# No. 51 MICROMETRIC MEASURES OF DOUBLE STARS 

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May 15, 1965

TThe present series of measures is the continuation of the observations published in Vol. IX, Part II, of the Publications of the Yerkes Observatory (1960). The four different telescopes at which these measurements were made are, in order of aperture, the 36 -inch reflector of the McDonald Observatory, the 40 -inch refractor of the Yerkes Observatory, the 82 -inch reflector of the McDonald Observatory, and lastly, the 84 -inch reflector of the Kitt Peak National Observatory. Rather than indicating which instrument was used for each measure, the dates of observation are listed to enable the reader to identify the telescope employed:

| $\quad$ Date | Telescope |
| :--- | :---: |
| Before 1959 | $36^{\prime \prime}$ |
| 59.624 to .660 | 36 |
| 59.950 | 82 |
| 60.523 to .601 | 82 |
| 60.870 | 40 |
| 61.097 | 40 |
| 61.201 to .265 | 36 |
| 61.330 to .428 | 40 |
| 61.551 to .580 | 40 |
| 61.864 | 36 |
| 62.288 | 36 |
| 62.285 to .313 | 82 |
| 62.373 | 40 |
| 62.690 to .710 | 36 |
| 62.953 to .959 | 36 |
| 63.334 to .359 | 36 |
| 64.090 | 82 |
| 64.404 to .408 | 82 |
| After 64.78 | 84 |

The telescopes generally used for double-star work are of smaller aperture than the four mentioned above. Therefore, I devoted my attention whenever possible to the very close binaries and to the others that passed near periastron. Only when
the seeing conditions were not adequate was the micrometer used on wider pairs, preference being given to those that are too faint to be reliably measurable with smaller instruments.

The star positions are referred to the equinox of 1950. For identification, the BD number has been added except when the star is bright enough to have a Greek letter designation. The magnitudes are on the Harvard scale and mostly taken (as were the spectra) from the new Index Catalogue (IDS) (Jeffers, van den Bos, and Greeby, 1963).

Comparisons have been made with the latest available orbits and are given in the notes. Also added are dynamical parallaxes whenever the accumulated material suffices for a good determination. For them the precepts of H. N. Russell and C. E. Moore were used (1940). Many of these dynamical parallaxes supersede earlier determinations from shorter arcs.

Acknowledgments. This work has been supported by Grant No. GP-2335 from the National Science Foundation which is hereby gratefully acknowledged.

## REFERENCES

Jeffers, Hamilton M., van den Bos, Willem H., and Greeby, Frances M. 1963, "Index Catalogue of Visual Double Stars, 1961.0," Publications of the Lick Observatory, 21, parts 1 and 2 (Mount Hamilton: University of California Printing Department).
Russell, Henry N. and Moore, C. E. 1940, "The Masses of the Stars with a General Catalogue of Dynamical Parallaxes," Astrophysical Monographs (Chicago: University of Chicago Press).
Van Biesbroeck, G. 1960, "Measurements of Double Stars," Publications of the Yerkes Observatory, 9, part 2 (Chicago: University of Chicago Press).

| $\underline{2} \mathrm{Mlb} 106$ |  |  |  |
| :---: | :---: | :---: | :---: |
| $0^{\text {h }} 0 . \mathrm{m} 0$ | $+59^{\circ} 26^{\prime}$ |  | 10.5-10.7 |
| 59.638 | 43.5 | 2.:06 |  |
| 59.646 59.950 | 41.7 41.4 | 2.10 |  |
| $\underline{60.618}$ | 44.8 | $\underline{2.11}$ |  |
| 59.96 | 42.9 | 2.02 | 4 n |
| $13 \beta 861$ |  |  | $+68^{\circ} 1422$ |
| $0^{h_{0}}{ }^{\text {m }} 5$ | $+69^{\circ} 25^{\prime}$ |  | 10.5-10.9 |
| 59.638 | 180.7 | 1".56 | $\Delta m=0.4$ |
| $\begin{array}{r}59.646 \\ 59.950 \\ \hline\end{array}$ | 176.6 178.3 | 1.64 <br> 1.75 | $\Delta \mathrm{m}=0.4$ |
| 59.74 | 178.5 | 1.65 | 3 n |

Change questionable. The $\Delta_{m}$ is certainly unch smaller than $1 \times 3$ given in ADS.

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No change since 1911.
Identified as Par. ph. $+22^{\circ}, 0^{\mathrm{h}} 0^{\mathrm{m}}$, No. 185 which gives the corrected position.


Conteau's orbit (1959) gives the residuals -0.8 and $0!.00$.

| 62 J 143 |  | $+12^{\circ} 5065$ |  |
| :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} \mathrm{m}^{\text {m }} 5$ | +12 ${ }^{\circ} 59^{\prime}$ |  | 11.8-11.7 |
| 59.942 | 89.5 | 2.45 |  |
| 59.964 | 88.2 | 2.65 |  |
| 60.722 | 89.7 | 2.44 |  |
| $\underline{60.870}$ | 89.0 | $\underline{2.61}$ |  |
| 60.37 | 89.1 | 2.54 | 4 n |
| Distance increasing. 2.54 4n |  |  |  |
| $80 \mathrm{Kü} 3$ |  | $+19^{\circ} 2$ |  |
| $0^{h_{4} \mathrm{~m}_{6} 9}$ | +20 ${ }^{\circ} 13^{\prime}$ | 10.2-10.2 |  |
| 61.664 | 80:7 | 0'!93 |  |
| 61.718 | 79.4 | 0.96 |  |
| 61.820 | 81.6 | 0.82 |  |
| 62.645 | 78.8 | 0.91 |  |
| 62.710 | 80.3 | 0.92 |  |
| 64.740 | 78.0 | 1.06 |  |
| 65.034 | 80.5 | 1.08 |  |
| 62.90 | 79.9 | 0.95 |  |
| Unchanged since 1901. 0.95 n |  |  |  |
| 97 Mlb 37 |  | $+56^{\circ} 5$ |  |
| $0^{\mathrm{h}_{5} \mathrm{~m}_{\text {. }} 9}$ | +56 ${ }^{\circ} 5^{\prime}$ | 10.9-11.0 |  |
| 59.624 | 199.9 | 3:00 |  |
| 59.638 | 198.5 | 2.96 |  |
| 59.646 | 198.7 | 3.00 |  |
| 59.950 | 198.1 | $\underline{2.95}$ |  |
| 59.71 | 198.8 | 2.98 | $4 \pi$ |
| Unchanged since 1918. |  |  |  |
| 134 Kr 1 |  | $+57^{\circ} 23$ |  |
|  | +57 ${ }^{\circ} 34^{\prime}$ | 10.0-10.5 |  |
| 61.580 | 192.8 | 2':48 |  |


| 61.716 | 191.9 | 2.51 |  |
| :--- | :--- | :--- | :--- |
| $\frac{61.842}{}$ | $\underline{194.2}$ | $\underline{2.54}$ |  |
| 61.71 | 193.0 | $\underline{2.51}$ | $3 n$ |

The increase in distance indicates $d p=0!023$.


| $0^{\text {h }}$ 9 ${ }^{\text {m }}$. | $+28^{\circ} 9^{\prime}$ |  |
| :---: | :---: | :---: |
| 59.638 | 87.3 | 0"70 |
| 59.646 | 87.6 | 0.67 |
| 59.950 | 85.4 | 0.59 |
| 60.601 | 87.5 | 0.60 |
| 60.615 | 86.6 | 0.58 |
| 61.718 | 83.7 | 0.64 |
| 61.771 | 86.4 | 0.59 |
| $\underline{61.820}$ | 88.9 | 0.69 |
| 60.73 | 86.7 | 0.63 |

Slow decrease in angle and increase in distance corresponding to $\mathrm{dp}=0!0041$.

| $148 \beta 1026$ |  | $+52^{\circ} 19$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} 9^{\mathrm{m}_{5}}$ | $+53^{\circ} 21^{\prime}$ |  | 7.2-8.0 | Fo |
| 59.656 | $332: 3$ | 0! 36 |  |  |
| 59.939 | 335.9 | 0.40 |  |  |
| 59.950 | 329.3 | 0.42 |  |  |
| 60.618 | 336.7 | 0.37 |  |  |
| 60.04 | $\begin{gathered} 333.6 \\ (1955)^{0.39} \\ \text { gives the } \end{gathered}$ |  | 4n |  |
| Baize's orbit and +0.04 . |  |  | e residuals $+6: 0$ |  |
| 202 Es 865 | $+51^{\circ} 33$ |  |  |  |
| $0^{\mathrm{h}} 13 .{ }^{\text {m }} 2$ | $+52^{\circ} 16^{\prime}$ | 10.9-11.1 |  |  |
|  | AB |  |  |  |
| 59.638 | 356:9 | 1.141 |  |  |
| 59.656 | 355.1 | 1.23 |  |  |
| 59.950 | 356.1 | 1.23 |  |  |
| 60.618 | 359.0 | 1.12 |  |  |
| 59.97 | 356.8 | 1.25 | $4 n$ |  |
|  | AC |  |  |  |
| 60.618 | 72:8 1 | 11'90 |  |  |

$A C$ is unchanged but there is a large change in $A B$.

$$
\begin{aligned}
& 207 \sum 13 \\
& +76^{\circ} 5 \\
& \begin{array}{ccccc}
0^{h_{13}} 13.4 & +76^{\circ} 40^{\prime} & & 6.7-7.2 & \text { B9 } \\
59.624 & 68.3 & 0.96 & & \\
59.638 & 65.5 & 0.80 & & \\
\frac{59.646}{59.64} & \frac{67.5}{67.1} & \underline{0.86} & & \\
\hline
\end{array}
\end{aligned}
$$

How completely undeterminate the elements still are is shown by the nearly similar residuals from the orbit of Rabe (1961) with a period of 720 years and that of Heintz (1960) with a period of 1600 years:

| Rabe | +3.6 | +0.04 |
| :--- | :--- | ---: |
| Heintz | +3.5 | 0.00 |

274 J 630

| $0^{\mathrm{h}} 17 \mathrm{~m}_{6}$ | $+21^{\circ} 13^{\prime}$ | $10.0-10.6$ |  |
| :--- | :--- | :--- | :--- |
| 62.940 | $118 \% 4$ | $2 . .55$ |  |
| $\frac{64.037}{63.49}$ | $\frac{117.0}{117.7}$ | $\frac{2.67}{2.61}$ | 2 n |


| 285 AC 1 |  |  | +32 ${ }^{\circ} 48$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} 18 . \mathrm{m}_{3}$ | +32 ${ }^{\circ} 42^{\prime}$ |  | 7.5-8.0 | FS |
| 59.624 | 290.3 | 1'60 |  |  |
| 59.638 | 290.2 | 1.50 |  |  |
| 59.646 | 288.5 | 1.49 |  |  |
| 59.64 | 289.7 | 1.53 | 3n |  |

Slow change in both coordinates making dp $=0!018$.

| 287 - 1093 |  | $+10^{\circ}$ |
| :---: | :---: | :---: |
| $0^{\mathrm{h}} 18 . \mathrm{m}_{3}$ | $+10^{\circ} 42^{\prime}$ | 7.0-7.9 |
| 61.820 | $103: 5$ | 1 n |

Long period binary with $\mathrm{dp}=0!009$.


The distance is hardily changed but the angular motion indicates a period of about 130 years and $\mathrm{dp}=0: 0030$.

| 344 Ho 210 |  |  | $+35^{\circ} 66$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} \mathbf{2 3}^{\text {m }}$. 0 | $+36{ }^{\circ} 12^{\prime}$ |  | 8.4-10.1 | F0 |
| 61.820 | $78: 7$ | 0!.97 |  |  |
| 62.718 | 72.8 | 0.93 |  |  |
| 62.959 | 74.0 | 0.97 |  |  |
| 65.040 | 79.7 | 0.93 |  |  |
| 63.13 | 76.3 | 0.95 | $4 n$ |  |
| Little change | 80 yea | . |  |  |
| 350 O§ 9 |  |  | $+55^{\circ} 72$ |  |
| $0^{\mathrm{h}} 23 .{ }^{\text {m }} 5$ | $+56^{\circ} 30^{\prime}$ |  | 6.9-9.9 | G0 |
| 61.716 | 50.7 | 2:43 |  |  |
| 61.718 | 52.2 | 2.26 |  |  |
| 62.701 | 50.3 | 2.25 |  |  |
| 62.718 | 54.2 | 2.23 |  |  |
| 62.959 | 54.3 | 2.15 |  |  |
| 64.730 | 50.8 | 2.26 |  |  |
| 65.037 | 54.8 | $\underline{2.28}$ |  |  |
| 63.08 | 52.5 | 2.27 | 7n |  |

The small change since 1847 indicates a dp of 0!.071.


371 Hu 1007
$+62^{\circ} 84$
$0^{\mathrm{h}} 25^{\mathrm{m}_{5}} \quad+63^{\circ} 28^{\prime} \quad 10.0-10.0 \quad$ G0

| 58.660 | 229.2 | 0.44 |  |
| :--- | :--- | :--- | :--- |
| 59.950 | 223.2 | 0.46 |  |
| 62.701 | 229.4 | 0.45 |  |
| 62.959 | $\frac{227.2}{21.07}$ | $\underline{0.37}$ |  |
| 627.2 | 0.43 | $4 n$ |  |

The slow change corresponds to $d p=0!0067$.



The residuals from my orbit (1954) are +2.3 and $-0!08$.

| $\underline{692} \beta 781$ |  |  | $+68^{\circ} 56$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} 48 . \mathrm{m}_{4}$ | $+68^{\circ} 3^{\prime}$ |  | 8.7-9.2 | A2 |
| 61.718 | 22.1 | 0'.92 |  |  |
| 61.842 | $\underline{23.0}$ | 1.14 |  |  |
| 61.78 | 22.6 | 1.03 | 2n |  |

The slow decrease in angle makes $d p=0!010$.

| 713 Hu 201 |  |  | $-14^{\circ} 152$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} 49^{\text {m }}$. 6 | $-13^{\circ} 30^{\prime}$ |  | 10.0-10.6 | K0 |
| 62.713 | $84: 3$ | $0 . .44$ |  |  |
| 62.940 | 88.5 | 0.52 |  |  |
| 62.959 | 85.0 | 0.44 |  |  |
| 64.905 | 87.1 | 0.48 |  |  |
| 64.947 | 89.4 | 0.55 |  |  |
| 63.69 | 86.9 | 0.49 | 5 n |  |

The residuals from Couteau's orbit (1961) are $-1: 2$ and $-0!05$.

| 715 A 2306 |  |  | $+16^{\circ} 83$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} \mathrm{Ca}^{\text {m }}$. 8 | $+17^{\circ} 24^{\prime}$ |  | 9.8-10.7 F8 |  |
| 59.646 | $302: 8$ | 1:29 |  |  |
| 59.964 | 299.7 | 1.19 |  |  |
| 60.774 | 304.2 | 1.31 |  |  |
| 60.870 | 303.7 | 1.14 |  |  |
| 60.31 | $\begin{aligned} & 302.6 \\ & \text { since } 1910 . \end{aligned}$ |  | $4 n$ |  |
| Hardly changed since 1910. |  |  |  |  |
| 746 OS 20 |  |  |  |  | $+18^{\circ} 122$ |  |
| $0^{h_{5}} 1^{\mathrm{m}} 9$ | $+18^{\circ} 55^{\prime}$ |  | 6.1-7.2 | A0 |
| 60.722 | 252:8 | 0.'54 |  |  |
| 60.765 | 251.1 | 0.50 |  |  |
| 60.774 | 252.5 | 0.56 |  |  |
| 61.820 | 249.0 | 0.52 |  |  |
| 61.842 | 251.5 | 0.47 |  |  |
| 61.921 | $\underline{253.9}$ | 0.49 |  |  |

The residuals from Couteau's orbit (1963) are $+2: 7$ and +0.02 .

| 763 A 2209 |  |  | $+18^{\circ} 127$ |
| :---: | :---: | :---: | :---: |
| $0^{h_{52}}{ }^{\text {m }} 6$ | $+18^{\circ} 40^{\prime}$ |  | 10.2-10:2 |
| 60.765 | $333: 9$ | 1:54 |  |
| 60.774 | 333.6 | 1.72 |  |
| 60.870 | 337.5 | 1.76 |  |
| 60.80 | 335.0 | 1.67 | $3 n$ |
| Hardly changed | in 50 yea |  |  |


| 789 Но 307 |  |  | $+31^{\circ} 147$ |
| :---: | :---: | :---: | :---: |
| $0^{\mathrm{h}} 54.0$ | $+31^{\circ} 56^{\prime}$ |  | 10.1-10.3 |
| $\begin{array}{r} 64.730 \\ 65.037 \\ \hline \end{array}$ | $\begin{aligned} & 86^{\circ}: 9 \\ & 87.8 \\ & \hline \end{aligned}$ | $\begin{aligned} & 2!86 \\ & 3.05 \\ & \hline \end{aligned}$ |  |
| 64.88 | 87.4 | 2.96 | 2 n |
| The increase | distance | makes | $s \mathrm{dp}=0!01$ |


| 795 |  | $+53^{\circ} 184$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $0^{\mathrm{H}} 44^{\mathrm{h}} \mathrm{m}_{6}$ | $+54^{\circ} 8^{\prime}$ | $9.0-9.5 \mathrm{GO}$ |  |  |
| 60.601 | 150.8 | 0.46 |  |  |
| 62.701 | 156.8 | 0.46 |  |  |
| 62.940 | 153.1 | 0.46 |  |  |
| $\frac{62.964}{62.30}$ | $\frac{153.5}{153.5}$ | $\frac{0.48}{0.47}$ | 4 n |  |

To the increase in angle and decrease in distance corresponds dp = 0! 019.
$828 \beta 867$
$+11^{\circ} 130$
$0^{h_{5}} 7^{m_{6}} 6$
$+11^{\circ} 40^{\prime}$
9.3-9.8 F5
$59.656 \quad 147.7 \quad 0.31$
$59.964 \quad 147.2 \quad 0.28$
$59.81 \quad 147.4 \quad 0.30 \quad 2 n$
Large change in both coordinates making dp $=$ 0.028 .

836 A $2901 \quad+68^{\circ} 64$
 $\mathrm{dp}=0,005$


| $\beta 1161$ |  |  | $+51^{\circ} 216$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} 0^{\mathrm{m}} 0$ | +51 ${ }^{\circ} 32^{\prime}$ |  | 7.2-8.0 | B5 |
| 62.701 | 359:6 | 0'.40 |  |  |
| 62.940 | 362.5 | 0.42 |  |  |
| 64.730 | 358.8 | 0.45 |  |  |
| 65.040 | 361.4 | 0.44 |  |  |
| 63.85 | 360.6 | 0.43 | 4n |  |

To the slow increase in angle corresponds $d p=$ 0",0067.

869 J 874

$$
1^{h_{0} m_{6}} \quad+22^{\circ} 16^{\prime} \quad 9.5-12.0
$$

| $\begin{array}{r} 62.953 \\ 62.959 \\ \hline \end{array}$ | $\begin{aligned} & 305: 0 \\ & 302.6 \\ & \hline \end{aligned}$ | $\begin{array}{r} 2!67 \\ 2.50 \\ \hline \end{array}$ |  |
| :---: | :---: | :---: | :---: |
| $\begin{gathered} \overline{62.96} \\ \text { Slow decrease } \end{gathered}$ | $\begin{gathered} \overline{303.8} \\ \text { in angle. } \end{gathered}$ | 2.58 | 2n |
| 896 AG 14 |  |  | $+20^{\circ} 154$ |
| $1^{\text {h }}$ 2 ${ }^{\text {ma }} 9$ | $+20^{\circ} 52^{\prime}$ |  | 9.7-10.1 K0 |
| 59.964 62.959 | 111.1 99.4 | $\begin{aligned} & 0!36 \\ & 0.32 \end{aligned}$ |  |

A pair to watch as it closes in and speeds up in angle. dp comes out 0.017 .

| 902 - 1228 |  | $+12^{\circ} 133$ |  |
| :---: | :---: | :---: | :---: |
| $1^{h_{3} \mathrm{~m}_{2}}$ | $+13^{\circ} 3^{\prime}$ |  | 9.9-11.0 60 |
| 59.646 | 259.1 | 0:76 |  |
| 59.964 | 264.8 | 0.74 |  |
| 60.774 | 266.2 | 0.73 |  |
| 61.820 | $\underline{264.5}$ | 0.78 |  |
| 60.55 | 263.6 | 0.75 | $4 n$ |

Practically unchanged in 70 years.

| 967 - 1162 |  |  | $+35^{\circ} 215$ |
| :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} \mathrm{m}^{\mathrm{m}} \mathrm{B}$ | $+35^{\circ} 41^{\prime}$ |  | 9.9-10.1 A5 |
| $\begin{array}{r} 60.618 \\ 62.959 \\ \hline \end{array}$ | $\begin{array}{r} 145.8 \\ 149.0 \\ \hline \end{array}$ | $\begin{aligned} & 0.34 \\ & 0.33 \\ & \hline \end{aligned}$ |  |
| 61.79 | 147.4 | 0.34 | 2n |
| Hardly changed in 70 years. |  |  |  |
| 989 Hu 1024 |  |  | +50 ${ }^{\circ} 240$ |
| $1^{\text {h }} 10^{\text {m. }} 2$ | $+50^{\circ} 50^{\prime}$ |  | 9.4-10.2 |
| $\begin{aligned} & 59.646 \\ & 59.964 \end{aligned}$ | $\begin{array}{r} 206.6 \\ 204.5 \\ \hline \end{array}$ | $\begin{aligned} & 0: 85 \\ & 0.68 \end{aligned}$ |  |
| 59.80 | 205.6 | 0.76 | 2 n |


| 999 - 1100 |  |  | $+60^{\circ} 193$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} 11^{\text {m }} 6$ | $+60^{\circ} 41^{\prime}$ |  | 8.3-8.3 | F5 |
| 59.687 | 53.3 | $0!30$ |  |  |
| 59.964 | 52.4 | 0.36 |  |  |
| 60.601 | 56.0 | 0.37 |  |  |
| 60.08 | 53.9 | 0.34 |  |  |

Muller's orbit (1958) gives the residuals $+8 \% 4$ and -0.005 .

| $1^{\text {h }} 16{ }^{\text {m/3 }}$ | $+64^{\circ} 2^{\prime}$ |  | 9.7-10.3 |
| :---: | :---: | :---: | :---: |
| 59.638 | $235: 6$ | 2!08 |  |
| 59.646 | 238.8 | 1.97 |  |
| 59.964 | $\underline{235.3}$ | 1.82 |  |
| 59.75 | 236.6 | 1.96 | 3n |

To the advance in angle corresponds $\mathrm{dp}=0 \%, 024$. The star was identified as Vat. ph. $63^{\circ} 3288$ which gives the corrected position.

| 1097 ß4 |  |  | $+10^{\circ} 168$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} 18.7$ | $+11^{\circ} 17^{\prime}$ |  | 7.4-7.9 | F0 |
| 60.774 | $134: 3$ | $0!28$ |  |  |
| 64.090 | 131.7 | 0.32 |  |  |
| 65.040 | 131.4 | 0.27 |  |  |
| 63.30 | 132.5 | 0.29 | $3 n$ |  |

Muller's orbit (1961) gives the residuals $+6: 8$ and $-0!02$.

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$\frac{65.037}{64.68} \quad \frac{211: 2}{214.0} \quad \frac{0.34}{0.34} 3 n$

Van den Bos' orbit (1963) gives the small residuals +0.1 and $-0!.01$.

| $\underline{1145} \beta 1102$ |  |  | $+59^{\circ} 251$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} 24^{\mathrm{m}} .2$ | $+60^{\circ} 2^{\prime}$ |  | 8.6-10.3 | B2 |
|  | BC |  |  |  |
| 59.646 | $339: 2$ | 1.06 |  |  |
| 60.601 | 342.8 | 0.82 |  |  |
| 62.957 | 339.7 | 0.81 |  |  |
| $\text { Hardly }{ }^{61.07} \text { changed }$ | 340.6 | 0.90 | $3 n$ |  |
|  | in 72 y |  |  |  |

$$
1161 \text { AC } 14 \quad+42^{\circ} 308
$$



Change questionable after a century.

| 1177 A 2214 |  |  | $+19^{\circ} 243$ |
| :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} 26{ }^{\text {m }}$ 5 | $+19^{\circ} 48^{\prime}$ |  | 10.2-10.3 |
| 58.077 | 198.2 | 0'.74 |  |
| 59.687 | 197.7 | 0.82 |  |
| 60.588 | 198.5 | 0.73 |  |
| 60.774 | 194.2 | 0.68 |  |
| 59.78 | 197.2 | 0.74 | 4n |

To the small change in angle corresponds $d p=$ 0.,006.

