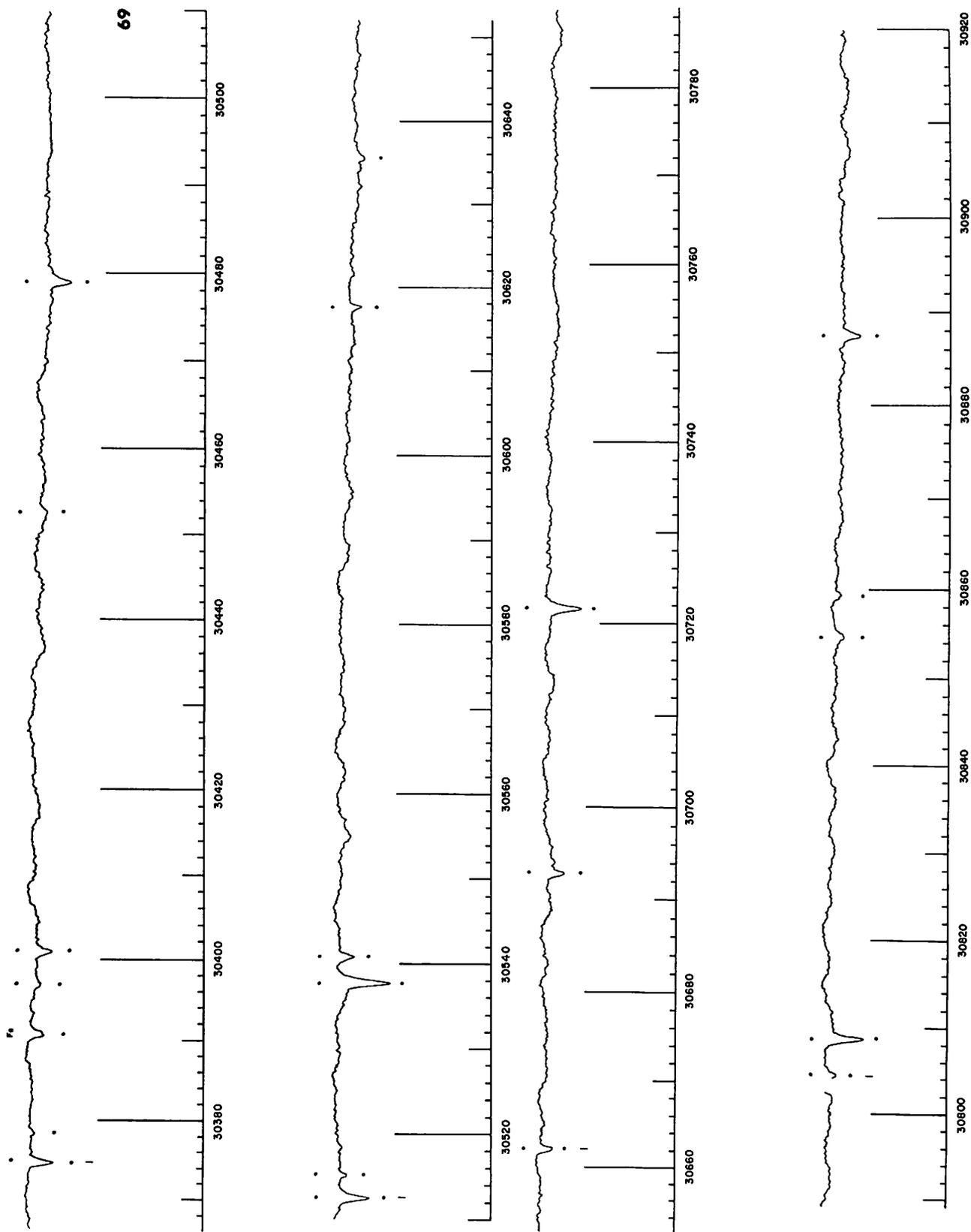


Fig. 20 Solar Spectrum  $\lambda\lambda$  30366-30920, in four strips (cf. Table 1).

TABLE I  
 SOLAR SPECTRUM RECORDS, 4-IN SPECTROMETER, NASA CV-900 JET  
 GRATING 600 l/mm; GRATING BLAZE, JULY 19: 1.6  $\mu$ , JULY 30, AUG. 1: 2.5  $\mu$   
 FILTER 1.8  $\mu$ ; DETECTOR 0.10 mm

FIG.	CHART	1968 DATE	UT	ALT. (FT.)	OUTSIDE TEMP. (°C)	CABIN ALT. (FT.)	GAIN	SLIT (mm)	$\tau$ (SEC.)
1	50 c	July 19	19:04	39,000	-52	8500	5-6	0.18	0.12
	d	July 19/30	19:07/18:26	39,000/40,800	-52/-55	8500/8900	5-5	0.18/0.15	0.12
2	51 a	July 30	18:29	41,000	-55	8900	5-5	0.15	0.12
	b	July 19	19:09	39,000	-52	8500	5-6	0.18	0.12
	c	July 30	18:32	41,100	-55	8900	5-5	0.15	0.12
	d	July 19	19:12	39,000	-52	8500	5-6	0.18	0.12
3	52 a	July 30	18:35	41,200	-55	8900	5-5	0.15	0.12
	b	July 19	19:16	39,000	-52	8500	5-6	0.18	0.12
	c	July 30	18:38	41,300	-55	8900	5-5	0.15	0.12
	d	July 19	19:19	39,000	-52	8500	5-6	0.18	0.12
4	53 a	July 30	18:41	41,400	-55	8900	5-5	0.15	0.12
	b	July 19	19:22	39,000	-52	8500	5-6	0.18	0.12
	c	July 30	19:25	41,500	-59	8900	5-6	0.15	0.12
	d	July 19	19:25	39,000	-52	8500	5-6	0.18	0.12
5	54 a	July 30	18:47/19:28	41,500	-55/-59	8900	5-5/5-6	0.15	0.12
	b	July 19	19:29	39,000	-52	8500	5-6	0.18	0.12
	c	July 30	19:31	41,500	-59	8900	5-6	0.15	0.12
	d	July 19	19:37	39,000	-52	8500	6-2	0.18	0.12
6	55 a	July 30	19:34	61,500	-59	8900	5-6	0.15	0.12
	b	July 30	18:55	41,700	-55	8900	5-6	0.15	0.12
	c	July 30	19:38	41,500	-59	8900	5-6	0.15	0.12
	d	July 30	18:58	41,600	-56	8900	5-6	0.15	0.12
7	56 a	July 30	19:42	41,500	-60	8900	5-6	0.15	0.12
	b	July 30	19:01	41,600	-57	8900	5-6	0.15	0.12
	c	July 30	19:45	41,500	-60	8900	5-6	0.15	0.12
	d	July 30	19:05	41,500	-58	8900	5-6	0.15	0.12
8	57 a	July 30	19:48	41,500	-60	8900	5-6	0.15	0.12
	b	July 30	19:08	41,500	-58	8900	5-6	0.15	0.12
	c	July 30	19:51	41,500	-60	8900	5-6	0.15	0.12
	d	July 30	19:11	41,500	-59	8900	5-6	0.15	0.12
9	58 a	July 30	19:57	41,500	-60	8900	6-2	0.15	0.12
	b	July 30	19:14	41,500	-59	8900	5-6	0.15	0.12
	c	July 30	20:00	41,500	-59	8900	6-2	0.15	0.12
	d	July 30	19:17	41,500	-59	8900	5-6	0.15	0.12