1183 A 191
$+22^{\circ} 236$

| $1^{\mathrm{h}_{27} \mathrm{~m}_{0}}$ | $+22^{\circ} 34^{\prime}$ |  | $7.5-7.7$ | $A 0$ |
| :---: | :--- | :--- | :--- | :--- |
|  |  |  |  |  |
| 59.687 | 161.9 | 0.26 |  |  |
| 60.588 | 162.9 | 0.26 |  |  |
| 60.744 | 161.6 | 0.19 |  |  |
| $\frac{65.040}{61.51}$ | $\frac{162.3}{162.2}$ | $\frac{0.23}{0.24}$ | 4 n |  |

The residuals from Nuller's orbit (1958) are $+5: 9$ and $0!00$.

| $1^{\mathrm{h}} 36.4$ | $-18^{\circ} 12^{\prime}$ | 12-12 |
| :---: | :---: | :---: |
| 61.743 | 15:8 | 2!56 |
| 62.340 | 9.4 | 2.76 |
| 62.959 | 9.5 | 2.58 |
| 64.905 | 1.8 | 2.55 |
| 64,947 | 2.4 | 2.40 |
| 63.38 | 7.8 | 2.57 5n |

The residuals from Luyten's orbit (1961) are -7.5 and -0.20 .

| $1294 \beta 508$ |  |  | +26 ${ }^{\circ} 276$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1^{\mathrm{h}} 36 . \mathrm{m}$ | +26 ${ }^{\circ} 1^{\prime}$ |  | 9.8-10.3 | F8 |
| 60.774 | 63:6 | 0'.55 |  |  |
| 61.743 | 64.4 | 0.64 |  |  |
| 61.844 | 62.8 | 0.66 |  |  |
| 61.45 | 63.6 | 0.62 | 3n |  |

The longer arc makes $d p=0.4008$.

| 1345 A 1 |  |  | $-7^{\circ} 282$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1^{h_{39}}{ }^{\text {ma }} 9$ | $-7^{\circ} 0^{\prime}$ |  | 8.5-9.0 | F2 |
| 62.718 | 239:0 | 0.72 |  |  |
| 62.940 | 234.4 | 0.70 |  |  |
| 62.959 | 232.7 | 0.68 |  |  |
| 64.705 | 234.4 | 0.83 |  |  |
| 64.947 | $\underline{236.2}$ | 0.87 |  |  |
| 63.65 | 235.3 | 0.76 | 5n |  |

The longer arc makes $\mathrm{dp}=0!016$.


The 0!.034 proper motion of the bright component proves the physical connection. Hardly changed after 38 years.

| 1458 | A 2602 |  | $-4^{\circ} 281$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | $1 \mathrm{~h}_{47} \mathrm{~m}_{4}$ | $-4^{\circ} 22^{\prime}$ |  | 10.4-10.4 | K2 |
|  | 58.598 | 358.4 | 0.145 |  |  |
|  | 59.656 | 360.2 | 0.54 |  |  |
|  | 60.774 | 361,7 | 0.39 |  |  |
|  | 59.68 | 0.1 | 0.46 | 3 n |  |

The change in both coordinates makes $d p=$ 0.1006 .

| 1473 Ho 311 |  |  | $+23^{\circ} 246$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1 \mathrm{~h}_{48 . \mathrm{m}}^{4}$ | +24 ${ }^{\circ} 24^{\prime}$ |  | 7.6-7.8 | A5 |
| 60.870 | 203:8 | 0.'38 |  |  |
| 61.743 | 203.1 | 0.44 |  |  |
| 61.921 | $\underline{200.6}$ | 0.38 |  |  |
| 61.51 | 202.5 | 0.40 | 3 n |  |

To the slow increase in angle corresponds $\mathrm{dp}=0!0020$.

1499 J 671

$$
\begin{array}{llll}
1^{\mathrm{h}_{50 \mathrm{~m}_{1}}} & +21^{\circ} 37 & & 9.5-9.7 \\
61.738 & 155: 8 & 2!93 & \\
\frac{61.743}{61.74} & \frac{154.5}{155.2} & \frac{2.80}{2.87} & 2 \mathrm{n}
\end{array}
$$

No appreciable change in 50 years. In the Paris zone $+22^{\circ}$, $1^{\mathrm{K}} \mathrm{m}^{\mathrm{m}}$, No. 135 the star is marked as "double?"' The pair is also Paris $+21^{\circ}, 1 h_{4} 8^{\mathrm{m}}$, Nos. 45 and 46 making $1893.91 \quad 157.0$ 2!.50.

| 1503 - 260 |  |  | $+14^{\circ} 298$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $1_{2}{ }_{50}{ }^{\text {m }}$ 5 | $+15^{\circ} 11^{\prime}$ |  | 8.8-9.5 | F5 |
| 60.870 | $251: 4$ | 0'.97 |  |  |
| 61.664 | 252.6 | 1.26 |  |  |
| 61.738 | 255.7 | 1.24 |  |  |
| 61.811 | 252.6 | 1.13 |  |  |
| 61.820 | 253.8 | 1.24 |  |  |
| 61.921 | 252.9 | 1.14 |  |  |
| 61.64 | 253.2 | 1.16 | 6n |  | $d p=0!013$.

$$
\underline{1530} \text { A } 2407 \quad+2^{\circ} 296
$$

| $1^{\mathrm{h}_{52}{ }^{\mathrm{m}} .8}$ | $+2{ }^{\circ} 43^{\prime}$ |  | $9.6-11.6$ | KO |
| :--- | :--- | :--- | :--- | :--- |
|  |  |  |  |  |
| 62.959 | 190.2 | $0!69$ |  |  |
| $\frac{64.740}{63.85}$ | $\frac{193.4}{191.8}$ | $\frac{0.79}{0.74}$ | 2 n |  |

Both coordinates increase making dp $=0$ ', 007.


No change in distance but the increase in angle makes $\mathrm{dp}=0!007$.

| Ho 312 |  |  | $+25^{\circ} 349$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{h_{4} m_{0}} 0$ | $+25^{\circ} 28^{\prime}$ |  | 6.0-11.5 | B8 |
| 59.638 | $343: 8$ | 1!60 |  |  |
| 62.940 | 344.8 | 1.64 |  |  |
| 62.959 | 339.6 | 1.48 |  |  |
| 64.090 | 341.3 | $\underline{1.55}$ |  |  |
| 62.41 | 342.4 | 1.57 | 4n |  |

Slow increase in angle with no definite change in distance making dp $=0!005$. The $0!04$ proper motion of the bright component confirms the physical connection.

| 1674 Hu 16 |  |  | -10 438 |
| :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}} \mathrm{m}_{2}$ | $-10^{\circ} 19^{\prime}$ |  | 10.2-11.4 |
| 61.844 | $3: 9$ | 1!05 |  |
| 62.704 | 8.0 | 0.98 |  |
| 62.718 | 6.6 | 1.09 |  |
| 62.940 | 5.7 | 0.96 |  |
| $\underline{64.090}$ | 3.0 | 1.18 |  |
| 62.86 | 5.4 | 1.05 | 5n |

The longer arc makes $d p=0!012$.
1680 A 2325
$+0^{\circ} 358$

| $2^{h_{7} \mathrm{~m}_{2}}$ | $+0^{\circ} 34$ |  | $9.3-9$ |
| :--- | :--- | :--- | :--- |
| 59.656 | 118.0 | 0.31 |  |
| 61.743 | 125.9 | 0.29 |  |
| 61.844 | 121.6 | 0.29 |  |
| 64.090 | 118.7 | 0.33 |  |
| 64.740 | 121.8 | 0.27 |  |
| 62.41 | 121.2 | 0.30 | $5 n$ |

Only a slight increase in angle after 50 years.

| 1780 A 961 |  |  | +29 ${ }^{\circ} 393$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}} 17^{\mathrm{m}} .2$ | $+29^{\circ} 35^{\prime}$ |  | 9.3-9.3 | F5 |
| 62.704 | $329: 6$ | $0 . .47$ |  |  |
| 62.954 | 323.6 | 0.46 |  |  |
| 65.040 | 323.2 | 0.39 |  |  |
| 63.57 | 325.5 | 0.44 | 3 n |  |



$$
\frac{59.950}{59.84} \quad \frac{69: 9}{68.9} \quad \frac{1!07}{0.98} 3 n
$$

The angle is hardly changed in 55 years but to the increase in distance corresponds dp $=0!1013$.

| 1833 \ 257 |  |  | $+60^{\circ} 472$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}} 21{ }^{\mathrm{m}}{ }^{9}$ | $+61^{\circ} 20^{\prime}$ |  | 7.6-8.1 | B8 |
| 59.964 | 10:9 | 0.'29 |  |  |
| 62.704 | 12.1 | 0.31 |  |  |
| 62.959 | 12.5 | 0.26 |  |  |
| 64.740 | 12.7 | 0.28 |  |  |
| 65.040 | 10.0 | 0.25 |  |  |
| 63.08 | 11.6 | 0.28 | 5n |  |
| Residuals fro | Toledo's | orbit | (1954): |  |
| -1.1 and +0! |  |  |  |  |


| 1837 A 446 |  | $-6^{\circ} 473$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}} 22 . \mathrm{m}_{2}$ | - $6^{\circ} 07^{1}$ |  | 10.1-10.3 | F5 |
| 61.811 | $364: 1$ | 0.55 |  |  |
| 61.820 | 360.8 | 0.49 |  |  |
| 61.844 | 364.0 | 0.48 |  |  |
| 62.940 | 352.0 | 0.46 |  |  |
| 62.959 | 354.5 | 0.46 |  |  |
| 64.090 | 363.3 | 0.59 |  |  |
| 64.740 | 360.7 | 0.51 |  |  |
| 64.905 | 357.8 | 0.48 |  |  |
| 63.14 | 359.6 | 0.50 | 8 n |  |

To the slow increase in angle corresponds $d p=$ 0!'0031.

| 1865 |  |  | $+3^{\circ} 339$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}_{25} \mathrm{~m}_{1}}$ | $+4^{\circ} 12^{\prime}$ |  | $9.3-9.5$ | K 5 |

Van den Bos ' orbit (1962) gives the residuals $-0: 6$ and $+0!09$.

| Kui 8 |  |  | $+1{ }^{\circ} 431$ |
| :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}} 25^{\mathrm{m}} .4$ | $+1^{\circ} 44^{\prime}$ |  | 7.1-7 |
| 59.656 | 28:8 | 0.145 |  |
| 60.618 | 28.3 | 0.46 |  |
| 60.774 | 32.7 | 0.49 |  |
| 60.870 | 32.2 | 0.42 |  |
| 65.133 | 29.3 | 0.40 |  |
| 65.136 | $\underline{29.6}$ | 0.45 |  |
| 62.03 | 30.2 | 0.44 | 6n |

Both angle and distance have increased since 1939 making $\mathrm{dp}=0!005$.


The quadrant remained uncertain.

| 1942 Es 1611 |  |  | $+39^{\circ} 563$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}_{30} \mathrm{~m}_{1}}$ | +40 $0^{\prime}$ |  | 10.7-10.9 | K0 |
| 59.638 | 107:3 | 1.42 |  |  |
| 59.933 | 108.0 | 1.48 |  |  |
| 59.79 | 107.6 | 1.45 | 2n |  |

Change questionable after 42 years.


| 62.940 | 280.5 | 2.49 |  |
| ---: | ---: | ---: | ---: |
| 62.959 | 278.9 | 2.41 |  |
| 64.090 | 280.8 | 2.28 |  |
| 64.730 | 278.9 | 2.25 |  |
| 64.740 | $\underline{277.1}$ | $\underline{2.28}$ |  |
| 63.18 | 279.0 | 2.33 | $6 n$ |
| Slow increase in angle. |  |  |  |


| $2^{\mathrm{h}} 31 . \mathrm{m}_{3}$ | +52 ${ }^{\circ} 33^{\prime}$ |  |
| :---: | :---: | :---: |
| 59.638 | 83.6 | 0'.67 |
| 59.933 | 79.7 | 0.71 |
| 59.79 | 81.6 | 0.69 |

To the slow change in angle corresponds $d p=$ 0!0026.


No certain change in distance but the angle has increased by $70^{\circ}$ since 1900 making $\mathrm{dp}=$ $0!0025$.

| 1992 | A 1278 |  |  | $+45^{\circ} 641$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | $2^{\mathrm{h}} 35 . \mathrm{m} .0$ | $+45^{\circ} 51^{\prime}$ |  | 8.9-9.0 | F5 |
|  | 59.151 | $264: 7$ | 0.112 |  |  |
|  | 59.964 | 268.8 | 0.11 |  |  |
|  | 59.56 | 266.8 | 0.12 |  |  |
| Both pair | angle and to watch | distance | have | decreased | A |



| 61.664 | 64.3 | $2: 87$ |  |
| :--- | :--- | :--- | :--- |
| 61.771 | 64.2 | 2.88 |  |
| 61.811 | 63.9 | 2.77 |  |
| 61.820 | 63.2 | 3.01 |  |
| $\frac{61.921}{61.80}$ | $\underline{63.7}$ | $\frac{2.84}{23.9}$ | 2.87 |
| $5 n$ |  |  |  |

No certain change in distance but the slow increase in angle makes $d p=0!011$.


No change in 52 years.
$\underline{2034} O \sum 43$
$+25^{\circ} 436$
$2^{\mathrm{h}} 37^{\mathrm{m}} .8$
8.3-9.9 F5

| 60.774 | 23.4 | 0.199 |  |
| :--- | :--- | :--- | :--- |
| 60.870 | 24.7 | 0.95 |  |
| 61.664 | 20.8 | 0.99 |  |
| 61.743 | 20.7 | 1.11 |  |
| $\frac{61.820}{61.37}$ | $\frac{23.4}{22.6}$ | $\frac{1.03}{1.01}$ |  |
| $5 n$ |  |  |  |

Heintz' orbit (1961) gives the residuals $+4: 9$ and -0.06.

2051

| Hu 539 |  | $+48^{\circ} 737$ |
| :---: | :---: | :---: |
| $2^{\mathrm{h}} 38^{\mathrm{m} .9} 9$ | $+49^{\circ} 12^{\prime}$ | $9.2-9.4 \quad \mathrm{F2}$ |


| 58.506 | 48.1 | 0.130 |  |
| :--- | :--- | :--- | :--- |
| 62.701 | 49.6 | 0.29 |  |
| 62.704 | 46.8 | 0.27 |  |
| 64.090 | $\frac{42.8}{62.00}$ | 46.8 | 0.34 |
| 0.30 | $4 \pi$ |  |  |

No change in distance but to the decrease in angle corresponds $\mathrm{dp}=0!0022$.

2063 Es 1812
$+60^{\circ} 556$
$2^{\mathrm{h}_{39} \mathrm{~m}_{7}}$
$+60^{\circ} 43^{\prime}$

| 59.638 | 90.8 | 2.11 |  |
| :--- | :--- | :--- | :--- |
| 61.842 | 88.7 | 2.22 |  |
| 62.701 | 89.8 | 2.31 |  |
| 62.704 | 88.2 | 2.21 |  |
| 62.940 | $\underline{89.7}$ | $\underline{2.22}$ |  |
| 61.96 | 89.4 | 2.21 | $5 n$ |

Little change in 40 years.


The increase in angle gives a $d p=0!014$.



The angle has hardly changed in 130 years but to the increase in distance corresponds dp 0":017.

2236
A 2413
$+1^{\circ} 515$

| $2^{h_{54}{ }^{m} .6}$ | $+1{ }^{\circ} 41$ |  | $8.5-8$ |
| :---: | :---: | :---: | :---: |
| 61.664 | 39.5 | 0.47 |  |
| 61.743 | 35.5 | 0.47 |  |
| 61.779 | 34.6 | 0.53 |  |
| 61.820 | 39.5 | 0.48 |  |
| 61.921 | 39.3 | $\underline{0.49}$ |  |
| 61.79 | 37.7 | 0.49 | $5 n$ |

Muller's orbit (1952) gives the residuals $+1: 5$ and $+0!.05$.


The residuals from Baize's orbit (1958) are $+1: 7$ and +0.08 .

| 2279 OS 49 |  |  | $+17^{\circ} 471$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $2^{\mathrm{h}_{57} \mathrm{~m}_{7}}$ | +17 ${ }^{\circ} 49^{\prime}$ |  | 7.0-10.0 | A0 |
| 61.664 | 57.2 | 2.00 |  |  |
| 61.718 | 58.2 | 2.06 |  |  |
| 61.743 | 55.8 | 2.03 |  |  |
| 61.779 | 57.7 | 1.99 |  |  |
| 61.73 | 57.2 | 2.02 | $4 n$ |  |

To the slow decrease in angle and increase in distance corresponds $d p=0!.009$.

| 2373 A 2030 |  | $+4^{\circ} 501$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $3^{\mathrm{h}} 7^{\mathrm{m}} 0$ | $+5^{\circ} 0^{\prime}$ |  | 9.4-9.4 | G0 |
| 60.774 | 349.2 | $0 . .33$ |  |  |
| 62.701 | 348.2 | 0.33 |  |  |
| 62.940 | 352.5 | 0.32 |  |  |
| 62.14 | 350.0 | 0.33 |  |  |
| Couteau's orbit and $+0!01$. | (1962) | Ves | he resi | als | and +0.01 .


| 2377 OE 50 |  |  | $+70^{\circ} 230$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $3^{\mathrm{h}} 7^{m} 6$ | +71 ${ }^{\circ} 22^{\prime}$ |  | 8.5-8.5 | F8 |
| 61.220 | 182.2 | 11.48 |  |  |
| 62.707 | 181.6 | 1.39 |  |  |
| $\underline{62.959}$ | $\underline{179.8}$ | 1.34 |  |  |
| 62.30 | 181.2 | 1.40 | $3 n$ |  |
| The longer arc | makes dp | 0.0 |  |  |


| 2484 |  |  |  |  |
| ---: | :--- | :--- | :--- | :--- |
| Ho 320 |  | $+0^{\circ} 570$ |  |  |
| $3^{h} 18^{\mathrm{m}_{3}}$ | $+0^{\circ} 59^{\prime}$ | $8.0-10.5$ |  | GO |
| 61.718 | 171.7 | 2.40 |  |  |
| 61.743 | 171.9 | 2.26 |  |  |
| $\frac{61.779}{61.75}$ | $\frac{170.5}{171.4}$ | $\frac{2.20}{2.29}$ | 3 n |  |

The 0.15 yearly proper motion of the bright component establishes the physical connection. To the increase in distance corresponds $d p=0!031$.

| $\underline{2491} \sum 380$ |  |  | $+8^{\circ} 500$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $3^{\mathrm{h}} 19 . \mathrm{m}_{0}$ | $+8^{\circ} 35^{\prime}$ |  | 8.7-8.7 | G0 |
| 59.638 | $39: 7$ | 1!16 |  |  |
| 59.964 | 33.4 | 1.17 |  |  |
| 59.967 | 36.0 | 1.09 |  |  |
| 59.86 | 36.4 | 1.14 | 3 n |  |

The longer arc again makes $\mathrm{dp}=0.0014$.

2525 Ho 321
$+44^{\circ} 695$
$3^{\mathrm{h}} 22^{\mathrm{m}} \mathrm{m}_{2}+45^{\circ} 20^{\prime} \quad 7.7-10.2 \quad \mathrm{~B} 8$

| 61.220 | $12: 3$ | 0.74 |  |
| :--- | :--- | :--- | :--- |
| 61.811 | 11.4 | 0.64 |  |
| 62.704 | 10.8 | 0.61 |  |
| 62.940 | 14.4 | 0.56 |  |
| 62.948 | 15.2 | 0.57 |  |
| $\frac{64.730}{62.73}$ | 10.7 | $\underline{0.66}$ |  |
| 12.5 | 0.63 | 6n |  |

Slow increase in both coordinates making $\mathrm{dp}=$ 0!.011.

\[

\]

The residuals from Danjon's orbit (1938) are $+3: 1$ and -0.104 .

| Kui 15 |  |
| :---: | :---: |
| $3^{\mathrm{h}} 49^{\mathrm{m}} .3$ | $+6^{\circ} 594$ |
|  | $+6^{\circ} 23^{\prime}$ |
| $6.3-6.4 \quad$ B9 |  |


| 60.192 | 211.4 | 0.49 |  |
| :--- | :--- | :--- | :--- |
| 60.774 | 215.3 | 0.53 |  |
| 65.133 | 209.2 | 0.57 |  |
| 65.136 | 215.2 | 0.50 |  |
| $\underline{65.140}$ | $\underline{209.5}$ | $\underline{0.51}$ |  |
| 63.28 | 212.1 | 0.52 | $5 n$ |

The angle of 31 Tau has changed very little since 1930 but there is an increase in the distance.

| 2828 A 1293 |  |  | +52 ${ }^{\circ} 722$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $3^{\text {h/ }} 49{ }^{\text {m. }} 9$ | +53 ${ }^{\circ} 8^{\prime}$ |  | 8.5-8.9 | G0 |
| 64.090 65.040 | $\begin{aligned} & 214: 5 \\ & 215.6 \end{aligned}$ | $\begin{aligned} & 0.34 \\ & 0.38 \end{aligned}$ |  |  |
| 64.56 | 215.0 | 0.36 | 2n |  |

The residuals from Couteau's orbit (1962) are $+5: 1$ and $+0!04$.

2911 Hu 27
$+9^{\circ} 523$

| $3^{h_{56}} \mathrm{~m}_{3}$ | $+9^{\circ} 38^{\prime}$ |  |
| :---: | :---: | :---: |
| 57.992 | 261.1 | 0.126 |
| 57.995 | 262.7 | 0.33 |
| 58.020 | 265.6 | 0.29 |
| 59.964 | 259.2 | 0.27 |
| 59.967 | 259.2 | 0.28 |
| $\underline{60.186}$ | $\underline{260.6}$ | 0.26 |
| 59.02 | 261.4 | 0.28 |


| $\underline{2959} \sum 483$ |  |  | $+39^{\circ} 918$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $4^{\text {h }} 0{ }^{\text {m }} 7$ | $+39^{\circ} 23^{\prime}$ |  | 7.4-8.9 | G5 |
| $\begin{aligned} & 64.090 \\ & 65.040 \\ & \hline \end{aligned}$ | $\begin{aligned} & 104^{\circ} 2 \\ & 102.2 \\ & \hline \end{aligned}$ | $\begin{aligned} & 0!84 \\ & 0.65 \\ & \hline \end{aligned}$ |  |  |
| Th 64.56 <br> The residuals <br> -0.1 and +0. | $\overline{103.2}$ <br> rom Cout |  |  | 58) |



| 62.940 | 3.5 | $0!31$ |  |
| :--- | :--- | :--- | :--- |
| 62.948 | 7.7 | 0.34 |  |
| $\underline{62.959}$ | $\underline{8.2}$ | 0.29 |  |
| 62.26 | 6.4 | 0.32 | 4 n |

The residuals from Baize's orbit (1961) are $+12: 1$ and +0.401 .

| 3041 A 2801 |  | $-5^{\circ} 841$ |  |
| :---: | :---: | :---: | :---: |
| $4^{\mathrm{h}_{8} \mathrm{~m}_{2}}$ | $-5^{\circ} 0^{\prime}$ | $8.3-8.3 \quad \mathrm{GO}$ |  |
| 59.967 | 12.1 | 0.19 |  |
| $\frac{60.192}{60.08}$ | $\frac{11.8}{12.0}$ | $\frac{0.16}{0.18}$ |  |
|  | 2 n |  |  |

The residuals from Muller's orbit (1954) are $+0: 3$ and -0.03 .

| 3098 ¢ 511 |  |  | $+58^{\circ} 727$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $4^{\text {h }} 13.9$ | +58 ${ }^{\circ} 40^{\prime}$ |  | 7.4-7.9 | A0 |
| 59.967 | 138.4 | 0:38 |  |  |
| 60.198 | 142.4 | 0.36 |  |  |
| 61.220 | 133.7 | 0.38 |  |  |
| 61.256 | 133.9 | 0.43 |  |  |
| 62.228 | $\underline{131.0}$ | 0.38 |  |  |
| 60.97 | 135.9 | 0.39 | 5n |  |

The residuals from Baize's orbit (1958) are $+5: 4$ and +0.02 .


A pair to watch as it closes in.
 Muller's orbit (1963) gives the residuals +0.6 and +0.13 .

| $3174 \sum 535$ |  |  | $+11^{\circ} 601$ |
| :---: | :---: | :---: | :---: |
| $4^{\text {h }} 20{ }^{\text {m }}$. ${ }^{\text {a }}$ | $+11^{\circ} 16^{\prime}$ |  | 7.1-8.6 A2 |
| 60.774 | 302:9 | 1!26 |  |
| 61.080 | 303.3 | 1.40 |  |
| 61.094 | 305.3 | 1.25 |  |
| 61.097 | 305.2 | 1.22 |  |
| 61.108 | 305.1 | 1.22 |  |
| 61.03 | 304 | 1.27 |  |

Long period binary. The arc described since 1831 makes $\mathrm{dp}=0!018$.

| $4^{\text {h }} 22^{\text {ma }} 9$ | $+18^{\circ} 45^{\prime}$ |  | 8.2-8.8 G0 |
| :---: | :---: | :---: | :---: |
| 59.157 | $221: 6$ | 0'12 |  |
| 59.967 | 219.1 | 0.15 |  |
| 60.186 | 216.9 | 0.15 |  |
| 60.192 | $\underline{214.7}$ | 0.20 |  |
| 59.88 | 218.1 | 0.15 | 4 n |

Wierzbinski's orbit (1962) gives for that date which shows that his period is too long.


| 62.948 | 180.8 | 0.29 |  |
| :--- | :--- | :--- | :--- |
| $\frac{62.959}{62.95}$ | $\frac{177.6}{179.2}$ | $\underline{0.27}$ |  |

Baize's second orbit (1956) leaves the residuals $+6: 7$ and 0.00 .

| 3370 Hu 442 |  |  | +22 ${ }^{\circ} 728$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $4^{\text {h }} 37^{\text {m }}$. 0 | +22 ${ }^{\circ} 55^{\prime}$ |  | 9.3-9.9 | A2 |
| 59.967 | 11:0 | 0'39 |  |  |
| 60.186 | 12.8 | 0.39 |  |  |
| 60.774 | 16.3 | 0.35 |  |  |
| 62.948 | 15.8 | 0.34 |  |  |
| 62.959 | 11.1 | 0.40 |  |  |
| 61.37 | 13.4 | 0.37 | 5n |  |

The longer arc makes $d p=0!.008$.
$3390 \quad \sum 577$
$+37^{\circ} 957$
$4^{\mathrm{h}} 38^{\mathrm{m} \cdot} 8$

| 64.740 | 34.7 | 1.32 |  |
| :--- | :--- | :--- | :--- |
| 64.905 | 31.8 | 1.31 |  |
| 64.947 | 34.4 | 1.32 |  |
| $\underline{65.037}$ | $\underline{31.5}$ | $\underline{1.35}$ |  |
| 64.91 | 33.1 | 1.32 | 4 n |

The longer arc makes $d p=0.1014$.
The residuals from Popovic's premature orbit (1964) are +1.1 and +0.06 . Based on an arc of only $60^{\circ}$ it is still completely indeterminate.
3391 A 1013
$+59^{\circ} 826$

| $4^{\mathrm{h}} 39^{m} .0$ | $+59^{\circ} 26^{\prime}$ |  | $7.3-7.3$ | A3 |
| :--- | :--- | :--- | :--- | :--- |
| 60.190 | 356.5 | 0.26 |  |  |
| $\underline{60.198}$ | $\frac{355.3}{355.9}$ | $\underline{0.26}$ |  |  |
| 60.194 | 355.9 | $2 n$ |  |  |

The angular velocity is speeding up as the distance decreases.

$$
\begin{aligned}
& 3475 \beta 883 \\
& +10^{\circ} 654 \\
& 4^{\mathrm{h}} 48 . \mathrm{m}_{4} \\
& +10^{\circ} 59^{\prime} \\
& \text { 7.8-7.8 F5 } \\
& \begin{array}{llll}
65.037 & 60.7 & 0.30 & \\
65.130 & 56.2 & 0.26 & \\
\underline{65.140} & \underline{59.2} & \underline{0.26} & \\
65.10 & 58.7 & 0.27 & 3 n
\end{array}
\end{aligned}
$$

Wierzbinski's orbit (1961) leaves the residuals $+4: 6$ and -0.01 .

3476 Hu 553
$+51^{\circ} 985$
$4^{\text {h }} 48^{\text {m. }} .5$
$+51^{\circ} 19^{\prime}$
9.1-11.3

| 62.228 | $84: 3$ | 3.34 |  |
| :--- | :--- | :--- | :--- |
| 62.899 | 83.6 | 3.11 |  |
| 62.940 | 81.8 | 3.16 |  |
| 62.948 | 82.3 | 3.26 |  |
| 64.730 | 80.3 | 3.15 |  |
| $\underline{64.740}$ | $\underline{83.8}$ | $\underline{3.04}$ |  |
| 63.31 | 82.7 | 3.18 | $6 n$ |

No definite change in 60 years. The 0 ":034
proper motion of the bright component proves the physical connection.

| 3481 Hu 554 |  |  | $+49^{\circ} 1262$ |
| :---: | :---: | :---: | :---: |
| $4^{\mathrm{h}} 49^{\text {m }} 0$ | $+49^{\circ} 58^{\prime}$ |  | 9.4-10.9 |
| 62.289 | $307: 7$ | 2".08 |  |
| 62.940 | 310.7 | 2.07 |  |
| 62.948 | 309.8 | 1.86 |  |
| 62.950 | 310.5 | 2.28 |  |
| 62.953 | 308.2 | 2.22 |  |
| 64.730 | 310.8 | 2.02 |  |
| 64.947 | 310.1 | 1.92 |  |
| 63.39 | 309.7 | 2.06 | 7 n |
| No definite c | nge in 60 | years |  |

No definite change in 60 years.


To the very slow change corresponds $d p=0!0018$.

| 3535 Kr 25 |  |  | +56 ${ }^{\circ} 988$ |
| :---: | :---: | :---: | :---: |
| $4^{\text {h }}$ 53 ${ }^{\text {m }} 3$ | $+56^{\circ} 34^{\prime}$ |  | 9.7-10.2 |
| 61.201 | $118: 8$ | 3.30 |  |
| 61.204 | 119.1 | 3.11 |  |
| 61.220 | 121.8 | 3.00 |  |
| 61.245 | 119.3 | 3.21 |  |
| 61.253 | 119.0 | 3.06 |  |
| 61.256 | $\underline{120.1}$ | 3.04 |  |
| 61.23 | 119.7 | 3.12 | 6 n |
| Very slow change. |  |  |  |
| $35420 \sum 91$ |  | $+2^{\circ} 818$ |  |
| $4^{\text {h }}$ 53. ${ }^{\text {m }} 6$ | $+3^{\circ} 6^{\prime}$ |  | 8.5-9.0 в9 |
| 61.080 | 234:8 | 0".55 |  |
| 61.097 | $\underline{234.6}$ | 0.66 |  |
| 61.09 | 234.7 | 0.60 | 2 n |

The longer arc reduces $d p$ to $0!\cdot 003$.
3591 J 47
$+0^{\circ} 913$

| $4^{\text {h }} 57{ }^{\text {m }} 3$ | $+0^{\circ} 26^{\prime}$ | 10.6-10.9 |
| :---: | :---: | :---: |
| 65.130 | 302.1 | 4:36 |
| 65.133 | 300.5 | 4.36 |
| 65.136 | 300.3 | 4.36 |
| 65.13 | 301.0 | 4.39 3n |

Van den Bos' uniform linear motion gives the residuals +2.9 and -0.11 .

3602

- 1238
$+26^{\circ} 774$

$4^{h_{58} .{ }^{m}} 8$
65.130
65.133 65.13
$+20^{\circ} 46^{\prime}$
8.6-8.9
$\begin{array}{llll} & 243.8 & 0.35 & 2 n\end{array}$
Comparison with two orbits give the residuals:
Kumaritz (1956) $+4: 7 \quad-0.02$
Arend (1959) $-1.6 \quad-0.05$


## 3622 J 240

$$
4^{\mathrm{h}_{59} \mathrm{~m}_{4}} 4 \quad+35^{\circ} 1^{\prime} \quad 10.0-10.2
$$

| 60.198 | $204: 1$ | $1!57$ |  |
| :--- | :--- | :--- | :--- |
| 62.950 | 209.3 | 1.54 |  |
| 64.947 | 209.6 | 1.42 |  |
| 65.130 | 210.9 | 1.63 |  |
| 65.133 | 206.5 | 1.56 |  |
| $\underline{65.136}$ | $\underline{205.7}$ | $\underline{1.60}$ |  |
| 63.92 | 207.7 | 1.55 | $6 n$ |

No change in angle but distance increased.

$$
\begin{array}{ccc}
3658 & \sum 615 & +73^{\circ} 271 \\
5^{\mathrm{h}} 1^{\mathrm{m} .8} 8 & +73^{\circ} 32^{\prime} & 8.2-10.0
\end{array}
$$

| 61.256 | 7.3 | 11.34 |  |
| :--- | :--- | :--- | :--- |
| 61.265 | 7.6 | 1.42 |  |
| 62.228 | 6.8 | 1.37 |  |
| 62.288 | 7.7 | 1.34 |  |
| 62.964 | $\frac{5.6}{7.0}$ | $\underline{1.46}$ |  |
| 62.00 | 7.39 | $5 n$ |  |

The longer arc reduces $d p$ to $0!006$.

3678 Hu 1095

$$
+39^{\circ} 1169
$$

$5^{h_{3}} 3^{m} .0$ $+39^{\circ} 58^{\prime}$
8.1-9.3 A0

| 61.204 | 17.4 | $0!.35$ |  |
| :--- | :--- | :--- | :--- |
| 61.265 | 19.4 | 0.32 |  |
| $\underline{62.285}$ | $\underline{19.5}$ | $\underline{0.36}$ |  |
| 61.58 | 18.8 | 0.34 | $3 n$ |

Very slow increase in angle corresponding to $\mathrm{dp}=0$ 0'0016.
$3689 \sum 635$
$+54^{\circ} 862$
$5^{h_{3}{ }^{(1)} 8}$
$+54^{\circ} 55^{\prime}$
8.7-8.7 B9

| 61.097 | 302.7 | 0.182 |  |
| :--- | :--- | :--- | :--- |
| 61.201 | 298.7 | 0.98 |  |
| 61.204 | 299.3 | 1.02 |  |
| 61.220 | 302.7 | 0.87 |  |
| 61.245 | 29.7 | 1.02 |  |
| 61.253 | 302.3 | 0.96 |  |
| $\underline{61.256}$ | $\underline{300.5}$ | $\underline{1.03}$ |  |
| 61.21 | 300.8 | 0.96 | 7 n |

The slow change in 131 years shows a $d p=0.0066$.
3697 J 14

| $5^{\mathrm{h}} 4 . \mathrm{m}_{3}$ | $+27^{\circ}{ }^{\prime}$ |  | 9.6-9.9 |
| :---: | :---: | :---: | :---: |
| 61.097 | 229.5 | 2'.90 |  |
| 61.220 | 233.2 | 2.73 |  |
| 65.130 | 233.8 | 2.80 |  |
| 62.48 | 232.2 | 2.81 | $3 n$ |

Motion slow if any. Identified as Oxf. ph. $+27^{\circ} 10488$ and $+28^{\circ} 9407$ which gives the corrected declination.


[^0]| $5^{\text {h }} 10.9$ | $+1^{\circ} 55^{\prime}$ |  | 6.9-7.1 A2 |
| :---: | :---: | :---: | :---: |
| 64.905 | 232:6 | 0'.43 |  |
| 64.947 | 233.8 | 0.39 |  |
| 65.130 | 231.8 | 0.38 |  |
| $\underline{65.140}$ | $\underline{228.2}$ | 0.42 |  |
| 65.03 | 231.6 | 0.40 | 4n |

The residuals from Vanden Bos' orbit (1959) are +1.8 and +0.01 .

3816

| $5^{\mathrm{h}} 11^{\mathrm{m}} .6$ | $+31{ }^{\circ} \mathrm{K}^{\prime}$ |  | $9.5-9.8$ |
| :--- | :--- | :--- | :--- |
| 61.204 | 298.2 | $2{ }^{\prime} .68$ |  |
| 61.217 | 299.5 | 2.58 |  |
| 61.220 | 299.4 | 2.54 |  |
| $\frac{61.237}{61.22}$ | $\frac{295.3}{298.1}$ | $\frac{2.44}{2.56}$ | $4 n$ |

Change questionable. Identified in the 0xford Zones as $+31^{\circ} 10728$ and $+32^{\circ} 17365$.

| 3844 J 48 |  |  | $+1^{\circ} 940$ |
| :---: | :---: | :---: | :---: |
| $5^{\text {h }} 13^{\text {m, }} 1$ | $+1^{\circ} 10^{\prime}$ |  | 0.3-10.7 |
| $\begin{array}{r} 62.940 \\ 64.905 \\ \hline \end{array}$ | $\begin{array}{r} 3896 \\ 37.7 \\ \hline \end{array}$ | $\begin{array}{r} 2!35 \\ 2.21 \end{array}$ |  |
| 63.92 | 38.2 | 2.28 | 2n |

Slow increase in angle.
Capella H (Stearns 3)

| $5^{\text {h }} 13^{\text {m/ }} 7$ | +4547 ${ }^{\circ}$ |  | 10.5-10.5 |
| :---: | :---: | :---: | :---: |
| 59.964 | 145:2 | 3.13 |  |
| 59.967 | 143.4 | 3.18 |  |
| 60.198 | 143.1 | 2.94 |  |
| 65.037 | 152.0 | 3.22 |  |
| 65.040 | 152.4 | 3.34 |  |
| 65.130 | 149.9 | 3.19 |  |
| 65.133 | 151.2 | 3.18 |  |
| $\underline{65.209}$ | $\underline{149.7}$ | 3.30 |  |
| 60.04 | 143.9 | 3.08 | 3 n |
| 65.11 | 151.0 | 3.25 | 5 n |

The orbital motion gives a $\mathrm{dp}=0$ ".086 which confirms the physical connection with Capella.



| 3959 A 2641 |  |  | $+2^{\circ} 934$ |
| :---: | :---: | :---: | :---: |
| $5^{\text {h }} 20.0$ | + $2^{\circ} 34^{\prime}$ |  | 8.4-10.9 G5 |
| 59.137 | $345: 2$ | 1 1.05 |  |
| 59.157 | 342.1 | 0.99 |  |
| 60.186 | 335.0 | 0.89 |  |
| 62.940 | 330.4 | 1.10 |  |
| 60.35 | 338.1 |  |  |
| Residuals from | Baize's |  | (1958) are |

3982
A 846
$+74^{\circ} 241$
$5^{h_{2}} 1_{1}^{m_{1}}$

| 61.240 | 342.6 | 1.28 |  |
| :--- | :--- | :--- | :--- |
| 61.256 | 346.1 | 1.16 |  |
| 62.203 | 343.0 | 1.13 |  |
| 61.57 | 343.9 | 1.19 | $3 n$ |

No definite change after 57 years.

| 4020 A 848 |  |  | -0 $0^{\circ} 945$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $5^{\mathrm{h}} 23 . \mathrm{m}$ | $-0^{\circ} 35^{\prime}$ |  | 6.7-7.3 | B9 |
| 60.186 | $137: 3$ | 0.17 |  |  |
| 60.198 | 139.4 | 0.14 |  |  |
| 60.19 | 138.4 | 0.16 | 2n |  |

The angle increased by $100^{\circ}$ since 1904 and the distance diminished.

4032 Ho 226
$+27^{\circ} 771$

$$
\begin{array}{llll}
5^{\mathrm{h}_{23} 3^{\mathrm{m}} .9} & +27^{\circ} 34^{\prime} & & 8.6 \\
& & & \\
61.097 & 256: 6 & 0.73 & \\
61.100 & 255.1 & 0.67 & \\
\frac{61.152}{61.12} & \frac{256.4}{256.0} & \frac{0.72}{0.71} & 3 n
\end{array}
$$

8.6-8.6 F8

The longer arc makes $d p=0.007$.
4097 之 725
$-1^{\circ} 913$

The proper motion of 31 Orionis is negligible in right ascension but it is 0 '. 022 south according to Boss or 0.012 according to the Yale Zone. If the companion did not belong to the bright star this p.m. would have decreased the angle by 13 and 7 degrees, respectively, since Struve's measure in 1829. That there is no change proves the physical connection.

| 4166 - 1267 |  |  | $+30^{\circ} 942$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $5^{\mathrm{h}} 31{ }^{\mathrm{m}} .9$ | $+30^{\circ} 54^{\prime}$ |  | 8.8-8.8 | F5 |
| 61.201 | 206:3 | 0'.65 |  |  |
| 61.204 | 206.8 | 0.76 |  |  |
| 61.220 | $\underline{204.7}$ | 0.62 |  |  |
| 61.21 | 205.9 | 0.68 | 3 n |  |

Slow orbital motion corresponding to $d p=$ 0.005

$$
\begin{array}{ccc}
4180 & & \\
5^{\mathrm{h}} 32^{\mathrm{m}} \mathrm{~m}_{4} & -0^{\circ} 9^{\prime} & \\
61.00^{\circ} 1005 \\
61.152 & 285.1 & 10^{\prime \prime} .12 \\
6^{\circ} .168 & 286.2 & 10.05 \\
61.217 & 286.4 & 10.38 \\
\frac{61.237}{61.19} & \underline{285.8} & \frac{10.28}{285.9} \\
10.21 & 4 \mathrm{n}
\end{array}
$$

The proper motion is 0.018 according to Boss and 0.008 in the Yale catalogue. If the companion were a background star this would have changed the angle by 11 and 5 degrees, respectively, since Struve's first measure in 1831. The angle is practically unchanged, hence this is a physical pair.

| 4224 |  | $+8^{\circ} 1019$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $5^{\mathrm{h}_{35} 5^{\mathrm{m}} 2}$ | $+8^{\circ}{ }^{\circ} 55^{\prime}$ |  | $8.7-9.5$ | GO |
|  |  |  |  |  |
| 61.265 | 295.3 | 0.59 |  |  |
| 61.811 | 295.6 | 0.55 |  |  |
| 62.940 | 288.7 | 0.56 |  |  |
| 62.964 | 291.2 | 0.76 |  |  |

$\frac{64.090}{62.61} \quad \frac{292: 9}{292.7} \quad \frac{0.60}{0.61} \quad 5 n$

The slow change over the 48 year interval since the discovery makes $\mathrm{dp}=0!007$.

$+1: 2$ and $+0!09$.

| 4243 |  |  | $+37^{\circ} 1277$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $5^{h_{36} m_{5}} 112$ | $+37^{\circ} 56^{\prime}$ |  | $7.8-8.5$ | $B 9$ |
| 60.186 | 49.6 | 0.161 |  |  |
| 60.192 | 54.3 | 0.67 |  |  |
| 60.198 | $\underline{53.3}$ | $\underline{0.64}$ |  |  |
| 60.19 | 52.4 | 0.64 | $3 n$ |  |

The longer arc reduces dp to $0!0030$.


The residuals from Baize's orbit (1961) are +3.5 and +0.06 .

| 4313 Hu 1110 |  |  | $+37^{\circ} 1306$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $5^{h_{4}} 1^{\text {m }}$ 4 | +37 ${ }^{\circ} 33^{\prime}$ |  | 8.5-11.7 | F0 |
| 61.237 | $240: 3$ | 1'47 |  |  |
| 61.256 | 239.3 | 1.25 |  |  |
| 61.259 | 236.9 | 1.38 |  |  |
| 61.265 | 238.2 | 1.47 |  |  |
| 61.25 | 238.7 | 1.39 | 4n |  |

The longer arc reduces dp to $0: 015$.

4370 J 35
$+6^{\circ} 1021$
$5^{\mathrm{h}} 44^{\mathrm{m}} .1$
$+6^{\circ} 22^{\prime}$
10.3-10.5

| 62.940 | 9.8 |
| :--- | ---: |
| 62.953 | 11.9 |
| 62.964 | 8.3 |
| 64.905 | 10.9 |
| 63.44 | 10.2 |

0.91
1.15
1.11

| 1.11 |
| :--- |
| 1.13 |

1.08 4n

Slow orbital motion.

4472 B 1053
$+37^{\circ} 1347$ To the large change in
corresponds $\mathrm{dp}=0.025$.

| 4619 A 120 |  |  | $+25^{\circ} 1089$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $5^{\mathrm{h}} 59.8$ | $+25^{\circ} 53^{\prime}$ |  | 8.3-9.5 | F8 |
| 60.186 | 170:1 | 0'.68 |  |  |
| 61.097 | 171.8 | 0.77 |  |  |
| 61.201 | 169.7 | 0.85 |  |  |
| 61.204 | 170.7 | 0.72 |  |  |
| 61.231 | 170.9 | 0.64 |  |  |
| 61.234 | 172.9 | 0.78 |  |  |
| 61.03 | 171.0 | 0.74 | 6n |  |

Kui: 23
1 Gem


To the slow decrease in angle corresponds dp = $0!0017$.

| Kui 24 |  |  | $+17^{\circ} 1182$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $6^{\mathrm{h}} 11{ }^{\text {m }}$ 6 | $+17^{\circ} 55^{\prime}$ |  | 6.5-6.5 | A5 |
| 60.186 | 138:8 | 0:.43 |  |  |
| 60.192 | 142.6 | 0.41 |  |  |
| 60.198 | 136.7 | 0.43 |  |  |
| 60.19 | 139.4 | 0.42 | 3n |  |

Hardly changed since 1934.

| Rst 5225 |  |  | $+1^{\circ} 1275$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $6^{\mathrm{h}} 13 .{ }^{\text {m/3 }}$ | $+1^{\circ} 11^{\prime}$ |  | 7.1-7.1 | F5 |
| 60.198 | 269:5 | 0'. 20 | 1 n |  |
| 4971 A 2667 |  |  | $+2^{\circ} 1197$ |  |
| $6^{\mathrm{h}} 18 . \mathrm{m}$ | $+2^{\circ} 18^{\prime}$ |  | 6.5-6.8 A5 |  |
| 64.090 65.130 | 133.0 130.8 | 0.47 <br> 0.42 <br> 0 |  |  |
| 65.130 | 130.8 | 0.42 |  |  |
| 64.61 | 131.9 | 0.45 | the residuals |  |
| Lategan's orbit $+1: 7$ and -0.01 . | (1961) | gives t |  |  |
| 5042 O§ 139 |  |  | $+22^{\circ} 1323$ |  |
| $6^{\mathrm{h}} 22{ }^{\text {m/ }} 6$ | $+22^{\circ} 29^{1}$ |  | 7.8-10.3 | A3 |
| 59.151 | 238:2 | $0 \% 39$ |  |  |
| 60.198 | 242.6 | 0.41 |  |  |
| 61.234 | 239.2 | 0.37 |  |  |
| 61.265 | 237.2 | 0.44 |  |  |
| 60.46 | 239.3 | 0.40 | $4 n$ |  |

Heintz's orbit (1962) gives the residuals $+5: 6$ and $+0!04$.

| $5054 \beta 1191$ |  | $+18^{\circ} 1214$ |  |  |
| :---: | :--- | :--- | :--- | :--- |
| $6^{\mathrm{h}} 23^{\mathrm{m}} ._{2}$ | $+18^{\circ} 48^{\prime}$ | $6.9-13.9$ |  | KO |
| 64.090 | 300.8 | $1: 96$ |  |  |
| 64.905 | 309.0 | 1.90 |  |  |
| 65.130 | 302.1 | 2.02 |  |  |
| $\frac{65.140}{64.82}$ | $\frac{306.5}{304.6}$ | $\frac{2.13}{2.00}$ | 4 n |  |

This angle differs by 140 degrees from Burnham's angle in 1890, a motion that leads to a dp of 0":054: There were no intermediate observations. The $0 " 21$ yearly proper motion of the bright component proves the physical connection.

| 5079 J 1092 |  | $+22^{\circ} 1334$ |  |
| :---: | :---: | :---: | :---: |
| $6^{\mathrm{h}_{24} 4^{\mathrm{m}} .8}$ | $+22^{\circ} 53^{\prime}$ |  | $9.5-10.4$ |
| 62.953 | 234.9 | $4^{\prime \prime} .27$ |  |
| 64.905 | 235.1 | 4.30 |  |
| $\frac{65.136}{64.33}$ | $\frac{233.8}{234.4}$ | $\frac{4.46}{4.34}$ | 3 n |

The angle has decreased by 20 degrees since 1915 and the distance has more than doubled. This would correspond to an improbably large value of $\mathrm{dp}=0$ 0'031. Probably optical.

| Wor 6 |  |  | +52 ${ }^{\circ} 1088$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $6^{\mathrm{h}} 28 . \mathrm{m}_{3}$ | +52 ${ }^{\circ} 27^{\prime}$ |  | 10.3-10.4 | M0 |
| 61.234 | 152:2 | 0.98 |  |  |
| 61.240 | 149.3 | 0.84 |  |  |
| 61.256 | 154.5 | 0.91 |  |  |
| 61.259 | 151.5 | 0.90 |  |  |
| 61.265 | 153.4 | 0.79 |  |  |
| 61.25 | 152.2 | 0.88 | $5 n$ |  |
| 5159 A 2817 |  |  | $+7^{\circ} 1327$ |  |
| $6^{\mathrm{h}} 28.07$ | $+7^{\circ} 52^{\prime}$ |  | 9.3-9.3 | G5 |
| 59.967 | 135:8 | 0.14 |  |  |
| 60.172 | 128.3 | 0.14 |  |  |
| 60.07 | 132.0 | 0.14 | 2 n |  |

Popovic's orbit (1964) gives the large residuals $+27: 1$ and $-0!04$.

| 5212 Ho 234 |  |  | $-11^{\circ} 1536$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $6^{\mathrm{h}} 32 . \mathrm{m}_{2}$ | $-11^{\circ} 11^{\prime}$ |  | 8.2-8.2 | F0 |
| 61.811 | 120:9 | 0.34 |  |  |
| 62.964 | 123.5 | 0.36 |  |  |
| 64.090 | 118.4 | 0.38 |  |  |
| 65.037 | 118.0 | 0.35 |  |  |
| 65.130 | 122.4 | 0.36 |  |  |
| 63.81 | 120.6 | 0.36 | 5n |  |
| Baize's orbit | 1958) rev | arses | the qua | an | gives the residuals $-6: 0$ and +0.08 .