10	59 a	27592-27692	July 30	20:03	41,500	-59	8900	6-2	0.15	0.12
	b	27692-27796	July 30	20:06	41,500	-59	8900	6-2	0.15	0.12
	c	27796-27903	July 30	20:10	41,500	-59	8900	6-2	0.14	0.12
	d*	27796-27903	Aug. 1	18:19	39,000	-54	9300	6-1	0.30	0.12
11	60 a	27903-28011	July 30	20:13	41,500	-59	8900	6-2	0.14	0.12
	b*	27903-28011	Aug. 1	18:20	39,000	-54	9300	6-1	0.30/0.24	0.12
	c	28011-28116	July 30	20:16	41,500	-58	8900	6-2	0.14	0.12
	d*	28011-28116	Aug. 1	18:22	39,200	-54	9300	6-1	0.24	0.12
12	61 a	28116-28225	July 30	20:20	41,500	-58	8900	6-2	0.14	0.12
	b*	28116-28225	Aug. 1	18:24	39,400	-56	9300	6-1	0.24	0.12
	c	28225-28331	July 30	20:23	41,500	-58	8900	6-2	0.14	0.12
	d*	28225-28331	Aug. 1	18:26	39,600	-54	9300	6-1	0.24	0.12
13	62 a	28331-28435	July 30	20:26	41,500	-58	8900	6-2/6-3	0.14	0.12/0.24
	b*	28331-28435	Aug. 1	18:27	39,800	-54	9300	6-1	0.24	0.12
	c	28435-28539	July 30	20:30	41,500	-58	8900	6-3	0.14	0.24
	d*	28435-28539	Aug. 1	18:29	40,000	-54	9300	6-1	0.24	0.12
14	63 a	28539-28635	July 30	20:33	41,500	-58	8900	6-3	0.14	0.24
	b*	28539-28635	Aug. 1	18:31	40,200	-53	9300	6-1	0.24	0.12
	c	28635-28734	July 30	20:37	41,500	-58	8900	6-3	0.14	0.24
	d*	28635-28734	Aug. 1	18:32	40,400	-53	9300	6-1	0.24	0.12
15	64 a	28734-28832	July 30	20:40	41,500	-58	8900	6-3	0.24	0.24
	b*	28734-28832	Aug. 1	18:34	40,500	-53	9300	6-1/6-2	0.24	0.12
	c	28832-28927	July 30	20:43	41,500	-58	8900	6-3	0.14	0.24
	d*	28832-28927	Aug. 1	18:36	40,500	-53	9300	6-2	0.24	0.12
16	65 a	28927-29022	July 30	20:46	41,500	-58	8900	6-3	0.14	0.24
	b*	28927-29022	Aug. 1	18:37	40,500	-53	9300	6-2	0.24	0.12
	c	29022-29120	July 30	20:49	41,500	-58	8900	6-3	0.14	0.24
	d*	29022-29120	Aug. 1	18:39	40,500	-53	9300	6-2	0.24	0.12
17	66 a	29120-29206	July 30	20:52	41,500	-58	8900	6-3	0.14/0.20	0.24
	b*	29120-29265	Aug. 1	18:41	40,500	-53	9300	6-2	0.24	0.12
	c*	29265-29400	Aug. 1	18:44	40,300	-53	9300	6-2	0.24	0.12/0.24
	d*	29265-29400	Aug. 1	18:48	40,500	-53	9300	6-2/6-3	0.24	0.24
18	67 a*	29400-29565	Aug. 1	18:50	40,500	-53	9300	6-3	0.24	0.24
	b*	29565-29650	Aug. 1	18:53/19:21	40,800/41,500	-55/--58	9300	6-3	0.24	0.24
	c*	29650-29825	Aug. 1	18:55	41,200	-58	9300	6-3	0.24	0.24
	d*	29650-29825	Aug. 1	19:23	41,500	-58	9300	6-3	0.24	0.24
19	68 a*	29825-29909	Aug. 1	18:58/19:26	41,800/41,500	-59/--60	9300	6-3	0.24	0.24
	b*	29909-30065	Aug. 1	19:28	41,500	-60	9300	6-3	0.24	0.24
	c*	30065-30218	Aug. 1	19:31	41,500	-60	9300	6-3	0.24/0.38	0.24
	d*	30218-30366	Aug. 1	19:34	41,500	-60	9300	6-3	0.38	0.24
20	69 a*	30366-30510	Aug. 1	19:37	41,500	-60	9300	6-3	0.38	0.24
	b*	20510-30652	Aug. 1	19:40	41,500	-60	9300	6-3	0.38	0.24
	c*	30652-30789	Aug. 1	19:43	41,500	-60	9300	6-3	0.38	0.24
	d*	30789-30920	Aug. 1	19:47	41,500	-60	9300	6-3	0.38	0.24

\*Grating turned at double rate.