Evidently a short period binary but the measures are too scarce to define the orbit.


| 61.234 | $302: 0$ | 0.58 |
| :--- | :--- | :--- |
| $\underline{61.237}$ | $\underline{300.4}$ | $\underline{0.65}$ |
| 61.08 | 301.7 | 0.60 |
| $7 n$ |  |  |

The longer arc reduces dp to 0,008 .
 +23.7 and +0!.02.


| 5444 | J 802 |  |  | $-4^{\circ} 1653$ |
| :---: | :---: | :---: | :---: | :---: |
|  | $6_{4}{ }_{4}{ }^{\text {m }}$. 3 | - $4^{\circ} 11^{\prime}$ |  | 10.9-12.4 |
|  | 62.953 | 124:2 | 4.31 |  |
|  | 62.964 | 124.2 | 4.03 |  |
|  | 65.130 | 123.1 | 4.12 |  |
|  | 65.133 | 122.7 | 3.94 |  |
|  | 64.04 | 123.6 | 4.10 | 4n |
| Slow i | increase | angle. |  |  |


| 5455 |  |  | $+0^{\circ} 1604$ |
| :---: | :---: | :---: | :---: |
| $6_{45} \mathrm{~h}_{4} \mathrm{~m}_{2}$ | $+0^{\circ} 24^{\prime}$ |  | $7.4-7.9 \quad \mathrm{~A} 2$ |

Residuals from Heintz' orbit (1961) are $+7: 9$ and $+0!01$.


| 5524 M1b 119 |  |  | $+59^{\circ} 1029$ |
| :---: | :---: | :---: | :---: |
| $6^{\mathrm{h}} \mathrm{m}^{\mathrm{m}} \mathrm{l}_{1}$ | $+59^{\circ} 14^{\prime}$ |  | 10.6-10.7 |
| 60.148 | $306: 7$ | 2!'93 |  |
| 61.201 | 305.4 | 2.93 |  |
| 61.220 | 306.0 | 2.85 |  |
| 61.231 | 303.9 | 2.86 |  |
| 61.240 | 305.6 | $\underline{2.80}$ |  |
| 61.01 | 305.5 | 2.87 | 5n |

[^1]| $6^{\text {h }} 51{ }^{\text {ma }} 6$ | $+18^{\circ} 21{ }^{\prime}$ |  |
| :---: | :---: | :---: |
| 61.217 | 12.3 | 4:57 |
| 61.220 | 13.2 | 4.39 |
| 61.234 | 13.0 | 4.43 |
| 61.237 | 12.9 | 4.57 4.50 |
| 61.240 | 13.4 | 4.50 |
| 61.23 | 13.0 | 4.49 |

No definite change since 1911 . The star was identified as Par. ph. $+18^{\circ}, 6{ }^{\circ} \mathrm{h}^{\mathrm{m}}{ }^{\mathrm{m}}$, No. 287 .

| Wor 18 |  |  | $+27^{\circ}$ |
| :---: | :---: | :---: | :---: |
| $7{ }^{\text {h }}$ [ ${ }^{\text {m }} 6$ | +27 ${ }^{\circ} 33^{\prime}$ |  | 10.6-13.4 |
| $\begin{array}{r} 62.288 \\ 62.291 \\ \hline \end{array}$ | $\begin{array}{r} 298.0 \\ 297.6 \\ \hline \end{array}$ | $\begin{aligned} & 12!.05 \\ & 12.22 \\ & \hline \end{aligned}$ |  |
| 62.29 | 297.8 | 12.14 | 2n |
| 5841 J 703 |  |  |  |
| $7^{\text {¢ }}{ }^{\text {m }} 7$ | $+15^{\circ} 49^{\prime}$ |  | 9.7-9.7 |
| 62.953 | $118: 9$ | 6!.56 |  |
| 65.037 | 120.6 | 6.55 |  |
| 65.209 | 117.7 | 6.51 |  |
| 64.40 | 119.1 | 6.54 | 3 n |

Slow change in both coordinates.

| 5866 AG - |  |  | $+18^{\circ} 1524$ |
| :---: | :---: | :---: | :---: |
| $7^{\text {h }}$ 9 ${ }^{\text {m }}$. | $+18^{\circ} 44^{\prime}$ |  | 9.1-9.1 |
| 60.186 | $190: 3$ | 0.93 |  |
| 61.097 | 189.8 | 0.96 |  |
| 61.220 | 190.8 | 1.06 |  |
| 61.223 | 191.8 | 0.95 |  |
| 61.231 | 189.8 | 1.02 |  |
| 60.99 | 190.5 | 0.98 | 5n |



To the increase in angle corresponds $d p=$ 0':014.


Change inmaterial.


$6175 \sum 1110$
Castor

| $7^{\text {h }} 31 .{ }^{\text {m }} 4$ | $+32^{\circ} 0^{\prime}$ |  | 2.0-2.8 AO |
| :---: | :---: | :---: | :---: |
| 61.094 | 164:0 | 1'99 |  |
| 61.097 | 164.2 | 1.98 |  |
| 61.152 | 165.3 | 2.03 |  |
| 61.201 | 167.0 | 2.17 |  |
| 61.214 | 166.4 | 2.06 |  |
| 61.231 | 165.8 | 2.22 |  |
| 61.237 | 165.6 | 2.17 |  |
| 61.242 | 166.1 | 2.14 |  |
| 61.245 | 165.5 | 2.07 |  |
| 62.067 | 162.0 | 2.00 |  |
| 62.069 | 163.3 | 1.95 |  |
| 62.203 | 162.7 | 2.05 |  |
| 62.225 | 160.0 | 2.09 |  |
| 62.228 | 162.7 | 2.18 |  |
| 62.285 | 164.4 | 2.09 |  |
| 62.288 | 164.7 | 2.16 |  |
| 62.291 | 160.8 | 2.03 |  |
| 62.953 | 159.3 | 2.05 |  |
| 63.337 | 161.9 | 1.91 |  |
| 63.340 | 160.5 | 1.93 |  |
| 63.351 | 160.9 | 1.96 |  |
| 63.362 | 160.6 | 1.92 |  |
| 64.905 | 153.0 | 2.02 |  |
| 65.037 | 148.2 | 2.11 |  |
| $\underline{65.040}$ | 149.0 | 2.10 |  |
| 61.19 | 165.5 | 2.09 | 9 n |
| 62.21 | 162.6 | 2.07 | 8 n |
| 63.27 | 160.6 | 1.95 | 5 n |
| 64.99 | 150.1 | 2.08 | 3n |

Comparison with Rabe's orbit (1958) gives: $61.19+1.5-0.04$

| 62.21 | +1.8 | 0.00 |
| ---: | ---: | ---: |
| 63.28 | +3.1 | -0.06 |

$\begin{array}{lll}64.99 & -0.9 & +0.14\end{array}$

## 6185 OE 175

$+31^{\circ} 1620$

| $7 \mathrm{~h}_{32 \mathrm{~m}} \mathrm{~m}$ |  | $+31^{\circ} \mathrm{H}^{\prime}$ |  |
| :---: | :---: | :---: | :---: |
|  |  | $5.8-6$ |  |
| 59.964 | 330.5 | 0.37 |  |
| 60.186 | 329.1 | 0.36 |  |
| 60.198 | 330.2 | 0.41 |  |
| 60.201 | 332.3 | 0.36 |  |
| 65.133 | 328.6 | 0.33 |  |
| 65.136 | 329.3 | 6.33 |  |
| 60.14 | 330.5 | 0.38 | 4 n |
| 65.13 | 329.0 | 0.33 | 2 n |

The angle has hardly changed in 118 years but after passing a maximum in the first half of the century the distance is now decreasing rapidly, showing that the orbit will be very elongated.


| 64.905 | 318:7 | 21.43 |  |
| :---: | :---: | :---: | :---: |
| Hardly changed | 318.6 | 2.44 | 4n |
|  | in 47 y |  |  |
| 6263 § 1126 |  |  | $+5^{\circ} 1742$ |
| $7{ }^{\text {h }} 37^{\text {m }}$ 5 | $+5^{\circ} 21^{\prime}$ |  | 6.4-6.7 A0 |
| 61.097 | 158.4 | 1! 11 |  |
| 61.100 | 160.7 | 1.19 |  |
| 61.168 | 159.1 | 1.25 |  |
| 61.223 | 159.8 | 1.19 |  |
| 61.228 | 156.0 | 1.29 |  |
| 61.231 | 161.1 | $\underline{1.19}$ |  |
| 61.17 | 159.2 | 1.20 |  |
| The longer arc | makes dp | $=0.10$ |  |


| 6291 ¢ 1130 |  |  | $+10^{\circ} 1599$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $7^{\mathrm{h}} 39^{\mathrm{m}} 0$ | $+9^{\circ} 49^{\prime}$ |  | 8.7-9.2 | G0 |
| 61.097 | $238: 3$ | 0.141 |  |  |
| 61.237 | 235.9 | 0.36 |  |  |
| 61.240 | 236.3 | 0.41 |  |  |
| 61.256 | 238.0 | 0.37 |  |  |
| 61.259 | 235.6 | 0.35 |  |  |
| 65.133 | 262.8 | 0.30 |  |  |
| 65.140 | 260.9 | 0.32 |  |  |
| 61.22 | 236.8 | 0.38 | 5n |  |
| 65.14 | 261.8 | 0.31 | 2n |  |

The angle has increased by 100 degrees since 1829 and the distance has closed in considerably making $d p=0!0105$. A pair to watch!

| 6307 Es 2158 |  |  | $+37^{\circ} 1757$ |
| :---: | :---: | :---: | :---: |
| $7^{\mathrm{h}} 40 . \mathrm{m}$ | $+37^{\circ} 3^{\prime}$ |  | 9.6-11.1-11.4 |
|  | AC |  |  |
| 65.130 | $331: 8$ | 2! 44 |  |
| 65.133 | 331.9 | 2.70 |  |
| 65.140 | 332.3 | 2.53 |  |
| 65.13 | 332.0 | 2.56 | $3 n$ |
|  | BC (Van | B.) |  |
| 59.137 | 146:9 | 1!20 |  |
| 59.964 | 147.2 | 1.06 |  |
| 60.198 | 145.9 | 1.05 |  |
| 65.130 | 146.1 | 1.20 |  |
| 65.133 | 147.6 | 1.23 |  |
| 65.140 | 146.8 | 1.15 |  |
| 62.45 <br> Unchanged. The from the astrog | 146.8 | 1.15 | $6 n$ |
|  | position | has | en corrected |
|  | raphic ca | talogu |  | from the astrographic catalogue.


| 6354 Hu 1247 |  |  | $+60^{\circ} 1082$ |
| :---: | :---: | :---: | :---: |
| $7 \mathrm{~h}_{4} 3^{\mathrm{m}_{7}}$ | $+60^{\circ} 25^{\prime}$ |  | 7.7-7.7 F5 |
| 62.288 | 308:0 | 0.124 |  |
| 62.305 | 314.5 | 0.24 |  |
| 65.136 | 280.7 | 0.24 |  |
| 65.140 | 278.7 | 0.24 |  |
| 62.30 | 311.2 | 0.24 | 2n |
| 65.14 | 279.7 | 0.24 | 2n |
| $\begin{gathered} \text { Residuals from } \\ 62.30 \\ 65.14 \end{gathered}$ | $\begin{gathered} \text { Baize's } \\ -1.4 \\ -5.1 \end{gathered}$ | $\begin{gathered} \text { orbit } \\ +0.05 \\ +0.01 \end{gathered}$ | (1961) are: |
| 6369 A 1331 |  |  | $+54^{\circ} 1175$ |
| $7^{\mathrm{h}} 44^{\mathrm{m}} 8$ | $+53^{\circ} 48^{\prime}$ |  | 8.6-13.1 |
| 60.201 | 253.6 | 4:74 |  |
| 61.231 | 255.4 | 4.84 |  |
| 61.234 | 254.8 | 4.62 |  |
| 61.256 | 253.4 | 4.69 |  |
| 60.98 | 254.3 | 4.72 | $4 \pi$ |

The 0.55 proper motion establishes the physical connection. The relative motion is mostly in distance making dp as large as 0!.059.

6374 J 418


Slow increase in angle corresponding to $d p=$ 0.007.

| $\underline{6406} \sum 1136$ |  |  | $+65^{\circ} 599$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $7^{\mathrm{h}} 48 . \mathrm{m}$ | $+65^{\circ} 2^{\prime}$ |  | 7.6-11.3 | K3 |
| 61.217 | 218.2 | 5.'69 |  |  |
| 61.228 | 215.9 | 5.90 |  |  |
| 61.231 | 216.0 | 5.62 |  |  |
| 61.234 | 219.9 | 5.72 |  |  |
| 61.242 | 216.0 | 5.90 |  |  |
| 61.23 | 217.2 | 5.77 | 5n |  |

So far there is no clear indication of curvature in the relative path. Probably optical because the big change since 1830 would lead to an improbable $d p=0.108$.


| 64.905 | $283: 4$ | 0.59 |  |
| :--- | :--- | :--- | :--- |
| 65.037 | 279.0 | 0.48 |  |
| 65.140 | $\underline{286.8}$ | $\underline{0.55}$ |  |
| 65.03 | 283.1 | 0.54 | $3 n$ |

Woolley-Symms' orbit (1933) leaves the residuals $-4^{\circ}: 1$ and +0.102 .

$$
\begin{aligned}
& 6428 \text { J } 1100 \\
& 7^{h_{50}} \mathrm{~m}_{4}+17^{\circ} 40^{\prime} \quad 10.0-12.0 \\
& 62.953 \quad 267.7 \quad 4!09 \\
& \frac{65.209}{64.08} \quad \frac{269.0}{268.4} \quad \frac{3.99}{4.04} \\
& \text { Slow increase in both coordinates. }
\end{aligned}
$$

The angular decrease accelerates as the distance diminishes. The corresponding dp is 0!008.

| 6476 J 69 |  |  | $+3^{\circ} 1844$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $7^{h_{53} \mathrm{~m}_{8}}$ | $+3^{\circ} 32^{\prime}$ |  | 9.5-12.1 | A |
| 61.240 | 258.4 | 2..55 |  |  |
| 61.265 | 259.6 | 2.49 |  |  |
| 62.940 | 261.1 | 2.37 |  |  |
| 62.953 | 260.8 | 2.54 |  |  |
| 65.136 | 260.2 | 2.38 |  |  |
| 65.209 | $\underline{259.9}$ | $\underline{2.46}$ |  |  |
| 63.12 | 260.0 | 2.46 | 6n |  |

Slow change in both coordinates indicating a dp = $0!.016$.

6516 ₹ 1165
$+55^{\circ} 1240$
$7^{h_{58} \mathrm{~m}_{2}}$
$+54^{\circ} 46^{\prime}$
8.2-10.5 A0
$60.188 \quad 284: 0 \quad 0.38$
$\frac{65.140}{62.66} \quad \frac{282.8}{283.4} \quad \frac{0.45}{0.42} \quad 2 n$
The longer arc makes $d p=0.1004$.


Slow orbital motion corresponding to $d p=$ 0",0030.

| 6582 A 1971 |  |  | $-0^{\circ} 1904$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $8^{h_{3}{ }^{\text {m }} 5}$ | - $0^{\circ} 38^{\prime}$ |  | 9.1-9.2 | G5 |
| 61.228 | $26: 4$ | 0!.92 |  |  |
| 61.237 | 28.4 | 0.88 |  |  |
| 61.245 | 23.6 | 1.00 |  |  |
| 61.24 | 26.1 | 0.93 | 3n |  |

To the large change in both coordinates corresponds $\mathrm{dp}=0!008$.

| 6616 J 375 |  |  | $+12^{\circ} 1780$ |
| :---: | :---: | :---: | :---: |
| $8{ }^{\mathrm{h}} \mathrm{m}_{1}$ | $+12^{\circ} 21^{\prime}$ |  | 11.3-11.3-14.5 |
|  | AB |  |  |
| 61.231 | 146.2 | 7!26 |  |
| 61.240 | 147.4 | 7.41 |  |
| 61.296 | 147.2 | 7.46 |  |
| 61.24 | 146.9 | 7.38 | 3n |
|  | AC |  |  |
| 61.231 | 348:3 | 13:'96 |  |
| 61.240 | 348.7 | 13.93 |  |
| 61.256 | 347.9 | 14.10 |  |
| 61.24 | 348.3 | 14.00 | 3n |

The marked motion in $A B$ seems to indicate an optical system.

6619 J 734
$+8^{\circ} 1979$

| $8^{\text {h }}{ }^{\text {m. }} 2$ | $+7^{\circ} 58^{\prime}$ |  | 10.7-11.0 |
| :---: | :---: | :---: | :---: |
| 61.231 | 237:0 | 2!.38 |  |
| 61.234 | 238.1 | 2.16 |  |
| 61.237 | 240.4 | 2.26 |  |
| 61.240 | 238.2 | 2.23 |  |
| 61.245 | 239.1 | 2.11 |  |
| 61.24 | 238.6 | 2.23 | 5n |

No definite change.


To the increase in angle corresponds $d p=$ $0: 014$.

| $6650 \times 1196$ |  | $\zeta$ Cancri |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $8^{\mathrm{h}} 9^{\mathrm{m}} 3$ | $+17^{\circ} 48^{\prime}$ |  | 5.6-6.0-6.3 | G2 |
|  | AB |  |  |  |
| $62.940$ | 356:9 | 11.28 |  |  |
| $62.959$ | 358.3 | 1.18 |  |  |
| 62.95 | 357.6 | 1.23 | $2 n$ |  |
|  | AC |  |  |  |
| 62.940 | $85: 5$ | 5:'88 |  |  |
| 62.959 | 84.8 | 5.91 |  |  |
| 62.95 | 85.2 | 5.90 | 2n |  |
| Gasteyer's orb | t (1954) | for $A B$ | gives the | esid- |
| uals $+2: 7$ and | $0.08$ |  | , |  |

6671 B1244 $+2^{\circ} 1904$

| $8^{\mathrm{h}} 11^{\text {m }}$. | $+2^{\circ} 8^{\prime}$ |  | 8.3-8.5 |
| :---: | :---: | :---: | :---: |
| 61.231 | $18: 5$ | 1:01 |  |
| 61.234 | 20.6 | 1.02 |  |
| 61.257 | 20.5 | 0.97 |  |
| 61.240 | 20.1 | 1.03 |  |
| 61.245 | 20.4 | 1.02 |  |
| 61.24 | 20.0 | 1.01 | 5n |
| longer ar | makes dp | $0 \cdot 01$ | 10. |

The longer arc makes dp $=0.010$.
6677 J 377

| $8^{h} 11^{m} .5$ | $+7^{\circ} 7$ | $9.4-10.0$ |  |
| :--- | :--- | :--- | :--- |
| 62.940 | 18.6 | $1!84$ |  |
| $\frac{65.136}{64.04}$ | $\frac{16.0}{17.3}$ | $\frac{2.00}{1.92}$ | 2 n |

Slow decrease in angle.
6718 J 422

| $8^{\mathrm{h}} 14^{\mathrm{m}} .8$ | $-0^{\circ} 47$ |  | $9.6-10.0$ |
| :---: | :---: | :---: | :---: |
|  |  |  |  |
| 61.256 | $76: 0$ | $4: 95$ |  |
| 62.953 | 77.8 | 5.10 |  |
| 63.348 | 76.2 | 5.11 |  |
| $\frac{65.136}{63.17}$ | $\frac{75.5}{76.4}$ | 5.04 |  |
|  | 5.05 | 4 n |  |

Distance increased. $0^{\circ}, 8^{\mathrm{h}} 16^{\mathrm{m}}$ No. 162 and $-1^{\circ}$, $8 \mathrm{hh}^{2}{ }^{\text {mim }}$, No. 179.

```
6719 A 337
\(8^{\mathrm{h}} 14^{\mathrm{m} .8}-5^{\circ} 13^{\prime} \quad 8.4-8.7 \mathrm{~F} 2\)
\(\begin{array}{lll}61.231 & 132: 4 & 0.37 \\ 61.240 & 133.8 & 0.47\end{array}\)
\(61.240 \quad 133.8 \quad 0.42\)
\(\begin{array}{lll}61.256 & 136.7 & 0.38\end{array}\)
\(\frac{61.265}{61.25} \quad \frac{136.3}{134.8} \quad \frac{0.46}{0.41} 4 n\)
Baize's orbit (1960) gives the residuals -3.2 and +0.06 .
```



| Finsen 346 |  |  | $+9^{\circ} 1954$ |
| :---: | :---: | :---: | :---: |
| $8^{\text {h }} 17 .{ }^{\text {m }} 2$ | $+4^{\circ} 6^{\prime}$ |  | 7.1-7.1 G5 |
| 60.198 | 89:7 | 0'.26 |  |
| 62.940 | 90.4 | 0.26 |  |
| 65.037 | 90.4 | 0.24 |  |
| 65.133 | 91.9 | 0.27 |  |
| $\underline{65.140}$ | 87.2 | 0.26 |  |
| 63.69 | 89.9 | 0.26 | 5n |
| No change so far. |  |  |  |
| $\underline{6762} \sum 1216$ |  |  | $-1^{\circ} 2017$ |
| $8^{\mathrm{h}} 18 .{ }^{\text {m }} 8$ | - $1^{\circ} 26{ }^{\prime}$ |  | 6.9-7.6 A0 |
| 61.097 | 265.4 | 0'.51 |  |
| 61.220 | 269.7 | 0.57 |  |
| 61.228 | 270.6 | 0.59 |  |
| 61.231 | 265.3 | 0.63 |  |
| 61.234 | 270.7 | 0.59 |  |
| 62.285 | 266.7 | 0.65 |  |
| 61.38 | 268.1 | 0.59 | 6n |

The residuals from Ekenberg's orbit (1945) are +6.3 and +0.03 .

| 6775 Hu 854 |  |  | $+65^{\circ} 630$ |
| :---: | :---: | :---: | :---: |
| $8{ }^{\text {h }}{ }_{20}{ }^{\text {m }} 1$ | $+65^{\circ} 38{ }^{\prime}$ |  | 9.8-10.1 ко |
| 60.198 | 215:3 | 1".52 |  |
| 60.201 | 214.9 | 1.68 |  |
| 61.201 | 216.3 | 1.63 |  |
| 61.234 | 215.6 | 1.62 |  |
| 61.237 | 213.7 | 1.53 |  |
| 61.240 | 214.6 | 1.65 |  |
| 61.259 | $\underline{214.2}$ | $\underline{1.56}$ |  |
| 60.94 | 214.9 | 1.60 | $7 n$ |

No appreciable change in 57 years.


| 6861 J 416 |  |  | $-3^{\circ} 2380$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $8^{\mathrm{h}} 29^{\text {m. }} 3$ | $-3^{\circ} 37^{\prime}$ |  | 9.0-10.1 | F8 |
| 62.940 | 190:1 | 0'.46 |  |  |
| 62.953 | 191.7 | 0.43 |  |  |


| 64.905 | 189.9 | 0.43 |  |
| :---: | :---: | :---: | :---: |
| $\begin{gathered} 63.60 \\ \text { Doubtful change. } \end{gathered}$ | 190.6 | 0.44 | $3 n$ |
| Couteau 10 |  |  | $+89^{\circ} 13^{\prime}$ |
| $8{ }^{\text {h }} 43 .{ }^{\text {m. }} 7$ | $+88^{\circ} 46^{\prime}$ |  | 7.1-10.4 A0 |
| 61.256 | 63.1 | 1"90 |  |
| 61.259 | 60.1 | 2.13 |  |
| 61.265 | 62.6 | $\underline{2.07}$ |  |
| 61.26 | 61.9 | 2.03 | 3 n |

The position is corrected from the Greenwich astrographic catalogue.

| 7037 J 74 |  | $+1^{\circ} 2181$ |
| :---: | :---: | :---: |
| $8{ }^{\text {h }} 47.7$ | $+1^{\circ} 36^{\prime}$ | 11.3-11.4 |


| 63.340 | 120.7 | 6.69 |
| :--- | :--- | :--- |
| 63.351 | 119.2 | 6.86 |
| $\underline{64.905}$ | $\underline{120.0}$ | $\underline{6.90}$ |
| 63.87 | 120.0 | 6.82 |

The relative change leads to an improbably large $\mathrm{dp}=0 .!12$. Probably optical.

$$
\begin{aligned}
& 7044 \text { Van de Kamp } 3 \quad+8^{\circ} 2131 \\
& 8^{\mathrm{h}} \mathrm{~h}^{\mathrm{B}^{\mathrm{m}} .0} 0+8^{\circ} 3^{\prime} \quad 10.3-10.4
\end{aligned}
$$

The increase in distance seems to have slowed down to a maximum. The relative motion corresponds to $\mathrm{dp}=0$ '!060.

| 7054 A 1584 |  |  | $+55^{\circ} 1297$ |
| :---: | :---: | :---: | :---: |
| $8{ }^{\text {h }} 49^{\text {m }}$. 4 | $+55^{\circ} 8^{\prime}$ |  | 8.2-8.2 G0 |
| 61.201 | $112: 5$ | 0'70 |  |
| 61.228 | 109.9 | 0.79 |  |
| 61.237 | 111.9 | 0.66 |  |
| 61.240 | 108.3 | 0.76 |  |
| 61.250 | 111.7 | 0.72 |  |
| 63.337 | 115.2 | 0.71 |  |
| 63.348 | 113.5 | 0.71 |  |
| 65.037 | 120.0 | 0.62 |  |
| 65.040 | 115.4 | 0.58 |  |
| 65.133 | 117.5 | 0.59 |  |
| 65.140 | $\underline{120.5}$ | 0.62 |  |
| 61.23 | 110.9 | 0.73 | 5n |
| 63.34 | 114.3 | 0.71 | 2 n |
| 65.09 | 118.4 | 0.60 | 4n |
| Dommanget's orbit (1962) gives the residuals: |  |  |  |
| 61.23 63.34 | ++ <br> + <br> +2.1 | +0.08 |  |
| 63.34 65.09 | + 2.3 | +0.07 |  |
| $7067 \quad \Sigma 1280$ |  |  | $+71^{\circ} 482$ |
| $8{ }^{h_{51}{ }^{\text {m }} 0}$ | $+71^{\circ} 0^{\prime}$ |  | 9.3-9.4 K5 |
| 65.037 | 78.1 | 2'24 |  |
| 65.040 | 78.6 | $\underline{2.22}$ |  |
| ```65.04 Comparison with residuals: orbit l orbit 2``` | 78.4 | 2.23 | (1956) gives the |
|  | Rabe's | orbits |  |
|  | - $4: 0$ | +0.40 |  |
|  | +9.3 | -0.20 |  |
| 7102 A 2968 |  |  | $+11^{\circ} 194$ |
| $8{ }^{\text {h }} 54.4$ | $+10^{\circ} 57^{\prime}$ |  | 9.2-9.2 F5 |
| 61.259 | $137 \% 1$ | 1"19 |  |
| 61.265 | $\underline{138.3}$ | 1.32 |  |
| 61.26 | 137.7 | 1.26 | 2n |
| No change. |  |  |  |



| 63.337 | 215.3 | 0.90 |  |
| :--- | :--- | :--- | :--- |
| $\underline{64.947}$ | $\underline{216.3}$ | $\underline{0.82}$ |  |
| 63.74 | 216.4 | 0.87 | 3 n |

Slow change.

| 7341 A 2477 |  |  | $+18^{\circ} 2182$ |
| :---: | :---: | :---: | :---: |
| $9^{\text {h }} 20.7$ | $18^{\circ} 21{ }^{\prime}$ |  | 7.3-8.8 G0 |
| 63.348 | 304:5 | 0'37 |  |
| 65.136 | 308.6 | 0.31 |  |
| 65.140 | 306.8 | 0.34 |  |
| 65.209 | 311.5 | 0.35 |  |
| $\underline{65.215}$ | $\underline{310.6}$ | 0.33 |  |
| 64.81 | 308.4 | 0.34 | 5n |
| The angle has increased by 70 degrees in 52 years. The corresponding dp is 0 ':0023. |  |  |  |
| 59 Hu 869 + $+15^{\circ} 2043$ |  |  |  |
| $9^{\text {h }} 22^{\text {m/4 }} 4$ | $+15^{\circ} 2^{\prime}$ |  | 10.1-10.3 K0 |
| 60.198 | 281.5 | 0.40 |  |
| 61.265 | 280.6 | 0.49 |  |
| 62.228 | $\underline{281.3}$ | 0.47 |  |
| 61.23 | 281.1 | 0.45 | 3n |

Slow orbital motion corresponding to $d p=$ 0!.0043.