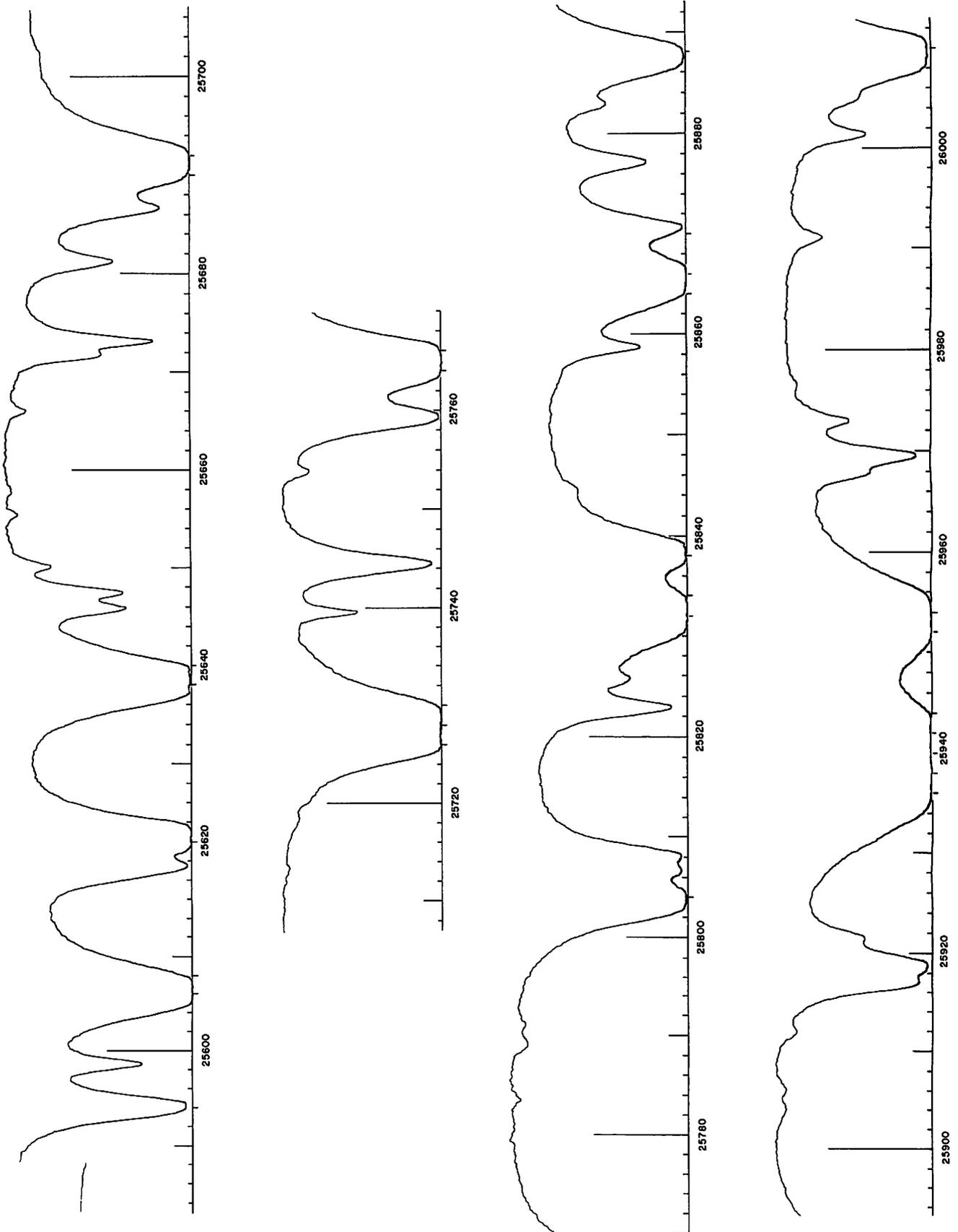


Fig. 21 Laboratory Spectrum of water vapor  $\lambda\lambda$  25583-26013, matching the solar spectrum Figs. 1, 2.

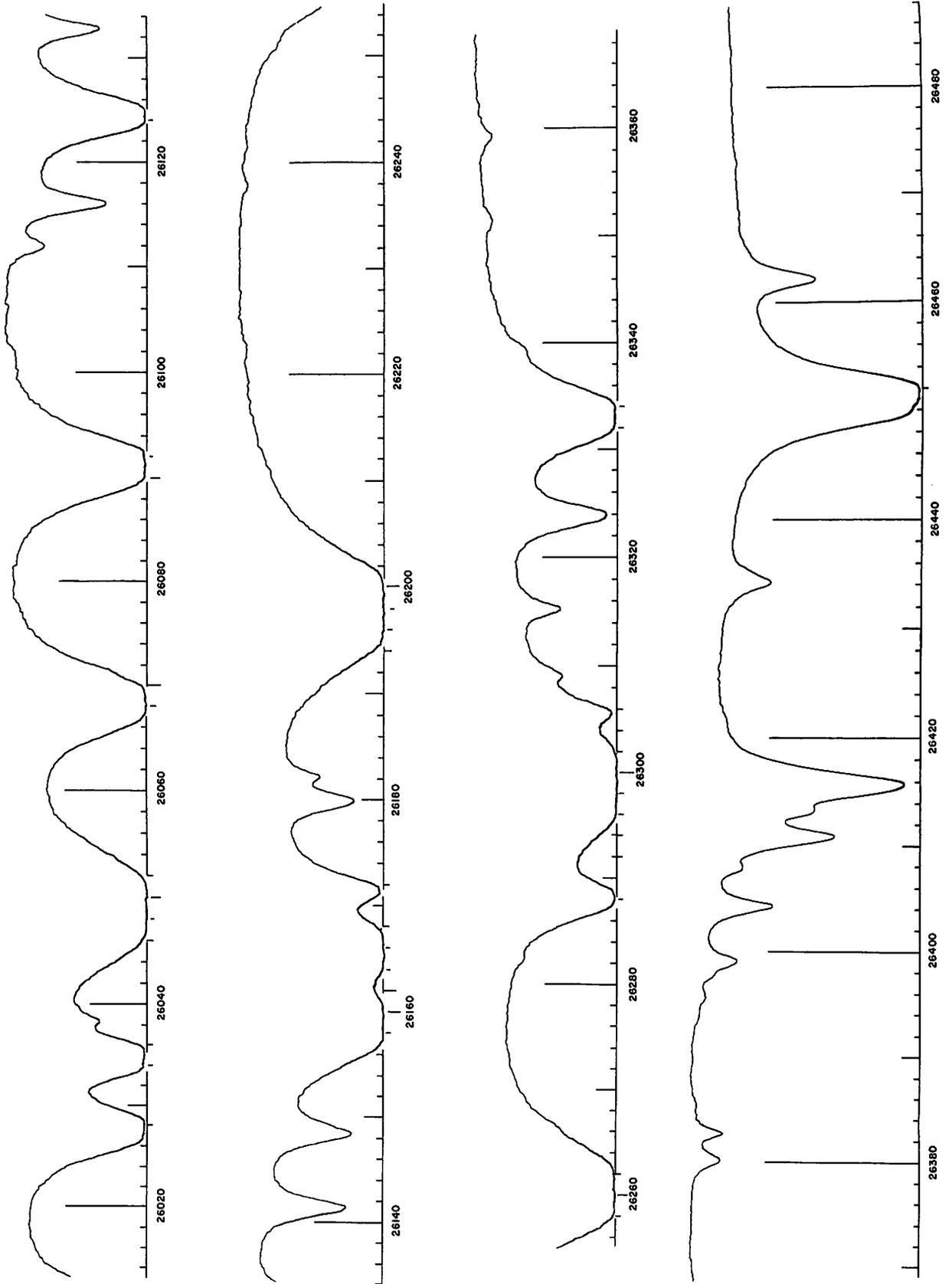


Fig. 22 Laboratory Spectrum of water vapor  $\lambda\lambda$  26013-26488, matching the solar spectrum Figs. 3, 4.

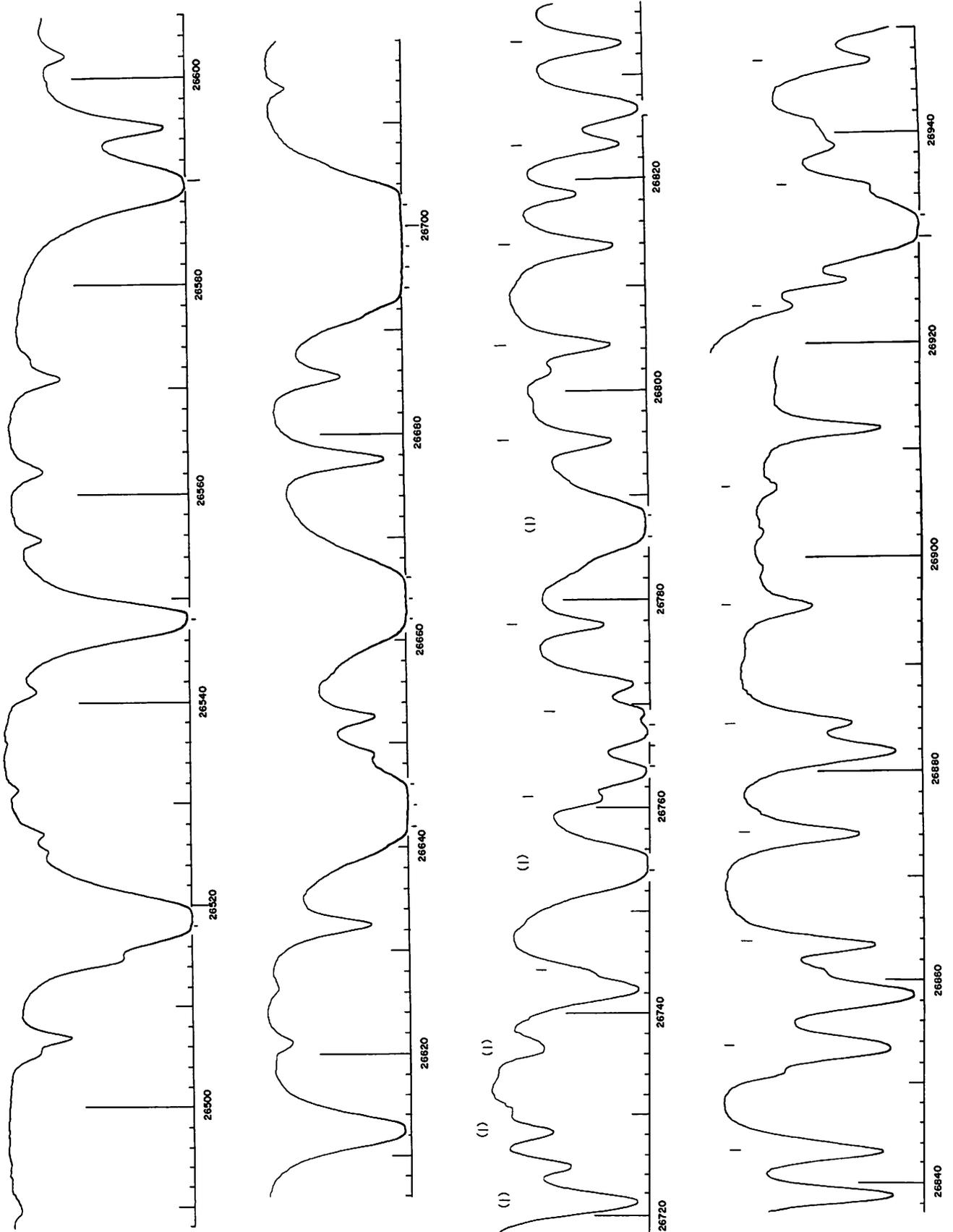


Fig. 23 Laboratory Spectrum of water vapor  $\lambda\lambda$  26488-26950, matching the solar spectrum Figs. 5, 6.