To the 58 degrees decrease in angle since 1910
corresponds an improbably large $d p=0!066$. Probably optical.


7541 Ho 369
$+37^{\circ} 2023$

| $9^{\mathrm{h}} 48^{\text {m }}$. ${ }^{\text {2 }}$ | $+36^{\circ} 43^{\prime}$ |  | 8.5-8.6 |
| :---: | :---: | :---: | :---: |
| 60.186 | 93:0 | 0.134 |  |
| 61.228 | 96.4 | 0.32 |  |
| 61.237 | 94.8 | 0.29 |  |
| 61.250 | 92.6 | 0.32 |  |
| 61.256 | 93.0 | 0.33 |  |
| 61.265 | 93.2 | 0.33 |  |
| 61.07 | 93.8 | 0.32 | 6n |

Since the quadrant is undeterminate the true nature of the orbit is still unknown.



Little change in angle but the distance has been halved since 1910. The corresponding dp is 0',0057.

| M1b 127 |  | $+57^{\circ} 1264$ |  |
| :---: | :---: | :---: | :---: |
| $10^{\mathrm{h}} 17^{\mathrm{m}} 9$ | $+56^{\circ} 40^{\prime}$ |  | 11.1-11.2 |
| 60.186 | $232: 3$ | 2".42 |  |
| 60.201 61.160 | 233.8 | 2.54 |  |
| 61.160 | $\underline{232.5}$ | $\underline{2.31}$ |  |
| 60.52 | 232.9 | 2.42 | $3 n$ |


| $7758 \times 1429$ |  |  | +25 ${ }^{\circ} 2247$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $10^{\mathrm{h}} 22^{\mathrm{m}} 3$ | +24* $53{ }^{\prime}$ |  | 7.0-9.0 | G5 |
| 61.097 | $202: 7$ | 0'.59 |  |  |
| 61.160 | 201.3 | 0.65 |  |  |
| $\underline{61.179}$ | $\underline{204.8}$ | 0.61 |  |  |
| 61.15 | 202.9 | 0.62 | 3 n |  |

The longer arc makes $d p=0.6011$.


| $10^{\mathrm{h}} 25^{\text {m. }} 0$ | $+36^{\circ} 58^{\prime}$ |  | 4.8-7.0 |
| :---: | :---: | :---: | :---: |
| 64.905 | 209:8 | 0'28 |  |
| 65.037 | 211.2 | 0.20 |  |
| 65.040 | $\underline{210.2}$ | 0.25 |  |
| 64.99 | 210.4 | 0.24 | 3 n |

Comparison with Baize's orbit (1950) gives the differences $+5: 8$ and -0'.02.

| 7831 A 2054 |  |  | $+46^{\circ} 1639$ |
| :---: | :---: | :---: | :---: |
| $10^{\mathrm{h}} 31 . \mathrm{m}$ | +46 ${ }^{\circ} 28^{\prime}$ |  | 9.5-9.5 |
| 62.285 | 204.7 | 0.19 |  |
| 62.301 | 204.0 | 0.20 |  |
| $\underline{62.313}$ | $\underline{205.8}$ | 0.22 |  |
| 62.30 | 204.8 | 0.20 | 3 n |

Motion questionable.

| 7844 A 2055 |  |  | $+45^{\circ} 1844$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $10^{\text {h }} 33.6$ | +44 ${ }^{\circ} 46^{\prime}$ |  | 8.9-8.9 | F8 |
| 62.285 | $153: 8$ | 0!36 |  |  |
| 62.292 | 152.0 | 0.34 |  |  |
| 62.301 | 155.7 | 0.34 |  |  |
| 62.940 | 157.5 | 0.36 |  |  |
| 63.334 | 158.9 | 0.35 |  |  |
| 65.136 | 158.3 | 0.36 |  |  |
| 65.140 | 158.8 | 0.38 |  |  |
| 65.212 | 157.2 | 0.39 |  |  |
| 63.58 | 156.5 | 0.36 | 8 n |  |

To the slow change in both coordinates corresponds $\mathrm{dp}=0.0042$.

## 7855

O $\sum 222$
$+60^{\circ} 1274$
$10^{\mathrm{h}_{3} 5^{\mathrm{m}} \mathrm{m}_{1} \quad+60^{\circ} 24^{\prime} \quad 7.0-11.0 \quad \mathrm{~F} 8}$

| 61.220 | 339.1 | 4.22 |  |
| :--- | :--- | :--- | :--- |
| 61.223 | 342.0 | 4.05 |  |
| 61.237 | 339.5 | 4.24 |  |
| 61.240 | 339.5 | 4.37 |  |
| 61.250 | 339.9 | 4.25 |  |
| 61.23 | 340.0 | $\frac{4.23}{}$ | $5 n$ |

After 114 years the motion is still questionable. The cormon proper motion of $0!21$ establishes the physical connection.


| 7871 OS 224 |  |  | $+9^{\circ} 2382$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $10^{\mathrm{h}} 37^{\mathrm{m}_{1}}$ | $+9^{\circ} 6^{\prime}$ |  | 8.1-10.1 | F5 |
| 61.265 | 205:9 | 0.43 |  |  |
| 61.330 | 199.4 | 0.45 |  |  |
| 61.335 | 202.3 | 0.45 |  |  |
| 62.228 | 202.9 | 0.46 |  |  |
| $\underline{62.285}$ | 199.2 | 0.50 |  |  |
| 61.69 | 201.9 | 0.46 | 5n |  |
| Baize's orbit and + 0 !. 01 . | (1958) ma | es the | residual | s +2:7 |



| 7881 A 1351 |  | $-1^{\circ} 2422$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $10^{h_{38}}{ }^{\text {m }}$, | - $1^{\circ} 56{ }^{\prime}$ |  | 10.0-10.2 | G5 |
| 61.234 61.335 |  | 0.34 0.34 |  |  |
| 61.335 <br> 62.313 | $\begin{array}{r} 356.4 \\ 356.7 \\ \hline \end{array}$ | 0.34 <br> 0.39 <br> 0.36 |  |  |
| 61.63 | 357.7 | 0.36 | $3 n$ |  |


| 7982 |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\beta 1076$ |  | $+1^{\circ} 2501$ |  |  |  |
| $10^{\mathrm{h}_{53} 3^{\mathrm{m}_{1}}}$ | $+1^{\circ} 0^{\prime}$ | $6.0-10.5$ |  |  | F 2 |
| 63.334 | 80.5 | 0.92 |  |  |  |
| 65.040 | $\frac{80.3}{80.4}$ | $\frac{1.02}{0.97}$ | 2 n |  |  |

Baize's orbit (1957) requires correction since the residuals are $+10 \% 9$ and $+0!19$.

| 8011 A 2376 |  | $+20^{\circ} 2541$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $10^{h_{57}}{ }^{\text {m }} 6$ | $+19^{\circ} 42^{\prime}$ |  | 10.4-10.9 | G5 |
| 62.285 | $9: 8$ | 0'.23 | 1 n |  |
| Unchanged in 51 years. |  |  |  |  |
| 8032 A 1590 |  |  | +55 ${ }^{\circ} 1439$ |  |
| $11^{\mathrm{h}} 0.6$ | $+54^{\circ} 49^{\prime}$ |  | 9.2-9.7 | G0 |
| 61.160 | 356:2 | 1'.16 |  |  |
| 61.201 | 355.0 | 1.10 |  |  |
| 61.220 | 355.4 | 1.24 |  |  |
| 61.224 | 356.7 | 1.16 |  |  |
| 61.237 | 355.0 | 1.23 |  |  |
| 61.21 | 355.7 | 1.18 | 5n |  |

Heintz' orbit (1963) gives the residuals $+1^{\circ} 1$ and +0.06

| 8039 J 1262 |  |  |  |
| :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} \mathrm{m}_{\mathrm{m}} .0$ | $+19^{\circ} 7^{\prime}$ |  | $9.6-9.6$ |
| 61.168 | 178.0 | 2.25 |  |
| 61.220 | 177.0 | 2.36 |  |
| 61.224 | 180.2 | 2.34 |  |
| 61.237 | 177.8 | 2.50 |  |
| $\frac{61.256}{61.22}$ | $\frac{178.0}{178.2}$ | $\underline{2.33}$ | 2.36 |
| $5 n$ |  |  |  |

Probably unchanged.

| Kui 47 |  |  | $\chi$ Leo |  |
| :---: | :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} \mathrm{m}_{4}$ | $+7^{\circ} 36^{\prime}$ |  | 4.7-11 | F0 |
| 59.151 | $264: 3$ | 3! 48 |  |  |
| 60.198 | 266.8 | 3.56 |  |  |
| 62.940 | 265.2 | 3.41 |  |  |
| 63.334 | 268.2 | 3.43 |  |  |
| 63.348 | 265.2 | 3.42 |  |  |
| 64.905 | $\underline{264.8}$ | 3.58 |  |  |
| 62.31 | 265.8 | 3.48 | 6 n |  |

The increase in both coordinates corresponds to a $\mathrm{dp}=0$ ". 052 .

| 8060 ~ 599 |  |  | $+2^{\circ} 2387$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $11 \mathrm{~h}_{4} \mathrm{~m}_{4}$ | $+2^{\circ} 14^{\prime}$ |  | 5.7-11.6 | A0 |
| 61.179 | 101.4 | 3!.00 |  |  |
| 62.220 | 104.5 | 3.05 |  |  |
| 62.223 | 100.1 | 2.88 |  |  |
| 62.234 | 101.2 | $\underline{2.93}$ |  |  |
| 61.96 | 101.8 | 2.97 | $4 n$ |  |

Long period binary. In the Lick Index Catalogue the magnitude is given erroneously as 8.1.

| 8085 A 2156 |  |  | $+35^{\circ} 2219$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $11{ }^{\mathrm{h}} 9 \mathrm{~m}_{3}$ | $+35^{\circ} 17^{\prime}$ |  | 8.3-9.1 | A2 |
| 61.237 | 236:3 | 0.34 |  |  |
| 61.330 | 234.6 | 0.36 |  |  |
| 62.285 | 234.1 | 0.38 |  |  |
| 62.305 | $\underline{232.3}$ | 0.37 |  |  |
| 61.79 | 234.3 | 0.36 | 4n |  |

Slow decrease in angle with little change in distance.

| $8094 \times 1517$ |  |  | $+20^{\circ} 2572$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 11^{\mathrm{m}} 1$ | +20 ${ }^{\circ} 25^{\prime}$ |  | 7.7-7.7 | G0 |
| 65.037 | 170:0 | 0.32 |  |  |
| 65.040 | 169.4 | 0.33 |  |  |
| 65.130 | 170.2 | 0.32 |  |  |
| 65.136 | 171.0 | 0.28 |  |  |
| 65.140 | 168.1 | 0.32 |  |  |
| 65.10 | 169.7 | 0.31 | 5 n |  |

It is now evident that the total angular motion is $120^{\circ}$ and that after the minimum distance around 1950 the pair is now opening up in the second quadrant. Long period binary.

| $8148 \times 1536$ |  |  | ¢ Le |
| :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 21 \mathrm{~m}_{3}$ | $+10^{\circ} 48^{\prime}$ |  | 4.1-7.3 |
| 60.198 | 213.1 | 0.'97 |  |
| 61.220 | 213.8 | 1.16 |  |
| 61.223 | 209.5 | 0.98 |  |
| 63.334 | 202.7 | 1.11 |  |
| 63.337 | 199.5 | 1.11 |  |
| 63.348 | 201.8 | 1.00 |  |
| 63.368 | 198.3 | 1.06 |  |
| 65.037 | 188.9 | 1.22 |  |
| 65.040 | 193.5 | 1.09 |  |
| 65.140 | 191.3 | 1.06 |  |
| 60.88 | 212.1 | 1.04 | $3 n$ |



Comparison with Couteau's orbit (1962) gives the residuals $+0: 8$ and $+0!.03$.

| 8182 A 7 |  |  | $-5^{\circ} 3300$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 27.2$ | - $6^{\circ} 2^{\prime}$ |  | 9.7-9.7 | F8 |
| 61.335 | $244 \% 5$ | 0.36 |  |  |
| 62.228 | 242.8 | 0.32 |  |  |
| 61.78 | 243.6 | 0.34 | 2n |  |
| The longer arc | kes dp | 0.'004 |  |  |


| 8189 O§ 234 |  |  | $+42^{\circ} 2214$ |
| :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 28{ }^{\mathrm{m}} 1$ | +41 $34^{\prime}$ |  | 7.6-8.0 F5 |
| 65.037 | 343.7 | 0!28 |  |
| 65.040 | 350.4 | 0.28 |  |
| 65.140 | 347.1 | 0.25 |  |
| 65.212 | 345.9 | 0.24 |  |
| 65.215 | 344.0 | 0.27 |  |
| 65.13 | 346.2 | 0.26 | 5 n |
| Muller's orbit | (1955) giv | s the | residuals |


| 8231 | $\Sigma 1555$ |  |  | $+28^{\circ} 2022$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | $11^{\mathrm{h}_{33} \mathrm{~m}_{7}}$ | $+28^{\circ} 3^{\prime}$ |  | 6.4-6.8 | A3 |
|  | 60.186 | $139: 0$ | 0:36 |  |  |
|  | 60.400 | 141.1 | 0.36 |  |  |
|  | 65.130 | 135.3 | 0.42 |  |  |
|  | 65.136 | 138.9 | 0.40 |  |  |
|  | 65.140 | 138.0 | 0.38 |  |  |
|  | 65.215 | 135.4 | 0.41 |  |  |
|  | 63.53 | 138.0 | 0.39 | $6 n$ |  |
| Long p | period or | al mot |  |  |  |

$8242 \mathrm{Kii} 39 \quad+48^{\circ} 1958$

| $11^{\mathrm{h}_{3}} 34^{\mathrm{m}} .6$ | $+47^{\circ} 45^{\prime}$ | $11.1-11.4$ | MO |
| :---: | :---: | :---: | :---: |
|  |  |  |  |
| 60.198 | 58.5 | $1^{\prime!} .94$ |  |
| $\frac{60.500}{60.35}$ | $\frac{60.8}{59.6}$ | $\frac{1.92}{1.93}$ | $2 n$ |

To the change in both coordinates corresponds $\mathrm{dp}=0!037$.

| 8302 B 602 |  |  | $+15^{\circ} 2378$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 44^{\mathrm{m}} 3$ | $+15^{\circ} 17^{\prime}$ |  | 8.0-10.5 | A5 |
| 62.228 | 106:0 | 0'.48 |  |  |
| 62.285 | 104.0 | 0.54 |  |  |
| 63.334 | 108.2 | 0.51 |  |  |
| 62.62 | 106.1 | 0.51 | 3 n |  |

The longer arc reduces dp to $0^{\prime \prime} .003$.

| 8311 ~ 603 |  |  | $+15^{\circ} 2381$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 46.1$ | $+14^{\circ} 34^{\prime}$ |  | 5.9-10.1 | A5 |
| 60.198 | $13: 0$ | 0.'62 |  |  |
| 61.335 | 8.4 | 0.64 |  |  |
| 62.285 | 8.7 | 0.66 |  |  |
| 65.130 | 7.0 | 0.72 |  |  |
| 65.140 | 3.8 | 0.70 |  |  |
| 61.27 | 10.3 | 0.64 | $3 n$ |  |
| 65.14 | 5.4 | 0.71 | 2n |  |

Comparison with Heintz' orbit (1963) gives the residuals: $\begin{array}{lll}61.27 & -0.5 & -0.09 \\ 65.14 & +0.4 & -0.09\end{array}$

8312 Kü 40
$+34^{\circ} 2259$
$11 \mathrm{~h}_{46} \mathrm{~m}_{3}$
$+33^{\circ} 59^{\prime}$
10.2-10.8

| 60.186 | 185.6 | $2!.88$ |  |
| :--- | :--- | :--- | :--- |
| 61.220 | 187.1 | 2.91 |  |
| 61.224 | 186.3 | $\underline{2.94}$ |  |
|  | ling.3 | 2.91 | $3 n$ |
| nged in 59 | years. |  |  |

Unchanged in 59 years.
8322 Es 1644

$$
11^{\mathrm{h}_{4} 8^{\mathrm{m}} 6} \quad+38^{\circ} 46^{\prime} \quad 10.0-10.1
$$

This star was not found $(60.186)$ in the catalogue position which is blank.

| 8323 Hu 730 |  |  | $+51^{\circ} 1$ |
| :---: | :---: | :---: | :---: |
| $11^{\mathrm{h}} 48 \mathrm{~m}_{8}$ | +50 ${ }^{\circ}{ }^{\prime \prime}$ |  | 10.5-10.5 |
| 62.203 | 178:0 | 0'.27 |  |
| 62.285 | 177.6 | 0.29 |  |
| 65.140 | 174.2 | 0.27 |  |
| 65.209 | 175.8 | 0.27 |  |
| 63.71 | 176.4 | 0.28 | 4 n |

The longer arc reduces dp to $0!006$.

| 8325 Hu 731 |  |  | $+48^{\circ} 1978$ |
| :---: | :--- | :--- | :---: |
| $11^{\mathrm{h}} \mathrm{H9}^{\mathrm{m} .4}$ | $+48^{\circ} 22^{\prime}$ |  | $9.6-9.8$ |
| 61.237 | 337.2 | 0.33 |  |
| 61.256 | 335.2 | 0.35 |  |
| 61.335 | 334.6 | 0.33 |  |
| 62.203 | 337.3 | 0.34 |  |
| 62.285 | 333.4 | 0.37 |  |
| 62.305 | $\underline{331.3}$ | $\underline{0.33}$ |  |
| 61.77 | 334.8 | 0.34 | $6 n$ |

The angular change corresponds to a period of about 150 years but the orbit is still undetermined.
 $+1: 2$ and $0!100$.

| 8345 Es 724 |  |  | $+51^{\circ} 1$ |
| :---: | :---: | :---: | :---: |
| $11^{h_{51}}{ }^{\text {m }} 7$ | $+50^{\circ} 50^{\prime}$ |  | 9.9-12.2 |
| 61.237 | $229: 7$ | 2!.64 |  |
| 61.256 | 226.6 | 2.80 |  |
| 61.259 | 226.1 | 2.63 |  |
| 61.322 | 224.8 | 2.77 |  |
| 61.330 | 228.4 | 2.60 |  |
| 61.335 | 228.0 | 2.59 |  |
| 61.29 | 227.3 | 2.67 | $6 n$ |

Unchanged.
Wor $20 \quad+16^{\circ} 2316$

| $11^{\mathrm{h}_{51} \mathrm{~m}_{9}}$ | $+15^{\circ} 32^{\prime}$ | $11.0-11.2 \quad$ M0 |  |
| :---: | :--- | :--- | :--- |
| 62.285 | $316^{\circ} .1$ | 1.45 |  |
| $\frac{62.301}{62.29}$ | $\frac{316.8}{316.4}$ | $\frac{1.46}{1.46}$ | 2 n |

Wor 21
$11^{\mathrm{h}} 56^{\mathrm{m}} .3 \quad+59^{\circ} 50^{\prime} \quad 11.8-13.4 \mathrm{M} 0$

| 62.203 | 279.3 | 0.199 |  |
| :--- | :--- | :--- | :--- |
| 62.285 | $\frac{277.5}{278.4}$ | $\frac{1.19}{1.09}$ | $2 n$ |

It is $611^{\circ} 22221$ and $60^{\circ} 22753$ in the Vatican Astrographic Catalogue.

| 8415 A 1358 |  | $+57^{\circ} 1354$ |  |
| :---: | :---: | :---: | :---: |
| $12^{\mathrm{h}} 2^{\mathrm{m} .8}$ | $+57^{\circ} 3^{\prime}$ |  | $9.8-10.0 \quad \mathrm{G5}$ |
| 65.136 | $235^{\circ} .4$ | $0^{\prime \prime} .70$ |  |
| 65.140 | 231.1 | 0.79 |  |
| $\frac{65.212}{65.16}$ | $\frac{233.1}{233.2}$ | $\frac{0.69}{0.73} 3 n$ |  |

Hardly changed in 59 years. The position is corrected from Aitken's General Catalogue.

| 8419 之 3123 |  |  | 44 |
| :---: | :---: | :---: | :---: |
| $12^{\text {h }} 3^{\text {m. }} 5$ | +68 ${ }^{\circ} 59^{\prime}$ |  | 7.9-7.9 |
| 61.256 | 99:3 | 0.129 |  |
| 61.322 | 96.5 | 0.32 |  |
| 61.330 | 98.9 | 0.31 |  |
| 61.30 | 98.2 | 0.31 | 3 n |

Baize's orbit (1944) gives the residuals $+5: 1$ and +0.03 .

| $\underline{8424}$ A 76 |  | +71 ${ }^{\circ} 603$ |  |
| :---: | :---: | :---: | :---: |
| $12^{\text {h }} 3^{\text {m }} 7$ | +70 ${ }^{\circ} 39^{\prime}$ |  | 10.0-11.1-12.1 KO |
| A | BC |  |  |
| 61.223 | $42: 9$ | 22'.29 |  |
| 61.330 | 43.0 | 22.38 |  |
| 62.203 | 42.5 | 22.21 |  |
| $\underline{62.285}$ | 42.7 | 22.13 |  |
| 61.76 | 42.8 | 22.25 | $4 n$ |
|  | BC |  |  |
| 61.223 | $321: 8$ | 1'.43 |  |
| 61.330 | 318.7 | 1.32 |  |
| 62.203 | 322.8 | 1.42 |  |
| 62.285 | 321.1 | 1.33 |  |
| 61.76 | 321.1 | 1.38 | 4n |

No change in the wide pair since 1900 but the angle of $B C$ has decreased making $d p=0!0012$.

| 8433 A 1998 |  |  | $+43^{\circ}$ |
| :---: | :---: | :---: | :---: |
| $12^{h_{5} \mathrm{~m}_{3}}$ | $+42^{\circ} 59^{\prime}$ |  | 9.6-9.6 |
| 61.179 | 15:2 | 0.136 |  |
| 61.223 | 13.0 | 0.33 |  |
| 62.203 | 19.5 | 0.29 |  |
| 62.285 | 18.2 | 0.33 |  |
| 61.72 | 16.5 | 0.33 | $4 n$ |




| 60.503 | 323.0 | 0.99 |  |
| :--- | ---: | ---: | ---: |
| 61.223 | 325.3 | 0.96 |  |
| 61.330 | 324.0 | 1.08 |  |
| $\underline{62.373}$ | $\underline{323.6}$ | $\underline{0.96}$ |  |
| 61.36 | 324.0 | 1.00 | 4 n |