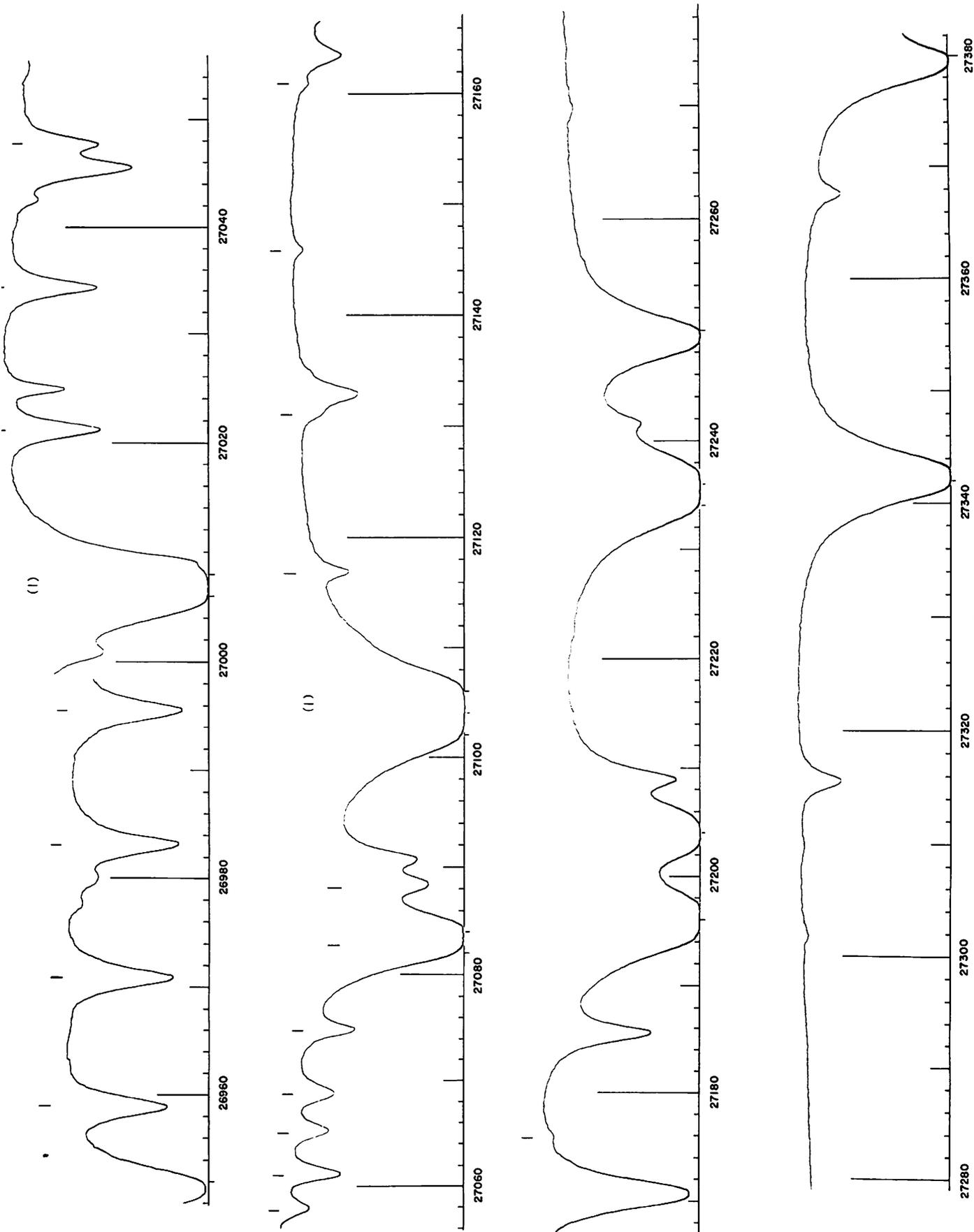


Fig. 24 Laboratory Spectrum of water vapor  $\lambda\lambda$  26950-27382, matching the solar spectrum Figs. 7, 8.

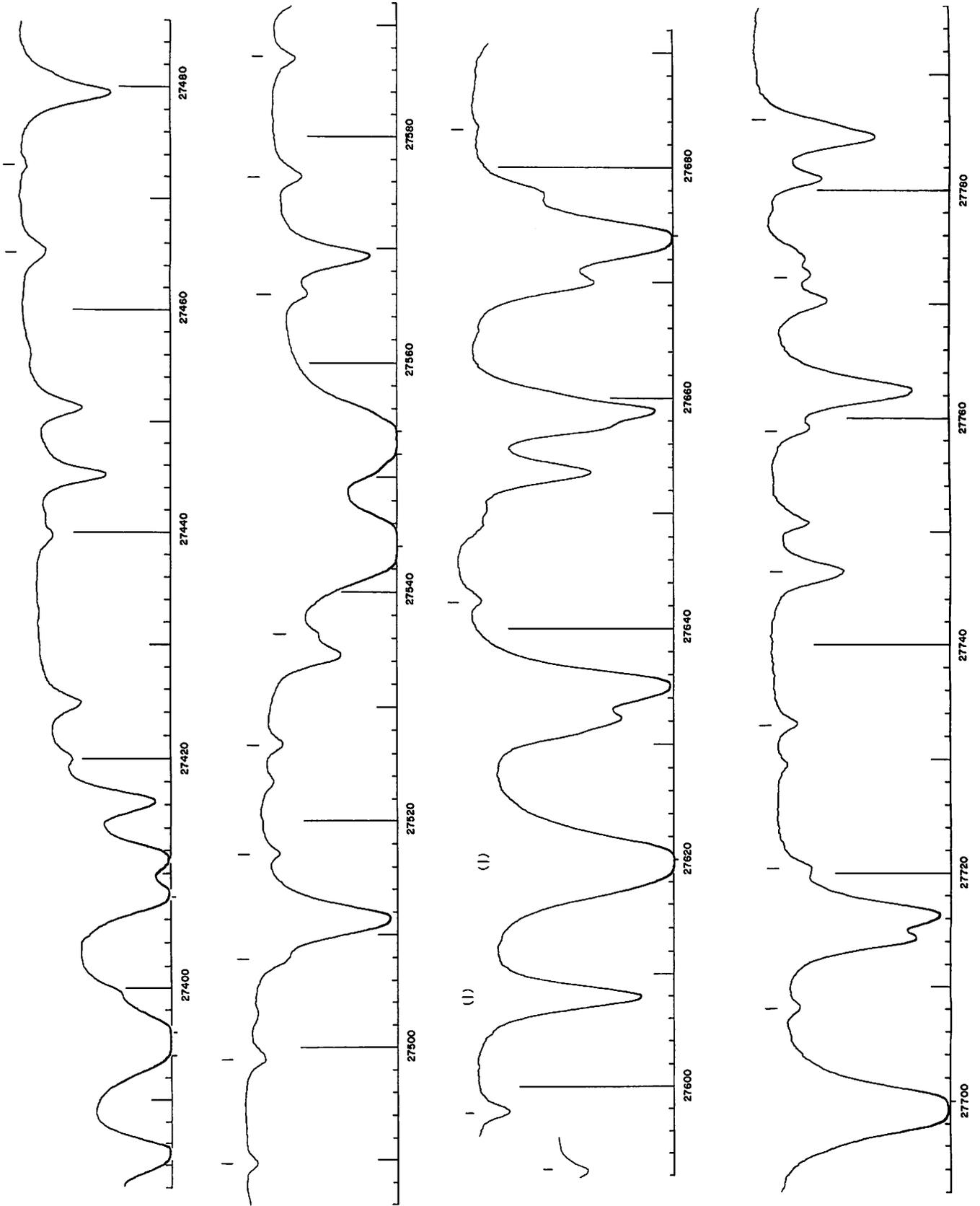


Fig. 25 Laboratory Spectrum of water vapor  $\lambda$  27382-27796, matching the solar spectrum Figs. 9, 10a, b.

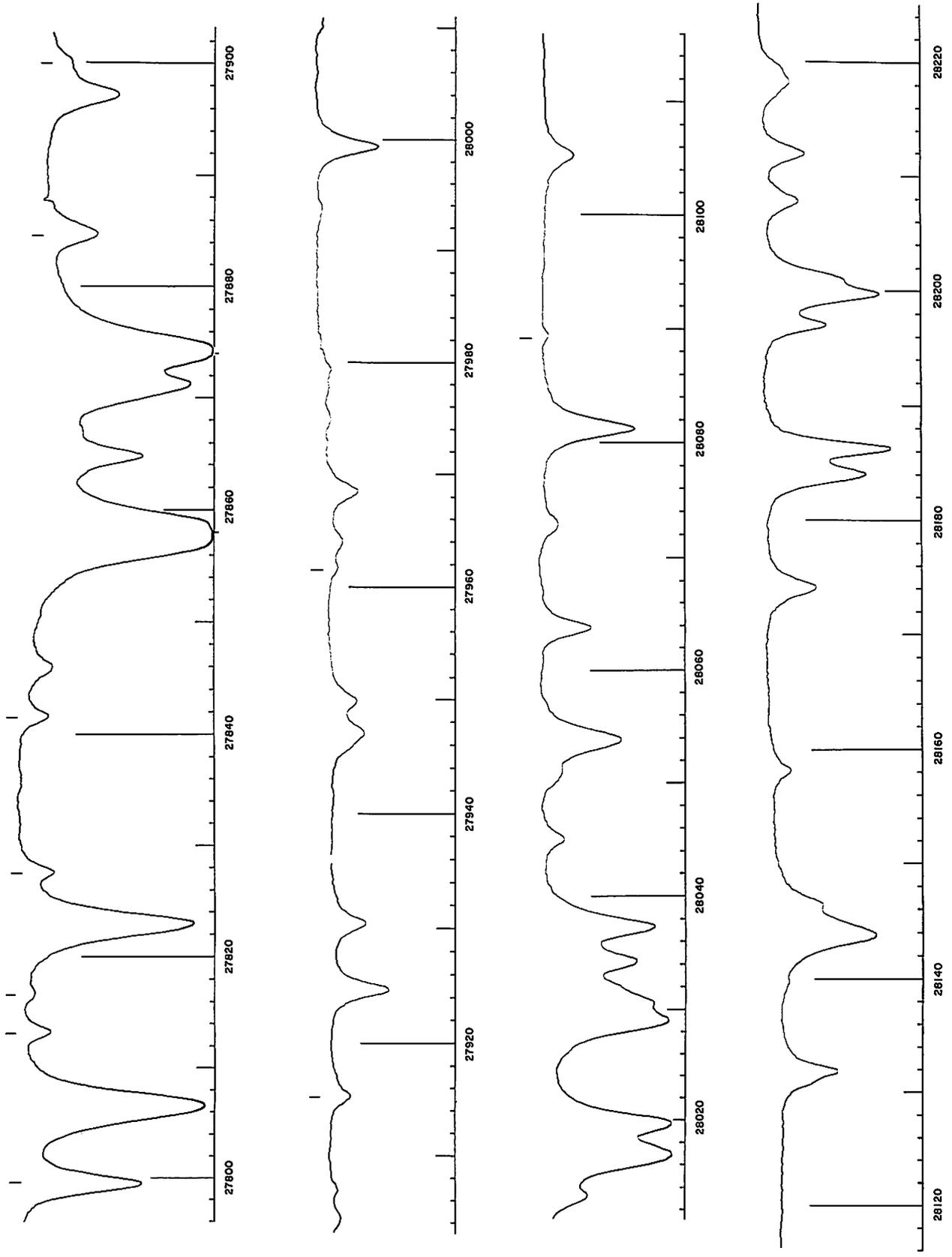


Fig. 26 Laboratory Spectrum of water vapor  $\lambda\lambda$  27796-28225, matching the solar spectrum Figs. 10c/d, 11, 12a/b.