No change so far.

| 8790 Mlb 170 |  |  |
| :---: | :---: | :---: |
| $13^{\mathrm{h}} 5 \mathrm{~m}_{6} 6$ | $+54^{\circ} 56{ }^{\prime}$ | 11.0-11.4 |
| 61.250 | 272.1 | 4:00 |
| 61.327 | 272.2 | 3.81 |
| $\underline{61.396}$ | $\underline{273.0}$ | 3.91 |
| Change questionable. This pair is Vab. ph $56^{\circ} 45825-6$ which indicates <br> $18.18 \quad 270: 0 \quad 3!30$ |  |  |
|  |  |  |
|  |  |  |
| $8804 \sum 1728$ ( 42 Comae |  |  |
| $13^{\text {h }} 7.6$ | $+17^{\circ} 47^{\prime}$ | 5.2-5.2 F5 |
| 64.404 | 11.3 | 0 O .43 |
| 64.410 | 12.6 | 0.46 |
| 65.037 | 12.6 | 0.42 |
| 64.62 | 12.2 | 0.44 3n |

Pavel's orbit (1944) gives the residuals $+0: 3$ and +0.11 .

| 8805 ß 608 |  |  | $+39^{\circ} 2614$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $13^{\mathrm{h}} 7 \mathrm{~m} .8$ | $+38^{\circ} 46^{\prime}$ |  | 6.2-11.2 | B9 |
| 61.220 | 273.5 | 1".29 |  |  |
| 61.223 | 274.8 | 1.29 |  |  |
| 61.237 | 273.4 | 1.26 |  |  |
| 61.250 | 272.2 | 1.48 |  |  |
| 61.256 | 269.8 | 1.29 |  |  |
| 61.330 | $\underline{270.1}$ | 1.39 |  |  |
| 61.25 | 272.3 | 1.33 | 6n |  |

A slow decrease in angle is now evident. To this change corresponds $\mathrm{dp}=0$ ".007.
Rst 3829
$-10^{\circ} 3635$

| $13^{\mathrm{h}} 12 \mathrm{~m}_{3}$ | $-11^{\circ} 6^{\prime}$ |  | $7.1-9.0$ |
| ---: | :--- | :--- | :--- |
|  | G0 |  |  |
| 61.250 | 290.7 | 0.83 |  |
| 61.423 | 287.8 | 0.83 |  |
| $\frac{61.426}{61.37}$ | $\frac{290.9}{289.8}$ | $\underline{0.90}$ |  |

Increase in both coordinates indicating a $d p=$ 0':042.

| 8862 Hu 644 |  |  | $+48^{\circ} 2108$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $13^{\mathrm{h}} 17{ }^{\text {m }} 6$ | $+48^{\circ} 2^{\prime}$ |  | 9.0-9.8 | K0 |
| 65.037 | 83.7 | 0.'96 |  |  |
| 65.130 | 84.7 | 1.11 |  |  |
| 65.136 | 83.7 | 1.09 |  |  |
| 65.212 | 83.5 | 0.93 |  |  |
| 65.13 | 83.9 | 1.02 | $4 n$ |  |

Heintz' orbit (1963) leaves the residuals $-5: 0$ and $+0: 04$.

| 8864 \ 1734 |  |  | $+3^{\circ} 2758$ |
| :---: | :---: | :---: | :---: |
| $13^{\mathrm{h}} 18.2$ | $+3^{\circ} 12^{\prime}$ |  | 6.7-7.4 |
| 63.334 | $180: 9$ | 1!38 |  |
| 63.348 | 179.9 | 1.38 |  |
| 63.359 | 181.4 | 1.42 |  |
| 64.404 | 181.4 | 1.35 |  |
| 65.130 | 183.0 | 1.52 |  |
| 65.136 | 181.4 | 1.43 |  |
| 65.209 | 181.1 | 1.46 |  |
| 64.27 | 181.3 | 1.42 | $7 n$ |

Orbital motion mostly in distance. The
longer arc makes dp 00.0076 .
-

8875



To the siow angular increase corresponds $\mathrm{dp}=$ $0!004$.

| $\underline{9220}$ A 1102 |  |  | $+69^{\circ} 743$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 17.0$ | $+69^{\circ} 28^{\prime}$ |  | 8.9-9.1 |  |
| 60.529 | 137:7 | $0!.22$ |  |  |
| 60.588 | 137.3 | 0.24 |  |  |
| 60.595 | 133.8 | 0.25 |  |  |
| 61.393 | 136.1 | 0.23 |  |  |
| 61.567 | 135.0 | 0.24 |  |  |
| 61.574 | 132.5 | 0.24 |  |  |
| 61.583 | 134.9 | 0.26 |  |  |
| 60.57 | 136.3 | 0.24 | 3 n |  |
| 61.53 | 134.6 | 0.24 | $4 n$ |  |
| Couteau's orbit | (1960) | gives | the resi | uals: |
| 60.57 | +3:8 | +0.02 |  |  |
| 61.53 | $+5.1$ | +0.02 |  |  |
| $\underline{9229} \boldsymbol{2} 1834$ |  |  | $+49^{\circ}$ | 294 |
| $14^{\mathrm{h}} 18{ }^{\mathrm{m}} 5$ | +48 ${ }^{\circ} 4^{\prime}$ |  | 7.9-8.0 | F8 |
| 63.334 | $101: 0$ | 1!.03 |  |  |
| 63.337 | 102.6 | 1.06 |  |  |
| 63.348 | 100.4 | 1.10 |  |  |
| 63.34 | 101.3 | 1.06 | $3 n$ |  |

Van den Bos' orbit (1939) still holds well The residuals are $-1: 3$ and $0!00$.
 $\beta 1111$
$+9^{\circ} 2882$
$14^{\mathrm{h}} 20^{\mathrm{m}} 9$
$+8^{\circ} 40^{\prime}$
7.4-7.7 A0

| 60.186 | 218.5 | 0.16 |  |
| :--- | :--- | :--- | :--- |
| 60.189 | 220.3 | 0.16 |  |
| 60.503 | 216.2 | 0.18 |  |
| 60.526 | 220.6 | 0.17 |  |
| 64.418 | 266.1 | 0.16 |  |
| 65.040 | $\underline{265.4}$ | $\underline{0.16}$ |  |
| 60.35 | 218.9 | 0.17 | $4 n$ |
| 64.73 | 265.7 | 0.16 | $2 n$ |

Van den Bos' earlier orbit (1938) gives smaller residuals than Couteau's more recent one (1958):

|  | Van den Bos |  | Couteau |  |  |
| :--- | :--- | :--- | :--- | :--- | :---: |
| 60.35 | -1.7 | -0.101 | -5.7 | -0.101 |  |
| 64.70 | -1.1 | -0.03 | -6.6 | -0.02 |  |



To the angular decrease corresponds $d p=0!.007$.

| 9285 Hu 1268 |  |  | $+36{ }^{\circ} 2496$ |
| :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 27^{\mathrm{m}_{5}}$ | $+36^{\circ} 26^{\prime}$ |  | 9.5-10.0 |
| 60.526 | 227:2 | 0'.12 |  |
| 60.529 | 234.6 | 0.14 |  |
| 60.606 | 230.5 | 0.12 |  |
| 60.609 | 235.8 | 0.14 |  |
| 60.57 | 232.0 | 0.13 | 4n |

To the large change in both coordinates corresponds a dp $=00^{\prime \prime}, 005$. This will be a very difficult pair for some years.

9318 - 941
$+0^{\circ} 3206$
$14^{\mathrm{h}_{33} \mathrm{~m}_{2}}$
$+0^{\circ} 28^{\prime}$
9.6-9.6 F8
$62.450 \quad 186.5 \quad 0.47$

| 62.466 | 182.7 | 0.43 |  |
| :--- | :--- | :--- | :--- |
| $\frac{64.418}{63.11}$ | $\frac{186.9}{185.4}$ | $\underline{0.53}$ |  |
| 0.48 | $3 n$ |  |  |

The longer arc reduces dp to $0!30030$.

| 9324 A 347 |  |  | $+48^{\circ} 2222$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 35^{\text {m }} 2$ | $+48^{\circ} 26^{\prime}$ |  | 8.5-8.7 | F2 |
| 64.418 | 292:7 | 0'.50 |  |  |
| 65.037 | 295.2 | 0.49 |  |  |
| 64.73 | 294.0 | 0.50 | 2n |  |

Comparison with the orbit of Guntzel-Lingner (1955) gives the small residuals $+0: 7$ and -0! 02 .

| 9378 OS 285 |  |  | $+42^{\circ} 2531$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 43^{\mathrm{m}} 6$ | +42 ${ }^{\circ} 35^{\prime}$ |  | 7.7-8.2 | F5 |
| 61.393 | 2:8 | $0!.24$ |  |  |
| 61.418 | 0.3 | 0.23 |  |  |
| 61.423 | 2.4 | 0.27 |  |  |
| 61.426 | 4.0 | 0.27 |  |  |
| 61.428 | 0.1 | 0.24 |  |  |
| 65.037 | 337.7 | 0.18 |  |  |
| 65.040 | 340.4 | 0.21 |  |  |
| 61.42 | 1.9 | 0.25 | 5n |  |
| 65.04 | 339.0 | 0.20 | 2n |  |
| Simonov's orbit | (1937) | gives | the resi | uals: |
| 61.42 | $+1: 0$ | +0'.03 |  |  |
| 65.04 | +11.0 | +0.04 |  |  |


| 9392 \ 1883 |  |  | $+6^{\circ} 2946$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 46 . \mathrm{m}_{4}$ | $+6^{\circ} 10^{\prime}$ |  | 7.5-7.5 | F8 |
| 61.396 | 100.5 | 0.!38 |  |  |
| 61.418 | 98.6 | 0.34 |  |  |
| 61.429 | 102.6 | 0.35 |  |  |
| 61.426 | 100.8 | 0.36 |  |  |
| 61.428 | 97.8 | 0.38 |  |  |
| 61.42 | 100.1 | 0.36 | 5n |  |

Baize's orbit (1961) gives the residuals $+4: 9$ and -0.01 .

| 9397 A 2983 |  |  | $+10^{\circ} 2747$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 46^{\text {m }} 8$ | $+10^{\circ} 26^{\prime}$ |  | 9.3-9.3 | G5 |
| 60.526 | $243: 3$ | 0'.12 |  |  |
| 60.529 | 244.0 | 0.12 |  |  |
| 60.606 | 246.4 | 0.13 |  |  |
| 60.615 | 246.9 | 0.13 |  |  |
| 60.618 | $\underline{245.7}$ | 0.14 |  |  |
| 60.58 | 245.3 | 0.13 | 5n |  |
| Comparison wi | Van den | Bos' | rbit (1 | 54) |
| gives the res | uals -13 | 1 and | -0.01. |  |


| 9400 A 1110 |  |  | $+8^{\circ} 2925$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 47 \mathrm{~m}^{\mathrm{m}}$ | $+8^{\circ} 11^{\prime}$ |  | 7.5-7.9 | F5 |
| 62.450 | 255:9 | 0.\%53 |  |  |
| 62.455 | 253.7 | 0.49 |  |  |
| 62.466 | 254.2 | 0.48 |  |  |
| 62.46 | 254.6 | 0.50 | 3 n |  |

The slow change gives a $\mathrm{dp}=0!$.005

| 1 A 1627 |  |  | $+40^{\circ} 2829$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 53.8$ | $+39^{\circ} 51^{\prime}$ |  | 8.7-8.7 | FO |
| 60.526 | 16:2 | $0: 25$ |  |  |
| 60.598 | 17.1 | 0.24 |  |  |
| 60.606 | 12.5 | 0.22 |  |  |
| 60.609 | 16.2 | 0.21 |  |  |
| 60.58 | 15.5 | 0.23 | $4 n$ |  |

Baize's orbit (1959) gives the residuals $-4: 4$ and +0.08 .


Slow orbital motion indicating a dp $=0!0027$.


| 9482 A 1629 |  |  | $+42^{\circ} 2557$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $14^{\mathrm{h}} 59{ }^{\text {m }} 4$ | +41 ${ }^{\circ} 5^{\prime}$ |  | 9.0-13.0 | K |
| 60.526 | $280: 7$ | 2!03 |  |  |
| 60.588 | 279.9 | 2.07 |  |  |
| 60.595 | 279.7 | 1.91 |  |  |
| 60.57 | 280.1 | 2.00 | $3 n$ |  |
| Hardly changed in 53 years. |  |  |  |  |
| 9483 Hu 1155 |  |  | $+15^{\circ} 2806$ |  |
| $14^{h_{59}}{ }^{m_{5}}$ | $+15^{\circ} 18^{\prime}$ |  | 9.4-10.9 F8 |  |
| 59.627 | $17: 8$ | 4:03 |  |  |
| 59.640 | 16.4 | 3.87 |  |  |
| 59.660 | 15.1 | 4.01 |  |  |
| 59.64 | 16.4 | 3.97 | $3 n$ |  |
| No evidence of | a change | in 54 | years. |  |


| $9494 \times 190$ |  |  | 44 Boo |  |
| :---: | :---: | :---: | :---: | :---: |
| $15^{\mathrm{h}} \mathrm{m}_{2}$ | $+47^{\circ} 5$ |  | 5.3-6.2 | G0 |
| 59.151 | $269{ }^{\circ} \mathrm{2}$ | 0!.99 |  |  |
| 59.422 | 267.7 | 1.06 |  |  |
| 59.632 | 268.7 | 1.10 |  |  |
| 59.656 | 266.3 | 0.96 |  |  |
| 59.663 | 268.9 | 0.96 |  |  |
| 60.500 | 273.9 | 0.99 |  |  |
| 61.179 | 272.1 | 0.92 |  |  |
| 61.217 | 271.8 | 1.00 |  |  |
| 61.220 | 273.5 | 0.95 |  |  |
| 61.223 | 271.7 | 0.95 |  |  |
| 61.237 | 272.8 | 1.04 |  |  |
| 61.486 | 274.6 | 0.96 |  |  |
| 62.288 | 278.3 | 1.06 |  |  |
| 62.285 | 276.0 | 0.11 |  |  |
| 62.305 | 275.1 | 0.99 |  |  |
| 62.373 | 280.2 | 0.96 |  |  |
| 62.450 | 279.2 | 0.92 |  |  |
| 62.698 | 274.1 | 1.00 |  |  |
| 63.334 | 281.4 | 0.80 |  |  |
| 63.337 | 282.0 | 0.82 |  |  |
| 63.348 | 277.0 | 0.90 |  |  |
| 63.359 | 279.3 | 0.79 |  |  |
| 64.415 | 285.5 | 0.80 |  |  |


| 64.418 | $283: 6$ | 0.'74 |  |
| :---: | :---: | :---: | :---: |
| 65.037 | 287.9 | 0.70 |  |
| 65.040 | 286.7 | 0.64 |  |
| 65.130 | 285.4 | 0.71 |  |
| 65.136 | 291.0 | 0.73 |  |
| 65.140 | 286.0 | 0.64 |  |
| 65.209 | 287.5 | 0.72 |  |
| 65.212 | 286.8 | 0.59 |  |
| 65.215 | 287.1 | 0.58 |  |
| 59.50 | 268.2 | 1.01 | 5n |
| 61.15 | 272.9 | 0.97 | 7n |
| 62.40 | 277.1 | 0.84 | 6 n |
| 63.34 | 279.9 | 0.83 | 4n |
| 64.42 | 284.6 | 0.77 | 2n |
| 65.14 | 287.3 | 0.66 | 8 n |
| Heintz' orbit | (1963) | leaves th | residuals: |
| 59.50 | + 1\%3 | +0!01 |  |
| 61.17 | +1.9 | +0.07 |  |
| 62.42 | + 2.3 | +0.04 |  |
| 63.44 | +1.4 | +0.09 |  |
| 64.42 | +1.9 | +0.08 |  |
| 65.07 | +1.3 | +0.01 |  |


| 9504 A 689 | $-1^{\circ} 3021$ |
| :---: | :---: |
| $15^{\mathrm{h}_{4} \mathrm{~m}_{6}}$ | $-2^{\circ} 5^{\prime}$ |


| 60.523 | 328.2 | 0.31 |  |
| :--- | :--- | :--- | :--- |
| 60.529 | 327.9 | 0.31 |  |
| 60.601 | 330.6 | 0.33 |  |
| 62.446 | 321.3 | 0.32 |  |
| 64.418 | 322.5 | $\underline{0.32}$ |  |
| 60.55 | 328.9 | 0.32 | $3 n$ |
| 63.43 | 321.9 | 0.32 | $2 \pi$ |

May be the quadrant is to be reversed which would indicate a $175^{\circ}$ decrease in angle.

| 9511 Hu 143 |  |  | $55^{\circ} 1733$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $15^{\mathrm{h}} \mathrm{m}_{\text {m }} 1$ | $+55^{\circ} 27^{\prime}$ |  | 6.0-10.3 | G5 |
| 60.526 | 149:8 | 0'63 |  |  |
| 60.588 | 152.6 | 0.65 |  |  |
| 60.592 | 151.0 | 0.67 |  |  |
| 61.220 | 148.7 | 0.66 |  |  |
| 61.223 | 148.6 | 0.69 |  |  |
| 61.237 | 141.2 | 0.72 |  |  |
| 61.250 | 141.7 | 0.56 |  |  |
| 61.256 | 147.8 | 0.66 |  |  |
| 64.418 | 146.4 | 0.70 |  |  |
| 65.136 | 148.6 | 0.63 |  |  |
| 65.212 | 143.7 | 0.55 |  |  |
| 60.57 | 151.1 | 0.65 | $3 n$ |  |
| 61.24 | 145.6 | 0.66 | $5 \pi$ |  |
| 64.92 | 146.2 | 0.63 | 3 n |  |

Slow increase in angle corresponding to dp = 0':0033.

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Hardly changed in 130 years.

slow orbital motion. The longer arc makes dp $=0!010$.

| $15^{\text {h }} 13{ }^{\text {m }}$ S | $+36^{\circ} 9^{\prime}$ |  | 9.5-10.0 |
| :---: | :---: | :---: | :---: |
| 64.418 | 80.4 | 0'.33 |  |
| Unchanged in 60 years. |  |  |  |
| 9600 Hu 146 |  | +21 ${ }^{\circ} 2759$ |  |
| $15^{\mathrm{h}} 18{ }^{\mathrm{m}} \mathrm{m}_{8}$ | $+21^{\circ} 15^{\prime}$ |  | 9.3-9.6 G0 |
| $\begin{aligned} & 62.450 \\ & 62.466 \end{aligned}$ | $\begin{aligned} & 14109 \\ & 137.4 \\ & \hline \end{aligned}$ | $\begin{array}{r} 0.47 \\ 0.50 \\ \hline \end{array}$ |  |
| 62.46 | 139.6 | 0.48 | 2 n |

he longer arc makes $\mathrm{dp}=0.009$.

| 9623 Hu 909 |  |  | $+61^{\circ} 1500$ |
| :---: | :---: | :---: | :---: |
| $15^{\text {h }} 22^{\text {m }} \cdot 2$ | $+61^{\circ} 10^{\prime}$ |  | 8.3-12.3 FS |
| 62.450 | 293.3 | 1 1.52 |  |
| 64.404 | 287.4 | 1.60 |  |
| 65.212 | 287.7 | 1.50 |  |
| $\underline{65.215}$ | $\underline{289.6}$ | 1.47 |  |
| 64.32 | 289.5 | 1.52 | $4 n$ |
| Hardly changed | in 58 yea |  |  |
| 9628 Hu 149 |  |  | $+54^{\circ} 1745$ |
| $15^{\mathrm{h}} 23^{\mathrm{m}} .2$ | +54 ${ }^{\circ} 23^{\prime}$ |  | 7.5-7.6 |
| 65.105 | 273.5 | 0".60 |  |
| 65.136 65.212 | 271.4 270.6 | 0.46 0.60 |  |
| 65.212 <br> 65.215 | 270.2 <br> 271 | $\begin{array}{r}0.60 \\ 0.49 \\ \hline\end{array}$ |  |
| 65.17 | 271.4 | 0.54 | 4n |

Slow orbital motion corresponding to $\mathrm{dp}=$ 0 '.009.

| 6 A 18 |  |  | $-5^{\circ} 4076$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $15^{\text {h }} 24^{\text {m. }} 0$ | - $5^{\circ} 29^{\prime}$ |  | 8.9-9.4 | F5 |
| 59.550 | $166: 0$ | 0':78 |  |  |
| 59.640 | 170.7 | 0.66 |  |  |
| 60.198 | 168.1 | 0.62 |  |  |
| 60.523 | 170.8 | 0.59 |  |  |
| 60.529 | 168.2 | $\underline{0.63}$ |  |  |
| 60.09 | 168.8 | 0.66 | 5n |  |

$\mathrm{dp}=0.0034$.

| 9643 A 1120 |  |  | $+10^{\circ} 2854$ |
| :---: | :---: | :---: | :---: |
| $15^{\mathrm{h}} 24^{\mathrm{m}} .9$ | $+9^{\circ} 53^{\prime}$ |  | 8.4-9.1 G0 |
| 59.550 | 333:3 | $0!28$ |  |
| 59.656 | 339.0 | 0.33 |  |
| 60.523 | 333.2 | 0.31 |  |
| 60.598 | 336.6 | 0.32 |  |
| 61.423 | 332.8 | 0.21 |  |
| 61.426 | 335.7 | 0.25 |  |
| 61.428 | 334.9 | 0.29 |  |
| 61.483 | 336.3 | 0.28 |  |
| 61.489 | 336.1 | 0.28 |  |
| 62.450 | 336.6 | 0.25 |  |
| $\underline{62.466}$ | 339.2 | 0.29 |  |
| 60.08 | 335.5 | 0.31 | 4 n |
| 61.45 | 335.2 | 0.26 | 5n |
| 62.46 | 337.9 | 0.27 |  |
| Comparison with the residuals: | Muller's | orbit | (1955) gives |
| 60.08 | $+1: 9$ | +0.03 |  |
| 61.46 | + 0.7 | -0.01 |  |
| 62.46 | + 2.8 | -0.01 |  |
| 9645 A 2074 |  |  | +18 ${ }^{\circ} 3024$ |
| $15^{\text {h }} 25{ }^{\text {m }}$. | $+17^{\circ} 48^{\prime}$ |  | 8.2-8.9 F8 |
| 59.550 | 267 \% 1 | $0 \cdot 33$ |  |
| 60.198 | 266.4 | 0.35 |  |
| 60.500 | 271.4 | 0.29 |  |
| 60.519 | 271.6 | 0.29 |  |
| 60.526 | 271.2 | 0.33 |  |





Slow change in angle but marked decrease in distance making $\mathrm{dp}=0$ 0'014.

| 9931 A 1798 |  | $+14^{\circ} 2999$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{h_{5} \mathrm{~m}_{6}}$ | $+14^{\circ} 33^{\prime}$ |  | 8.4-8.9 | F0 |
| 61.483 | $68: 0$ | 0.19 |  |  |
| 61.489 | 64.7 | 0.17 |  |  |
| 61.574 | 64.5 | 0.18 |  |  |
| 64.418 | 58.8 | 0.17 |  |  |
| 62.24 | 64.0 | 0.18 | 4 n |  |

Decrease in both coordinates making $d p=0!005$.

| Finsen 354 |  | $+10^{\circ} 2971$ |  |
| :---: | :---: | :---: | :---: |
| $16^{\text {h }} 9.1$ | $+9^{\circ} 50^{\prime}$ |  | 7.3-7.3 A3 |
| 60.523 | 87.1 | 0.12 |  |
| 60.526 | 89.9 | 0.13 |  |
| 60.529 | 87.2 | 0.13 |  |
| 60.598 | 88.8 | 0.12 |  |
| 60.54 | 88.2 | 0.12 | $4 n$ |
| 9952 A 1799 |  | $+15^{\circ} 2964$ |  |
| $16^{\text {h }} 9.0$ | $+15^{\circ} 15^{\prime}$ |  | 9.2-9.3 G5 |
| 59.539 | 137:8 | 0.45 |  |
| 59.613 | 140.3 | 0.50 |  |
| 59.616 | 140.1 | 0.53 |  |
| 59.663 62.466 | 139.9 134.8 | 0.52 |  |
| 62.543 | 136.3 | 0.45 |  |
| 65.204 | 137.1 | 0.53 |  |
| 65.215 | 135.0 | 0.46 |  |
| 59.61 | 139.5 | 0.50 | 4 n |
| 63.86 | 135.8 | 0.50 | 4n |

To the large change in both coordinates corresponds $\mathrm{dp}=0.4006$.

| 9961 O§ 306 |  |  | $+34^{\circ} 2745$ |
| :---: | :---: | :---: | :---: |
| $16^{\mathrm{h}} 10^{\mathrm{m}} \cdot 0$ | $+34^{\circ} 31^{\prime}$ |  | 8.4-9.9 G5 |
| 59.656 | 25.5 | 0!.33 |  |
| 59.687 | 26.0 | 0.27 |  |
| 60.519 | 26.4 | 0.27 |  |
| 60.523 | 30.4 | 0.32 |  |
| $\underline{60.595}$ | $\underline{29.6}$ | 0.27 |  |
| 60.20 | 27.6 | 0.29 | $5 n$ |
| Very slow angular motion. |  |  |  |
| 9970 इ 2028 |  |  | $+39^{\circ} 2963$ |
| $16^{\mathrm{h}} 11^{\text {m }} 1$ | $+39^{\circ} 29^{\prime}$ |  | 8.7-9.2 G0 |
| 59.422 | 162:9 | 0 O 28 |  |
| 59.656 | 161.0 | 0.32 |  |
| 59.660 | 160.6 | 0.30 |  |
| 59.58 | 161.5 | 0.30 | 3n |
| Slow orbital | tion. |  |  |
| $9974 \times 2023$ |  |  | +5 ${ }^{\circ} 3169$ |



Slow orbital motion. The longer arc makes $d p=$ 0 0.:0020.