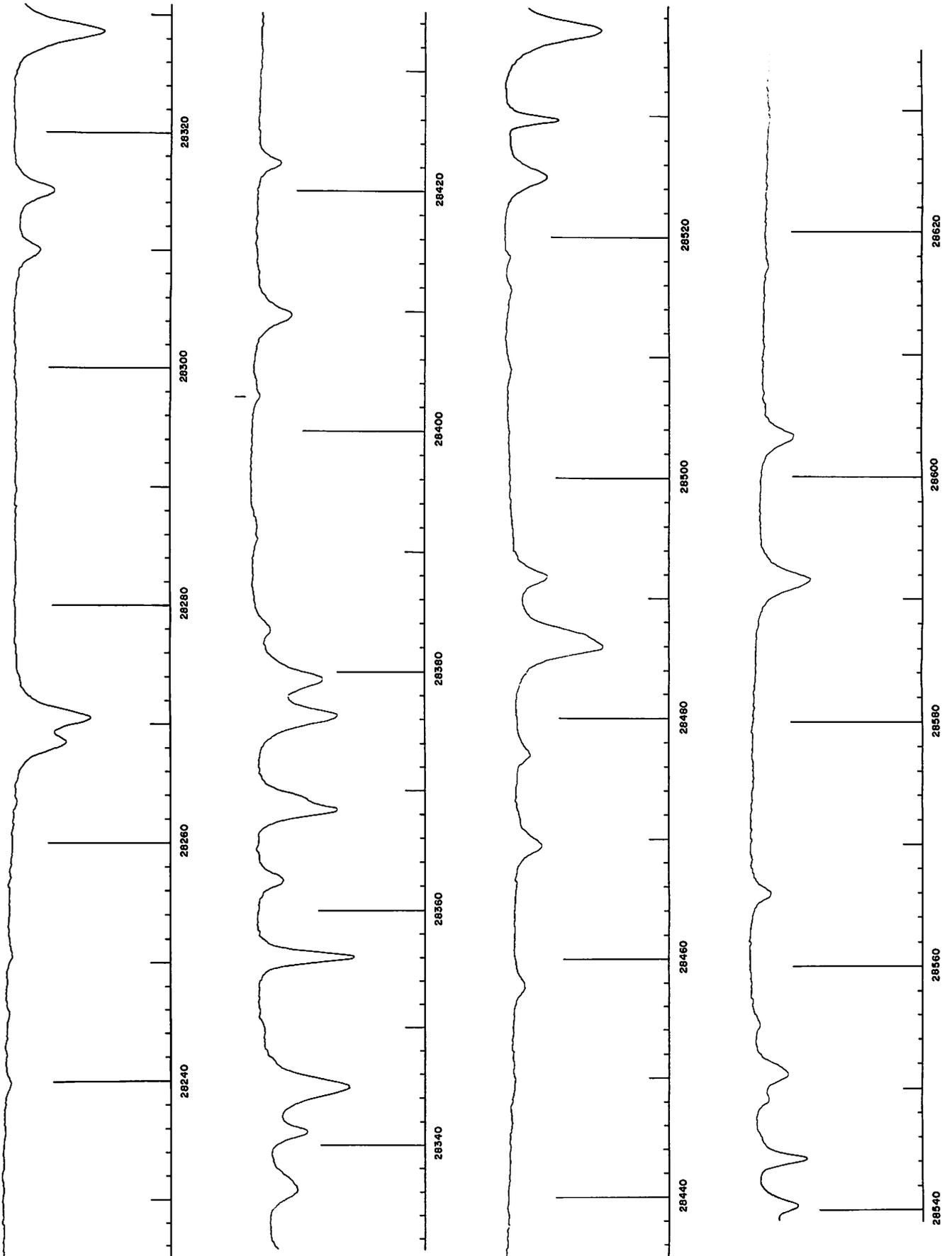


Fig. 27 Laboratory Spectrum of water vapor  $\lambda\lambda$  28225–28635, matching the solar spectrum Figs. 12c/d, 13, 14a/b.

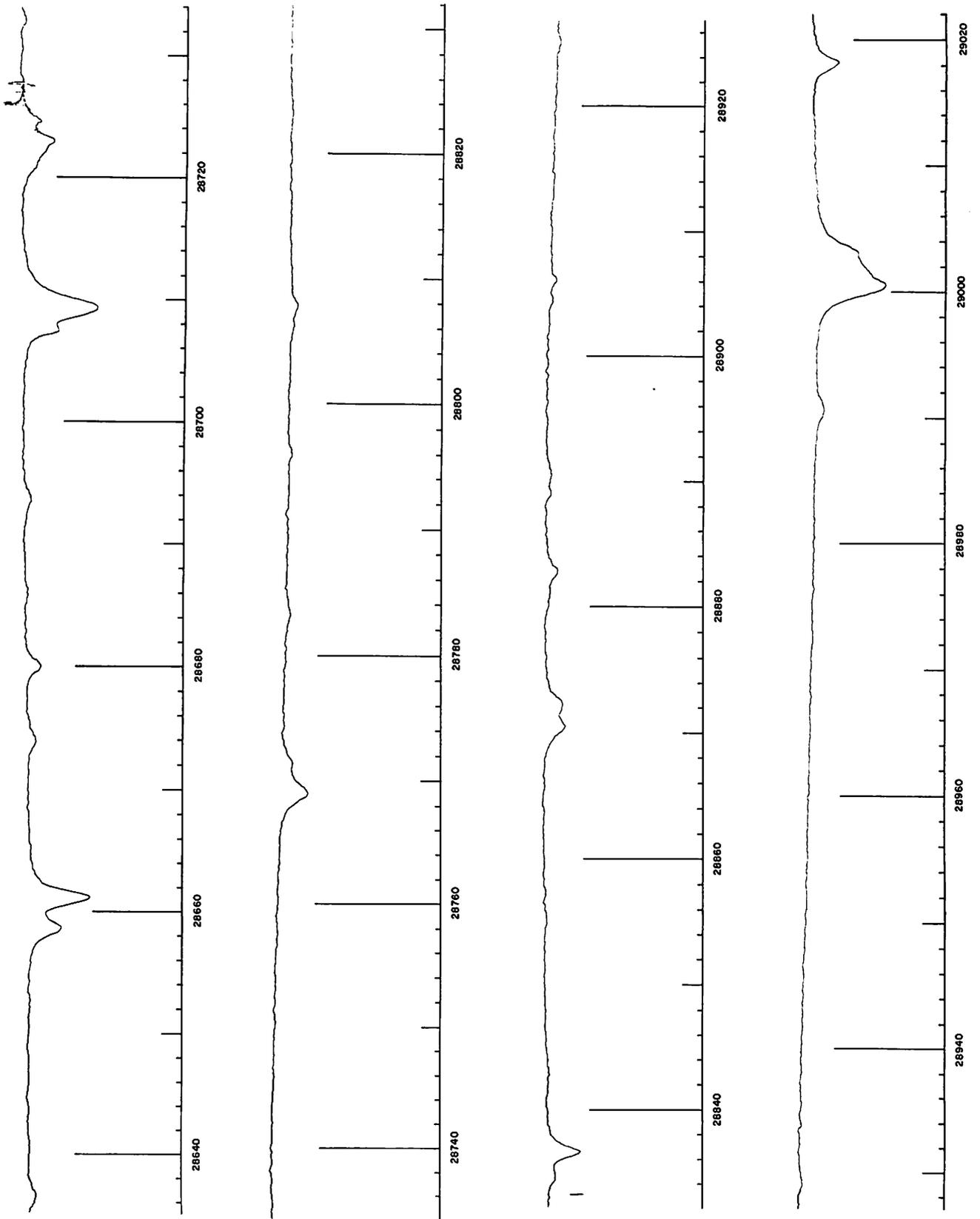


Fig. 28 Laboratory Spectrum of water vapor  $\lambda$  28635-29022, matching the solar spectrum Figs. 14c/d, 15, 16a/b.