Kui 72
$+40^{\circ} 3005$
$16^{\mathrm{h}} 18^{\mathrm{m}} \mathrm{m}_{2}$
$+39^{\circ} 50^{\prime}$
5.5-10.7 F2

| 62.455 |
| :--- |
| 62.466 |
| 65.215 |

$63.38 \quad 140.9 \quad 1.90 \quad 3 n$
Since 1935 the angle has slowly decreased but there is a marked increase in distance. The corresponding $d p$ is $0!013$.

| 10017 Hu 481 |  |  | $+23^{\circ} 2924$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{\text {h }} 19^{\text {m }} 1$ | +23 ${ }^{\circ}{ }^{\prime}$ |  | 8.0-9.9 | F8 |
| 59.634 | $174 \% 4$ | 0 0.77 |  |  |
| 59.672 | 178.0 | 0.76 |  |  |
| 60.526 | 177.5 | 0.63 |  |  |
| 60.592 | 173.9 | $\underline{0.68}$ |  |  |

60.11 176:0 0.71 4n To the change in both coordinates corresponds $\mathrm{dp}=0.010$.

| 10036 - 951 |  |  | $+33^{\circ} 2722$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{\mathrm{h}} 21 \mathrm{~m}_{6}$ | $+33^{\circ} 28^{\prime}$ |  | 9.6-9.8 |  |
|  | $\mathrm{AB}=\mathrm{VB}$ |  |  |  |
| 59.660 | $107: 3$ | 0.20 |  |  |
| 60.526 | 105.3 | 0.24 |  |  |
| 60.595 | 107.2 | 0.25 |  |  |
| 62.455 | 106.4 | 0.26 |  |  |
| 62.466 | 110.0 | 0.26 |  |  |
| 64.418 | 106.1 | 0.25 |  |  |
| 61.69 | 107.0 | 0.24 | $6 n$ |  |
| Slow decrease in angle. |  |  |  |  |
| $\underline{10046} \beta 950$ |  |  | -9\%4381 |  |
| $16^{\mathrm{h}} 22 \mathrm{~m}_{5}$ | - $9^{\circ} 45^{\prime}$ |  | 8.7-9.8 | F0 |
| 61.220 | $349: 4$ | 0'.90 |  |  |
| 61.250 | 350.0 | 1.14 |  |  |
| 61.483 | 347.5 | 0.96 |  |  |
| 61.489 | 349.2 | 1.08 |  |  |
| 61.497 | 350.6 | 0.94 |  |  |
| 61.567 | 349.9 | 1.20 |  |  |
| 61.42 | 349.4 | 1.04 | 6 n |  |
| Very slow decrease in angle. |  |  |  |  |
| $10054 \beta 625$ |  |  | $\omega$ Her |  |
| $16^{\mathrm{h}} 23.1$ | $+14^{\circ} 9^{\prime}$ |  | 4.5-11.5 A0 |  |
| 60.523 | 221:8 | 0!'98 |  |  |
| 60.529 | 224.5 | 0.97 |  |  |
| 60.53 | 223.2 | 0.98 | 2n |  |

This pair becomes very difficult as the distance decreases. To the change in both ccordinates corresponds $\mathrm{dp}=0$ 0'020.

10065
A 1859
$+12^{\circ} 3016$
$16^{\mathrm{h}} 24^{\mathrm{m}} .4$

| 60.523 | 49.4 | 0.16 |  |
| :--- | :--- | :--- | :--- |
| 60.526 | 50.4 | 0.20 |  |
| 60.595 | 47.8 | $\underline{0.20}$ |  |
| 60.55 | 49.2 | 0.19 | $3 n$ |
| nged in 52 | years. |  |  |

Unchanged in 52 years.
Rst 3949
$\nu$ Oph

| $16^{\mathrm{h}} 25^{\mathrm{m}} .1$ | $-8^{\circ} 16^{\prime}$ |  | $4.6-7.8$ |
| :---: | :---: | :---: | :---: |
|  | A 2 |  |  |
| 62.450 | $104: 0$ | $0 . .96$ |  |
| 62.455 | 100.7 | 0.91 |  |
| $\frac{62.553}{62.49}$ | $\frac{99.9}{101.5}$ | $\frac{1.12}{1.00}$ |  |
|  | $3 n$ |  |  |

The angle has increased by $50^{\circ}$ since 1935.
10087

| $16^{\mathrm{h}} 28^{\mathrm{m}} 4$ | $+2^{\circ} 6^{\prime}$ |  | 3.9-6.9 |
| :---: | :---: | :---: | :---: |
| 62.450 | $347: 8$ | 0.91 |  |
| 62.455 | 348.8 | 0.85 |  |
| 62.466 | 346.4 | 0.82 |  |
| 62.543 | 345.6 | 0.99 |  |
| 62.553 | 348.8 | 1.00 |  |
| 62.49 | 347.5 | 0.91 | 5 n |

Rabe's orbit (1948) leaves the residuals $+5: 6$ and $-0!03$.

| 10092 |  |  | $-6^{\circ} 4446$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{h_{29} \mathrm{~m}_{1}}$ | $-6^{\circ} 56^{\prime}$ |  | $7.3-7.3 \quad \mathrm{AO}$ |  |
| 59.151 | 259.3 | 0.13 |  |  |
| 59.687 | 256.6 | 0.16 |  |  |
| 60.500 | 253.0 | 0.13 |  |  |
| 60.523 | 257.7 | 0.13 |  |  |



To the slow increase in angle corresponds $\mathrm{dp}=0.0006$.

| - 953 |  |  |
| :---: | :---: | :---: |
| $16^{\mathrm{h}} 36^{\mathrm{m}} 9$ | $+69^{\circ} 53$ |  |
| 62.689 | $126: 4$ | 0'.38 |
| 62.707 | 129.9 | 0.44 |
| 62.710 | 126.0 | 0.39 |
| 62.70 | 127.4 | 0.40 |

The residuals from Baize's orbit (1953) are $+4: 9$ and $+0!01$.
 are $+12: 7$ and $+0!03$.

| 10165 Hu 487 |  |  | $+22^{\circ} 3007$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{\mathrm{h}} 40.0$ | $+21^{\circ} 57^{\prime}$ |  | 9.9-9.9 F8 |  |
| 59.663 | $44: 7$ | 0:38 |  |  |
| 60.523 | 43.1 | 0.36 |  |  |
| 60.529 | 43.0 | 0.34 |  |  |
| 60.24 | 43.6 | 0.36 | 3 n |  |
| To the slow change corresponds dp $=0!.0045$ |  |  |  |  |
| $10188 \Delta 15$ |  |  | $+43^{\circ} 2639$ |  |
| $16^{\mathrm{h}} 2^{\text {m }}{ }_{4}$ | $+43^{\circ} 34^{\prime}$ |  | 9.1-9.1 | K5 |



Residuals from Baize's orbit (1961) are $+2: 5$ and $+0!.02$.

| 10229 इ 2106 |  |  | $+9^{\circ} 3287$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{\mathrm{h}_{48} \mathrm{~m}_{7}}$ | $+9^{\circ} 30^{\prime}$ |  | 7.0-8.7 | F8 |
| 59.523 | 200:6 | 0'34 |  |  |
| 59.627 | 201.3 | 0.35 |  |  |
| 59.679 | 196.6 | 0.40 |  |  |
| 60.523 | 200.9 | 0.38 |  |  |
| 60.529 | 193.3 | 0.43 |  |  |
| 60.588 | 194.5 | 0.40 |  |  |
| 60.595 | 193.9 | 0.42 |  |  |
| 60.15 | 197.3 | 0.39 | $7 n$ |  |

Long period binary.
10230 O§ 315
$+1^{\circ} 3323$


To the slow orbital motion corresponds dp = 0!.013.

10276
$+57^{\circ} 1716$
$16^{\mathrm{h}_{5}} 5^{\mathrm{m}_{7}} \quad+57^{\circ} 16$
9.8-9.9 G0 " 40
$\begin{array}{lll}61.571 & 123.6 & 0.40 \\ 61.574 & 118.8 & 0.35\end{array}$
$61.580 \quad 121.8 \quad 0.34$
$\frac{61.583}{61.58} \quad \frac{118.4}{120.6} \quad \frac{0.41}{0.37}$
$61.58 \quad 120.6 \quad 0.37 \quad 4 n$
Baize's orbit (1961) gives the residuals $+19: 3$ and $+0!10$.

| 10279 | $\sum 2118$ |  |  | $+65^{\circ} 1159$ |  |
| ---: | :---: | :---: | :---: | :---: | :---: |
| $16^{h_{56} \mathrm{~m}_{2}}$ | $+65^{\circ} 7^{\prime}$ | $6.9-7.4$ | FO |  |  |
| 59.422 | 72.0 | 0.98 |  |  |  |
| 59.656 | 69.8 | 0.97 |  |  |  |
| 59.660 | 69.2 | 0.95 |  |  |  |
| 62.707 | 70.0 | 1.02 |  |  |  |
| 62.710 | 71.3 | 0.99 |  |  |  |
| 62.713 | 69.6 | 1.03 |  |  |  |
| 62.718 | 71.7 | 0.98 |  |  |  |
| 61.37 | 70.5 | 0.99 | $7 n$ |  |  |

Maximum separation in the first quadrant seems still far off.

| $10295 \beta 1298$ |  | $+9^{\circ} 3303$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{\mathrm{h}_{57} \mathrm{~m}_{1} .1}$ | $+9^{\circ} 46^{\prime}$ |  | $8.7-8.9 \quad \mathrm{FO}$ |  |
| 59.550 | $114^{\circ}{ }^{\circ} 1$ | $0^{\prime \prime} .36$ |  |  |
| 59.553 | 108.2 | 0.34 |  |  |


| 59.566 | $110: 0$ | 0.34 |  |
| :--- | :--- | :--- | :--- |
| 59.616 | 113.0 | 0.39 |  |
| $\frac{60.519}{59.76}$ | $\underline{114.5}$ | $\underline{0.41}$ |  |
| 112.0 | 0.37 | $5 n$ |  |

To the slow change in angle corresponds $d p=$ $0!0019$. On three of these nights the difference in magnitude of the components was estimated as 0 . 2 , certainly much less than 1 . 3 given by Burnham.

| 10305 Hu 163 |  |  | -12 ${ }^{\circ} 4641$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $16^{\mathrm{h}_{58} \mathrm{~m}_{4}}$ | $-12^{\circ} 9^{\prime}$ |  | 9.6-9.9 | G0 |
| 60.526 | 339:2 | 0'.21 |  |  |
| 60.529 | 337.7 | 0.20 |  |  |
| 60.598 | 338.9 | 0.21 |  |  |
| 60.55 | 338.6 | 0.21 | $3 n$ |  |

Hardly changed in 60 years.
$10341 \beta 823 \quad+0^{\circ} 3633$

| $17^{\mathrm{h}} \mathrm{m}^{\mathrm{m}} 0$ | $+0^{\circ} 43^{\prime}$ |  | 8.7-9.7 | G0 |
| :---: | :---: | :---: | :---: | :---: |
| 59.566 | $89: 8$ | 0'. 85 |  |  |
| 59.613 | 90.3 | 0.90 |  |  |
| 59.616 | 94.7 | 0.96 |  |  |
| 59.640 | 93.0 | 0.88 |  |  |
| 59.61 | 92.0 | 0.90 | 4n |  |

Arend's orbit (1955) gives the residuals $+2: 3$ and +0.06 .
 $\pm 2: 0$ and $-0!04$.

| 10403 A 2087 |  |  | $+17^{\circ} 31$ |
| ---: | :--- | :--- | :--- |
| $17^{\mathrm{h}} 11^{\mathrm{m} .4}$ | $+17^{\circ} 20^{\prime}$ |  | $9.9-9.9$ |
| 59.660 | 118.4 | $0^{\prime \prime} .14$ |  |
| 60.529 | 115.0 | 0.16 |  |
| 60.598 | 132.2 | 0.14 |  |
| 60.601 | 120.1 | 0.13 |  |
| 60.606 | 126.5 | 0.14 |  |
| 60.609 | 121.8 | 0.15 |  |
| 60.615 | $\underline{130.1}$ | $\underline{0.15}$ |  |
| 60.46 | 123.4 | 0.14 | $7 n$ |

Couteau's orbit (1965) gives the residuals $-4: 8$ and $-0: 01$.

| 10409 A 1147 | $+6^{\circ} 3373$ |
| :---: | :---: |
| $17^{\mathrm{h}} 11^{\mathrm{m} .8}$ | $+6^{\circ} 25^{\prime}$ |$\quad 9.8-10.0 \quad \mathrm{~A} 2$


| 60.523 | 358.5 | $0 . .21$ |  |
| :--- | :--- | :--- | :--- |
| 60.529 | 353.5 | 0.25 |  |
| $\frac{60.532}{60.53}$ | $\frac{354.1}{355.4}$ | $\underline{0.22}$ |  |
|  |  | 0.23 | $3 n$ |

The change is almost all in the distance.

| 10423 A 2592 | $-9^{\circ} 4525$ |
| :---: | ---: |
| $17^{\mathrm{h}} 12^{\mathrm{m} .9}$ | $-9^{\circ}{ }^{\circ} 5^{\prime}$ |$\quad 7.6-8.1$ F5


| 59.550 | 275.2 | 0.45 |  |
| :--- | :--- | :--- | :--- |
| 59.613 | 272.4 | 0.42 |  |
| 59.635 | 273.8 | 0.36 |  |
| $\frac{60.523}{59.83}$ | $\underline{272.1}$ | $\underline{0.39}$ |  |

The change in both coordinates leads to $\mathrm{dp}=0.1007$.

| 5 O 327 |  |  | $+56{ }^{\circ} 1959$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} 13 . \mathrm{m}$ | $+56^{\circ} 11^{\prime}$ |  | 8.5-8.8 | F2 |
| 59.656 | $41: 5$ | 0.15 |  |  |
| 59.660 | 42.6 | 0.14 |  |  |
| 60.523 | 48.6 | 0.15 |  |  |
| 60.526 | 42.5 | 0.13 |  |  |
| 60.529 | 47.8 | 0.16 |  |  |
| 60.598 | 48.3 | 0.16 |  |  |
| 60.601 | 44.9 | 0.15 |  |  |
| 59.66 | 42.0 | 0.14 | $2 \pi$ |  |
| 60.56 | 46.4 | 0.15 | 5 n |  |
| $\begin{gathered} \text { My orbit }(1960) \\ 59.66 \\ 60.56 \end{gathered}$ | $\begin{aligned} & \text { gives } \\ & +2.8 \\ & +2.5 \end{aligned}$ | the resi $+0.103$ $+0.04$ | duals: |  |
| 10435 HW 15 |  |  | $+26^{\circ} 29$ |  |
| $17^{\mathrm{h}} 14^{\mathrm{m}} .6$ | $+26^{\circ} 38^{\prime}$ |  | 9.7-10.7 | G5 |
| AB | ¢ 2145 |  |  |  |
| 59.660 | $27: 2$ | 0.15 |  |  |
| 60.526 | 29.8 | 0.15 |  |  |
| 60.601 | 31.3 | 0.16 |  |  |
| 60.606 | 30.8 | 0.16 |  |  |
| 60.35 | 29.8 | 0.16 | 4 n |  |

The change is mostly in the distance which has closed in considerably. The corresponding dp is $0!0065$.

| $10459 \beta 628$ |  | $+32^{\circ} 28$ |  |
| :---: | :--- | :--- | :--- |
| $17^{\mathrm{h}} 16^{\mathrm{m}} .5$ | $+32^{\circ} 43^{\prime}$ |  | $9.4-9.9$ |
| 59.640 | 309.0 | 01.43 |  |
| 59.660 | 307.3 | 0.42 |  |
| $\underline{59.663}$ | $\underline{306.8}$ | $\underline{0.46}$ |  |
| 59.65 | 307.7 | 0.44 | 3 n |

With little change in distance the angle has decreased by $60^{\circ}$ since 1878 making $d p=$ $0!005$.

10478 Hu 670
$+49^{\circ} 2617$

| $17^{\mathrm{h}} 18^{\mathrm{m}} 0$ | $+49^{\circ} 20^{\prime}$ | 10.7 |
| :---: | :---: | :---: |
| 60.523 | 27.9 | 0.13 |
| 60.526 | 23.5 | 0.14 |
| 60.529 | 22.2 | 0.15 |
| $\frac{60.601}{60.54}$ | $\underline{27.5}$ | $\underline{0.13}$ |
| 25.3 | 0.14 | $4 n$ |

The change is mostly in distance. The period may be short.

| 10480 A 2593 |  |  | $-6^{\circ} 4581$ |
| :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} 18.0$ | $-7^{\circ} 3^{\prime}$ |  | 9.5-9.7 G5 |
| 60.523 | 220.2 | 0.112 |  |
| 60.526 | 219.5 | 0.14 |  |
| 60.529 | 218.8 | 0.14 |  |
| 60.595 | 208.0 | 0.16 |  |
| 60.601 | 219.4 | 0.13 |  |
| 60.55 | 217.2 | 0.14 | $5 n$ |
| Couteau's orbit $+6: 3$ and $+0: 01$. | (1962) | gives | the residuals |
| $10513 \beta 1249$ |  |  | $+53^{\circ} 1938$ |
| $17^{\mathrm{h}} 21 \mathrm{~m}_{0}$ | $\div 53^{\circ} 54^{\prime}$ |  | 10.0-10.2 |
| 59.656 | 100:7 | $0 \% 20$ |  |
| 59.660 | 93.2 | 0.17 |  |
| 60.523 | 102.8 | 0.22 |  |


| 60.526 | 92.4 | 0.20 |  |
| :--- | :--- | :--- | :--- |
| 60.529 | 99.1 0.18  <br> 60.18  97.6 | 0.19 $5 n$ |  |

Change in both coordinates making $d p=0!005$.
10531 Hu 1179
$+38^{\circ} 2928$

| $17^{\mathrm{h}} 22^{\mathrm{m}_{4}}$ | $+38^{\circ} 38^{\prime}$ |  | $7.1-7.7$ | $\mathrm{F8}$ |
| :---: | :--- | :--- | :--- | :--- |
| 59.656 | 249.1 | $0 \prime 11$ |  |  |
| 59.660 | 252.9 | 0.13 |  |  |
| 59.663 | 262.8 | 0.11 |  |  |
| 59.687 | 255.2 | 0.13 |  |  |
| 60.526 | 258.1 | 0.12 |  |  |
| 60.529 | 255.2 | 0.11 |  |  |
| 60.595 | 248.5 | 0.13 |  |  |
| $\frac{60.601}{60.11}$ | $\underline{254.9}$ | $\underline{0.11}$ |  |  |
|  | 254.6 | 0.12 | $8 n$ |  |

To the slow decrease in angle and distance corresponds $\mathrm{dp}=0!0016$.

| 10585 A 351 |  |  | $+29^{\circ} 3029$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} \mathrm{m}^{\mathrm{m}} 4$ | $+29^{\circ} 26^{\prime}$ |  | 9.7-10.1 | K2 |
| 59.640 | $65: 0$ | 0'.55 |  |  |
| 59.656 | 61.6 | 0.59 |  |  |
| 59.660 | 62.4 | 0.60 |  |  |
| 62.450 | 67.5 | 0.49 |  |  |
| 62.466 | 66.8 | 0.53 |  |  |
| 59.65 | 63.0 | 0.58 | $3 n$ |  |
| 62.46 | 67.2 | 0.51 | 2n |  |
| Baize's orbit | (1954) gi | ves the | residuals |  |
| 59.65 | + 2.3 | +0'.07 |  |  |
| 62.46 | + 3.3 | -0.01 |  |  |
| 10617 A 1155 |  |  | $+74^{\circ} 713$ |  |
| $17^{\mathrm{h}} 29^{\mathrm{m}} 9$ | $+74^{\circ} 32^{\prime}$ |  | 8.3-10.9 | F2 |
| 60.523 | 301:0 | 0'.58 |  |  |
| 60.526 | 295.1 | 0.56 |  |  |
| 60.601 | 299.5 | 0.64 |  |  |
| 60.618 | 299.4 | 0.54 |  |  |
| 60.57 | 298.8 | 0.58 | $4 n$ |  |
| Slow orbital | notion mak | ing dp | $=0!.008$. |  |

10621
A 352
$+28^{\circ} 2771$
$17^{\mathrm{h}_{30} \mathrm{~m}_{1}}$
$+28^{\circ} 50^{\prime}$
8.9-8.9 F2

| 59.656 | 189.5 | 0.117 |  |
| :--- | :--- | :--- | :--- |
| 59.660 | 182.2 | 0.15 |  |
| 60.523 | 190.6 | 0.13 |  |
| 60.526 | 186.2 | 0.19 |  |
| $\frac{60.529}{60.18}$ | $\underline{185.7}$ | $\underline{0.20}$ |  |
| 186.8 | 0.17 | $5 n$ |  |

Baize's orbit (1958) gives the residuals $+10: 3$ and +0.101 .

| 10624 Hu 1181 |  |  | $+34^{\circ} 2990$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} 30^{\mathrm{m}} 8$ | $+34^{\circ} 47^{\prime}$ |  | 8.4-8.7 | G0 |
| 59.656 | $6: 8$ | 0.'14 |  |  |
| 59.660 | 11.4 | 0.13 |  |  |
| 60.523 | 10.3 | 0.12 |  |  |
| 60.526 | 10.6 | 0.12 |  |  |
| 60.529 | 8.4 | 0.13 |  |  |
| 60.601 | 12.5 | 0.12 |  |  |
| 60.25 | 10.0 | 0.13 | 6n |  |

The quadrant could not be ascertained at such a small distance. Orbital motion evident but its nature is stili indeterminate.