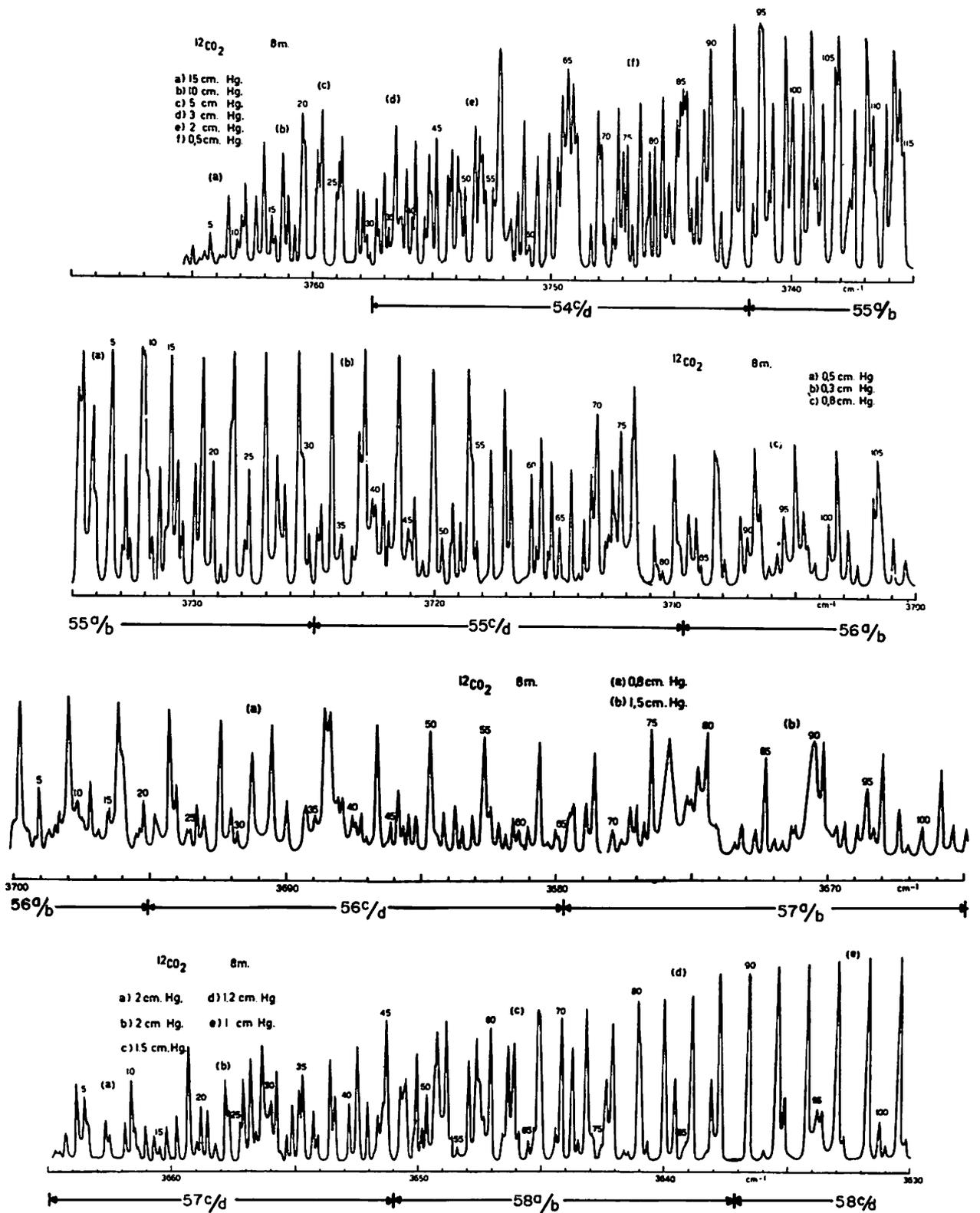


Fig. 29 Laboratory spectrum of  $\text{CO}_2$  by C. P. Courtoy. Scales in wavenumbers. Arrows indicate corresponding parts of solar records. (Reproduced with permission).

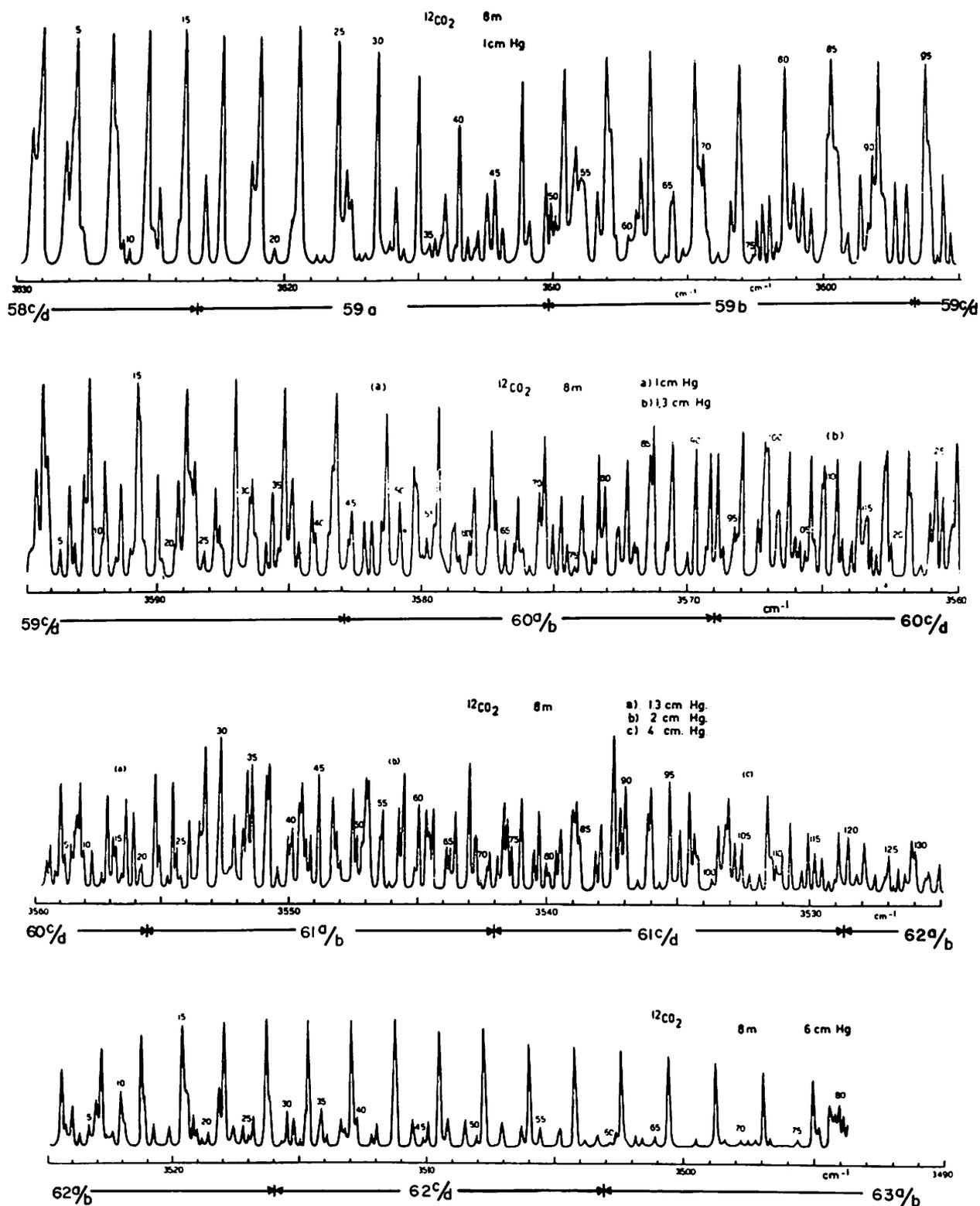


Fig. 30 Laboratory spectrum of  $\text{CO}_2$  by C. P. Courtoy. Scales in wavenumbers. Arrows indicate corresponding parts of solar records. (Reproduced with permission).

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