Kui 83
$+27^{\circ} 2853$
$17^{h_{3}}{ }^{\mathrm{m} .2}$
$+27^{\circ} 55^{\prime}$
11.3-11.5 K5

| 59.656 | 345.2 | 0.26 |
| :--- | :--- | :--- |
| 59.660 | 338.8 | 0.29 |
| 59.663 | 340.5 | 0.25 |
| 61.483 | 314.7 | 0.25 |


| $\begin{array}{r} 61.489 \\ 61.580 \\ \hline \end{array}$ | $\begin{aligned} & 313.7 \\ & 310.3 \\ & \hline \end{aligned}$ | $\begin{aligned} & 0!26 \\ & 0.24 \end{aligned}$ |  |
| :---: | :---: | :---: | :---: |
| 59.66 | 341.5 | 0.27 | 3 n |
| 61.52 | 312.9 | 0.25 | 3 n |
| Baize's orbit | (1963) giv | ves the | residuals: |
| 59.66 61.52 | +6:7 | +0.05 |  |
| 61.52 | + 3.5 | +0.04 |  |
| 10660 - 962 |  |  | 26 Dra |
| $17^{\text {h }} 34.5$ | $+61^{\circ} 55^{\prime}$ |  | 5.3-9.9 F8 |
| 59.627 | 150:1 | 1146 |  |
| 59.632 | 153.9 | 1.40 |  |
| 59.635 | 150.7 | 1.50 |  |
| 59.640 | 152.4 | 1.36 |  |
| 60.601 | 145.3 | 1.33 |  |
| 60.618 | 148.9 | 1.36 |  |
| 59.63 | 151.8 | 1.43 | 4 n |
| 60.61 | 147.1 | 1.35 |  |
| Baize's recent | orbit (19 | 965) ma | akes the |
| residuals: $59.63$ | $+3.1$ | +0'05 |  |
| 60.61 | + 2.3 | +0.08 |  |
| 10683 Hu 181 |  |  | $-15^{\circ} 4635$ |
| $17^{\text {h }} 36{ }^{\text {m }} 5$ | $-15^{\circ} 40^{\prime}$ |  | 9.8-10.2 K0 |
| 60.529 | $35: 8$ | 0.49 |  |
| 60.598 | 33.6 | 0.55 |  |
| 60.601 | 31.8 | 0.60 |  |
| $\underline{60.606}$ | 34.1 | 0.60 |  |
| 60.58 | 33.8 | 0.56 | $4 n$ |
| Orbital motion | evident. |  |  |
| 10686 Es 1257 |  |  | +45 ${ }^{\circ} 2574$ |
| $17^{\text {h }} 36{ }^{\text {m }}$ 8 | +45 ${ }^{\circ} \mathbf{2}^{\prime}$ |  | 10.1-10.6 |
| 61.217 | 295:9 | 2'.72 |  |
| 61.220 | 297.2 | 2.90 |  |
| 61.223 | 298.7 | 2.77 |  |
| 61.240 | 298.1 | 2.88 |  |
| 61.393 | 296.2 | $\underline{2.97}$ |  |
| 61.26 | 297.2 | 2.85 | 5n |
| Slow increase | in angle. |  |  |
| 10702 A 1160 |  |  | $-0^{\circ} 3342$ |
| $17^{\mathrm{h}} 38^{\mathrm{m}} .0$ | - $0^{\circ} 42$ ' |  | 10.3-10.3 A2 |
| 60.588 | 20.1 | 1 1'63 |  |
| 60.592 | 19.4 | 1.71 |  |
| $\underline{60.609}$ | 19.0 | 1.70 |  |
| 60.60 | 19.5 | 1.68 | 3 n |

To the slow increase in both coordinates corresponds $\mathrm{dp}=0.0085$.

| 10705 A 2685 |  | $-5^{\circ} 4480$ |  |
| :---: | :---: | :---: | :---: |
| $17^{\text {h }} 38{ }^{\text {m }} 1$ | - $5^{\circ} 19^{\prime}$ |  | 10.1-10.4 G8 |
| 60.609 | 183.1 | 0'.96 |  |
| 60.615 | 179.5 | 0.86 |  |
| 60.618 | 179.4 | 0.98 |  |
| 60.61 | 180.7 | 0.93 | $3 n$ |
| Slow orbital motion. |  |  |  |
| $\underline{10711}$ AG $211 \quad+20^{\circ} 3540$ |  |  |  |
| $17^{\text {h }} 38{ }^{\text {m }} 5$ | $+20^{\circ} 18^{\prime}$ |  | 9.5-10.0 F8 |
| 60.590 | 130:4 | 2.73 |  |
| 60.598 | 130.5 | $\underline{2.75}$ |  |
| 60.59 | 130.4 | 2.74 | 2 n |
| Unchanged. |  |  |  |
| 10730 AG 212 |  |  | $+5^{\circ} 3457$ |
| $17^{\text {h }} 40{ }^{\text {m }}$. 2 | $+5^{\circ} 21^{\prime}$ |  | 9.9-9.9 F0 |
| 60.588 | 26:3 | 2!.66 |  |


| 60.592 | 27.1 | $2 .!69$ |  |
| :--- | :--- | :--- | :--- |
| 60.595 | 25.7 | 2.54 |  |
| $\underline{60.598}$ | $\underline{26.8}$ | $\underline{2.67}$ |  |
| 60.59 | 26.5 | 2.64 | 4 n |

Unchanged.
10733 Hu $1283 \quad+12^{\circ} 3278$
$17^{\mathrm{h}}{ }_{40} \mathrm{~m}^{\mathrm{m}} 7 \quad+12^{\circ} 17^{\prime} \quad 9.5-10.3 \quad \mathrm{~A} 2$

| 60.609 | 351.4 | 0.37 |  |
| :--- | :--- | :--- | :--- |
| 60.615 | 352.0 | 0.39 |  |
| 60.618 | 348.8 | 0.38 |  |
| 61.574 | 348.7 | 0.34 |  |
| 61.580 | 347.2 | 0.36 |  |
| $\underline{61.588}$ | 351.4 | $\underline{0.37}$ |  |
| 61.10 | 349.9 | 0.37 | $6 n$ |

The slow change indicates a $\mathrm{dp}=0!, 0025$.

| 10734 HI 41 |  | +73 ${ }^{\circ} 786$ |  |
| :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} 40^{\mathrm{m}} \mathrm{m}_{8}$ | +72 ${ }^{\circ} 57^{\prime}$ |  | 9.0-9.4 F2 |
| 60.590 | 341.0 | 1 1'37 |  |
| 60.605 60.616 | $\begin{array}{r}339.2 \\ 338.6 \\ \hline\end{array}$ | 1.38 |  |
| $\underline{60.616}$ | 338.6 | 1.44 |  |
| $60.60$ <br> Unchanged. | 339.6 | 1.40 | 3n |
| 10752 Es 2168 |  | $+36^{\circ} 2931$ |  |
| $17^{\text {h }}{ }_{4} \mathrm{~m}^{\mathrm{m}} 1$ | $+36^{\circ} 57{ }^{\prime}$ |  | 11.0-11.6 G0 |
| 60.588 | 29199 | 6".91 |  |
| $60.595$ | 293.0 293.2 | 6.90 7.05 |  |
| 60.59 | 292.7 | 6.95 | $3 n$ |
| 10755 Hu 1286 |  | $+22^{\circ} 3201$ |  |
| $17^{\mathrm{h}} 42 \mathrm{~m}_{4}$ | $+22^{\circ} 38^{\prime}$ |  | 11.0-11.6 |
| 60.592 | $271{ }^{\circ} 2$ | 3".38 |  |
|  |  |  |  |
| 60.59 | 271.8 | 3.32 | 2n |

Change
$\underline{10756} \Sigma 2206 \quad+19^{\circ} 3416$

| $17^{\mathrm{h}} 42 \mathrm{~m}_{4}$ | $+19^{\circ} 1^{\prime}$ |  |
| :---: | :---: | :---: |
| 60.592 | 249.5 | 1:11 |
| 60.595 | 247.5 | 1.25 |
| 60.601 | 249.9 | 1.23 |
| 60.606 | 249.1 | 1.19 |
| 60.60 | 249.0 | 1.19 |

No definite change after 130 years.




| 60.595 | 218.4 | 2.61 |  |
| :--- | :--- | :--- | :--- |
| 60.601 | 217.6 | 2.94 |  |
| $\frac{60.606}{60.60}$ | $\frac{218.8}{218.3}$ | $\frac{2.79}{2.78}$ |  |
| $3 n$ |  |  |  |

Unchanged. The position is corrected from the astrographic catalogue.

10896
J 458

| $17^{\mathrm{h}} \mathrm{h}_{53^{m} .7}$ | $+8^{\circ} 32^{\prime}$ |  | $9.3-11.0$ |
| :---: | :--- | :--- | :--- |
| 60.595 | $116^{\circ} .2$ | $4^{\prime \prime} .33$ |  |
| 60.601 | 117.5 | 4.48 |  |
| $\frac{60.606}{60.60}$ | $\frac{115.3}{116.3}$ | $\frac{4.32}{4.38}$ | 3 n |

Probably unchanged.

| $\underline{10912}$ \ 2244 |  |  | $+0^{\circ} 3816$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} 54{ }^{\text {m }} 5$ | $+0^{\circ} 4^{\prime}$ |  | 6.8-7.0 | A2 |
| 59.550 | $281: 2$ | 0:38 |  |  |
| 59.566 | 283.9 | 0.34 |  |  |
| 59.613 | 286.6 | 0.36 |  |  |
| 59.656 | 289.4 | 0.34 |  |  |
| 59.663 | 286.7 | 0.32 |  |  |
| 61.483 | 285.8 | 0.30 |  |  |
| 61.497 | 287.7 | 0.32 |  |  |
| 61.551 | 282.2 | 0.36 |  |  |
| 61.556 | 286.0 | 0.30 |  |  |
| 62.450 | 287.6 | 0.29 |  |  |
| 62.466 | 288.7 | 0.31 |  |  |
| 62.553 | 291.6 | 0.31 |  |  |
| $\underline{62.600}$ | 286.5 | 0.29 |  |  |
| 59.61 | 285.6 | 0.35 | 5 n |  |
| 61.52 | 285.4 | 0.32 | $4 n$ |  |
| 62.52 | 288.6 | 0.30 | 4 n |  |
| The longer arc makes $\mathrm{dp}=0!0039$. |  |  |  |  |
| $10916 \beta 1299$ |  |  | $+10^{\circ} 3337$ |  |
| $17^{\mathrm{h}_{55} \mathrm{~m}_{1}}$ | $+10^{\circ} 58^{\prime}$ |  | 8.8-8.8 K5 |  |
| 59.656 | 343.5 | 0.12 |  |  |
| 59.660 | 348.2 | 0.12 |  |  |
| 60.523 | 357.4 | 0.13 |  |  |
| 60.601 | 351.8 | 0.14 |  |  |
| 60.618 | 353.0 | 0.15 |  |  |
| 59.66 | 345.8 | 0.12 | 2n |  |
| 60.58 | 354.1 | 0.14 | 3n ${ }^{3 n}$ residuals: |  |
| $\begin{gathered} \text { Baize's orbit } \\ 59.66 \\ 60.58 \end{gathered}$ | $\begin{gathered} (1959) g i \\ -8.8 \\ -7.9 \end{gathered}$ | $\begin{aligned} & \text { res the } \end{aligned}$ $0!00$ | residu | ls: |
| 10938 Hu 190 |  |  | $-13^{\circ} 4807$ |  |
| $17^{\mathrm{h}_{55} \mathrm{~m} 9}$ | $-13^{\circ} 4^{\prime}$ |  | 9.8-11 | G0 |
| 60.609 | $164: 8$ | 0'!54 |  |  |
| 60.618 | 164.4 | 0.66 |  |  |
| 60.621 | 165.2 | 0.63 |  |  |
| 60.62 | 164.8 | 0.61 | 3 n |  |

Motion in both coordinates making $d p=0!.011$.
Kui 84
$+5^{\circ} 3562$
$17^{\mathrm{h}_{56}} \mathrm{~m}_{0}+4^{\circ} 28^{\prime} \quad 10.9-11.5 \quad \mathrm{~K} 8$

| 60.598 | 350.8 | 0.14 |  |
| :--- | :--- | :--- | :--- |
| 60.606 | 351.5 | 0.15 |  |
| 60.609 | $\frac{355.6}{}$ | $\underline{0.14}$ |  |
| 60.60 | 352.6 | 0.14 | $3 n$ |

No clear evidence of motion so far.

| 10954 Hu 1185 |  |  | $+32^{\circ} 3024$ |
| :---: | :---: | :---: | :---: |
| $17^{\mathrm{h}} 57 \mathrm{~m} .2$ | $+32^{\circ} 28^{\prime}$ |  | 9.6-10.6 |
| 59.656 | $36: 3$ | 0.'12 |  |
| 59.660 | 34.8 | 0.13 |  |
| 59.663 | 36.3 | 0.15 |  |
| 59.66 | 35.8 | 0.13 | 3n |

The period may be short:

10977 ק 47

| $17^{\mathrm{h}_{58} \mathrm{~m} .7}$ | $-10^{\circ} 14^{\prime}$ |  | $8.5-10.5$ | $\mathbf{G 0}$ |
| ---: | :--- | :--- | :--- | :--- |
|  |  | 0.91 |  |  |
| 59.638 | 309.9 | 0.91 |  |  |
| 59.640 | 305.3 | 0.86 |  |  |
| $\frac{59.663}{59.65}$ | $\frac{304.4}{306.5}$ | $\frac{0.84}{0.87}$ | 3 n |  |

Change in both coordinates making $d p=0^{\prime \prime} .0037$.
$10990 \beta 1125 \quad 68$ Oph

| $17^{\mathrm{h}_{59} \mathrm{~m}_{2}}$ | $+1^{\circ} 18^{\prime}$ |  | 4.4-9.2 |
| :---: | :---: | :---: | :---: |
| 59.550 | $65: 3$ | 0.'65 |  |
| 59.566 | 70.8 | 0.70 |  |
| 60.609 | 69.3 | 0.58 |  |
| 59.91 | 68.5 | 0.64 | $3 n$ |

The longer arc makes $d p=0.009$.
$11002 \beta 635$


Very slow angular increase.

show that further correction of the elements is required.

| 11010 | $\beta 1127$ |  |  | +44 ${ }^{\circ} 2812$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | $8^{\mathrm{h}} 1.0$ | +44 ${ }^{\circ} 14^{\prime}$ |  | 7.2-9.3 | F2 |
|  | 62.686 | $88^{\circ} 1$ | 0..83 |  |  |
|  | 62.689 | 90.5 | 0.86 |  |  |
|  | 62.707 | 91.3 | 0.93 |  |  |
|  | 62.710 | 87.2 | 0.90 |  |  |
|  | 62.70 | 89.3 | 0.88 | 4n |  |

The longer arc makes $d p=0: 024$.
11022 Es $1416+44^{\circ} 2813$

$$
18^{h} 1^{\mathrm{m}} 6 \quad+44^{\circ} 42^{\prime} \quad 11.0-11.2
$$

| 60.598 | 72.9 | 1.67 |  |
| :--- | :--- | :--- | :--- |
| 60.601 | 75.4 | 1.80 |  |
| 60.606 | $\frac{76.2}{74.8}$ | $\frac{1.62}{1.70}$ |  |
| 60.60 |  |  |  |

Unchanged.
$\underline{11023} \Sigma 2275 \quad+39^{\circ} 3308$

| $18^{\mathrm{h}} 1^{\mathrm{m}_{6}}$ | $+39^{\circ} 21^{\prime}$ |  | $9.3-9$ |
| ---: | :--- | :--- | :--- |
| 59.651 | 157.6 | 0.131 |  |
| 59.660 | 155.6 | 0.28 |  |
| 60.615 | 155.8 | 0.34 |  |
| 60.621 | 158.6 | 0.28 |  |
| $\frac{64.418}{60.99}$ | $\underline{159.4}$ | $\underline{0.30}$ |  |
| 157.4 | 0.30 | $5 n$ |  |

Larger change especially in angle indicating $\mathrm{dp}=0!.009$. A pair to watch as it closes in. 11060 O 341
$+21^{\circ} 3302$
$18^{\mathrm{h}} \mathrm{hm}_{7}$
$+21^{\circ} 26^{\prime} \quad 7.2-8.5 \quad$ G0

| 59.656 | $91: 8$ | 0.19 |
| :--- | :--- | :--- |
| 59.660 | 96.1 | 0.16 |
| $\underline{59.663}$ | $\underline{95.9}$ | $\underline{0.18}$ |
| 59.66 | 94.6 | 0.18 |
| bit (1951) gives the residuals +6.7 and |  |  |

My orbi
-0.01 .
11067 Но 79
$+33^{\circ} 3025$
$18 \mathrm{~h}_{4} \mathrm{~m}_{3}$
$+33^{\circ} 25^{\prime}$
10.8-10.8

| 59.663 | 22.3 | 0.137 |  |
| :--- | :--- | :--- | :--- |
| 59.687 | 22.1 | 0.40 |  |
| 60.523 | 24.5 | 0.35 |  |
| 60.526 | 26.4 | 0.41 |  |
| $\underline{60.529}$ | $\underline{21.0}$ | $\underline{0.37}$ |  |
| 60.19 | 23.3 | 0.38 | 5 n |

To the increase in angle corresponds $d p=0.0026$.


No definite change.

| $\underline{11080}$ OE 524 | $+19^{\circ} 3533$ |  |  |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 5^{\mathrm{m}} .3$ | $+19^{\circ} 39^{\prime}$ |  | 7.7-9.0 A2 |
| 59.656 | 265:0 | 0.14 |  |
| 59.660 | 271.3 | 0.14 |  |
| 59.663 | 270.5 | 0.13 |  |
| 60.606 | 266.7 | 0.15 |  |
| $\underline{60.609}$ | $\underline{267.8}$ | 0.15 |  |
| 60.04 | 268.3 | 0.14 |  |
| My orbit (1960) $+6: 7$ and $+0 .: 01$. | gives | resi | duals |
| 11098 Hu 314 |  |  | $+18^{\circ} 3566$ |
| $18^{\text {h }} 6 .{ }^{\text {m/4 }}$ | $+18^{\circ} 38^{\prime}$ |  | 8.6-8.7 AO |
| 62.450 | $115: 8$ | 0 O 33 |  |
| 62.466 | 112.2 | 0.34 |  |
| $\underline{62.600}$ | 115.2 | 0.30 |  |
| 62.51 | 114.4 | 0.32 |  |

$11111 \Sigma 2281$
73 Oph
$18^{\mathrm{h}} 7^{\mathrm{m}_{1}}$ $+3^{\circ} 59^{\prime}$ 5.9-7.4
61.571
$\begin{array}{r}61.571 \\ 61.580 \\ 61.588 \\ \hline\end{array}$
61.58

Heintz' orbit (1959) gives the residuals $+16: 6$ and $+0!02$.

| 11127 - 132 |  |  | $-19^{\circ} 4886$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 8^{\text {m }} .3$ | $-19^{\circ} 52^{\prime}$ |  | 6.9-7.3 | A2 |
| 62.689 | 200:1 | 1.42 |  |  |
| 62.707 | 197.8 | 1.60 |  |  |
| 62.718 | 198.8 | 1.67 |  |  |
| 63.334 | 199.7 | 1.55 |  |  |
| 62.86 | 199.1 | 1.56 | 4n |  |

The longer arc makes $\mathrm{dp}=\dot{0}^{\prime \prime} .016$.

| 11149 Ho 82 |  |  | +33 ${ }^{\circ} 3044$ |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} \mathrm{m}_{\text {m }} 9$ | $+33^{\circ} 26^{\prime}$ |  | 6.4-6.7-9.8 A2 |
| $A B=B 2545$ |  |  |  |
| 59.656 | $246: 6$ | 0.12 |  |
| 59.663 | 240.5 | 0.12 |  |
| 60.523 | 246.8 | 0.11 |  |
| 60.526 | 244.3 | 0.13 |  |
| 60.529 | 242.9 | 0.13 |  |
| 60.18 | 244.2 | 0.12 | $5 n$ |
| $A B-C$ |  |  |  |
| 59.656 | 219:1 | 0'.81 |  |
| 59.663 | 218.3 | 0.83 |  |
| 60.523 | 219.8 | 0.75 |  |
| 60.526 | 216.8 | 0.70 |  |
| 60.529 | $\underline{218.5}$ | 0.85 |  |
| 60.18 | 218.5 | 0.79 | 5n |
| The slow change | in $A B-C$ | indica | tes a dp $=00.00$ |

The slow change in $A B-C$ indicates a dp $=00^{\prime \prime}, 006$.
11163 Hu $65 \quad+84^{\circ} 409$
$18^{\mathrm{h}} 10^{\mathrm{m}} 7 \quad+84^{\circ} 36^{\prime} \quad 9.7-10.5$ G5

| 62.689 | 262.4 | $2!.01$ |  |
| :--- | :--- | :--- | :--- |
| 62.704 | 264.8 | 1.89 |  |
| 62.707 | 260.2 | 1.91 |  |
| $\frac{64.736}{63.21}$ | $\frac{263.3}{262.7}$ | $\underline{2.03}$ |  |
| 1.96 | $4 n$ |  |  |

The longer are makes $d p=0!.021$.

| $\beta 1091$ |  |  | $+38^{\circ} 3109$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 10^{\mathrm{m}} 9$ | $+38^{\circ} 35^{\prime}$ |  | 9.4-9.4 | F8 |
| 60.601 | $346: 2$ | 0.631 |  |  |
| 60.615 | 347.6 | 0.34 |  |  |
| 60.61 | 346.9 | 0.32 |  |  |

The slow decrease in both coordinates makes dp $=0!005$.


Hardly changed in 56 years.

| 11225 Hu 1291 |  |  | $+36^{\circ} 3076$ |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 14^{\mathrm{m}} 5$ | $+36^{\circ} 24^{\prime}$ |  | 9.3-10.0 |
| 59.687 | $303: 7$ | 0'.34 |  |
| 60.523 | 300.3 | 0.34 |  |
| 60.526 | 301.2 | 0.31 |  |
| 60.529 | 303.5 | 0.30 |  |
| 62.707 | 300.9 | 0.30 |  |
| 60.79 | 301.9 | 0.32 | 5n |

Both coordinates slowly decrease indicating $\mathrm{dp}=0.006$.

| 11234 A 241 |  |  | $+26^{\circ} 3211$ |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 15^{\mathrm{m}} .2$ | $+26^{\circ} 39^{\prime}$ |  | 10.2-10.5 |
| 59.550 | 289:0 | 0.774 |  |
| 59.640 | $\underline{290.9}$ | 0.59 |  |
| 59.60 | 290.0 | 0.66 | 2n |
| Hardly changed in 58 years. |  |  |  |
| 11239 A 577 |  |  | $+43^{\circ} 2936$ |
| $18^{\mathrm{h}} 15^{\mathrm{m}} \mathrm{m}$ | +43 ${ }^{\circ} 54^{\prime}$ |  | 9.1-12.6 |
| 58.658 | $287: 6$ | 0..88 |  |
| 59.616 | 291.8 | 0.81 |  |
| 59.656 | 286.7 | 0.80 |  |
| 59.660 | 290.5 | 0.86 |  |
| 60.523 | 290.5 | 0.70 |  |
| 59.62 | 289.4 | 0.81 | 5n |

Little change after 56 years.

| 11247 A 578 |  |  | $+43^{\circ} 2938$ |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 16^{\mathrm{m}} 3$ | $+43^{\circ} 49^{\prime}$ |  | 9.2-9.9-13.4 |
|  | AB |  |  |
| 58.658 | 309:7 | $0 \% .23$ |  |
| 59.656 | 309.8 | 0.28 |  |
| 59.660 | 307.0 | 0.27 |  |
| 60.523 | 307.8 | 0.28 |  |
| 60.529 | 306.5 | 0.24 |  |
| 59.81 | 308.2 | 0.26 | 5n |
|  | AB-C |  |  |
| 58.658 | 169:1 | 1'84 |  |
| 59.656 | 168.9 | 1.86 |  |
| 59.660 | 167.8 | 2.01 |  |
| 60.523 | 168.3 | 1.89 |  |
| 60.529 | 169.4 | 1.94 |  |
| 59.81 | 168.7 | 1.91 | 5n |

The $80^{\circ}$ decrease in the angle of the close pair in 57 years makes $\mathrm{dp}=0$ ".0036. There is no appreciable change in $A B-C$.

| 11260 Hu 197 |  |  | $+10^{\circ} 3473$ |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 17^{\mathrm{m}} \cdot 3$ | $+10^{\circ} 15^{\prime}$ |  | 8.5-9.6 G0 |
| 59.550 | 214:2 | 0'.37 |  |
| 59.663 | 211.5 | 0.41 |  |
| 60.598 | 216.0 | 0.44 |  |
| 60.601 | 215.6 | 0.43 |  |
| 62.450 | 212.1 | 0.36 |  |
| 62.466 | 212.5 | 0.44 |  |
| 62.553 | 216.2 | 0.41 |  |
| $\underline{62.600}$ | $\underline{213.0}$ | 0.45 |  |
| 60.10 | 214.3 | 0.41 | 4n |
| 62.52 | 213.4 | 0.42 | 4 n |
| Comparison with | Baize's | orbit | (1955) gives: |
| 60.08 | +199 | -0.'03 |  |
| 62.51 | + 5.9 | -0.02 |  |
| 11284 Ho 430 |  |  | $+20^{\circ} 3750$ |
| $18^{\mathrm{h}} 19^{\mathrm{m}} .5$ | $+20^{\circ} 29^{\prime}$ |  | 9.8-10.3 AS |
| 59.553 | 193.3 | 2.41 |  |
| 59.566 | 194.2 | 2.34 |  |
| 59.613 | 194.1 | 2.46 |  |
| $\underline{59.640}$ | 196.0 | $\underline{2.42}$ |  |
| 59.59 | 194.4 | 2.41 | $4 n$ |
| No change since | 1890. |  |  |
| 11300 Hu 581 |  |  | $+14^{\circ} 3507$ |
| $18^{\text {h }} 20{ }^{\text {m }}$ 6 | $+14^{\circ} 57^{\prime}$ |  | 8.8-9.4 |
| 59.663 | $126: 9$ | 0 0.34 |  |
| 59.687 | 128.8 | 0.33 |  |
| 60.526 60.529 | 126.8 | 0.33 |  |
| 60.529 | 128.0 | 0.32 |  |


| 62.707 | 127:9 | 0.30 |  |
| :---: | :---: | :---: | :---: |
| $60.62 \quad 127.7 \quad 0.32 \quad 5 n$ <br> Residuals from Baize's orbit (1957) are $-3: 3$ and +0.02 . |  |  |  |
|  |  |  |  |
| $\underline{11305}$ A $700 \quad+45^{\circ} 2702$ |  |  |  |
| $18^{\mathrm{h}} 21{ }^{\text {m }} 0$ | +45 ${ }^{\circ} 3^{\prime}$ |  | 10.0-10.1 |
| $\begin{aligned} & 60.523 \\ & 60.529 \\ & 60.601 \\ & \hline \end{aligned}$ | $106: 1$ | 0'.13 |  |
|  | 108.3 | 0.16 |  |
|  | 1.01 .4 | 0.17 |  |
| $60.55 \quad 105.3 \quad 0.15 \quad 3 n$ <br> The uncertainty of the quadrant makes the large change hard to interpret. |  |  |  |
|  |  |  |  |
| 11311 OE 353 |  |  |  |
| $18^{\mathrm{h}} 21^{\mathrm{m}}{ }_{5}$ | $+71^{\circ} 19^{\prime}$ |  | 4.4-6.1 A0 |
| 60.52960.601 | $301 \% 2$ | 0'27 |  |
|  | 305.1 | 0.31 |  |
| 60.606 | 302.0 | 0.28 |  |
| 60.609 | 306.3 | 0.30 |  |
| 60.612 | 306.6 | 0.30 |  |
| 61.567 | 304.7 | 0.28 |  |
| 60.75 | 304.3 | 0.29 | 6n |

Since 1856 the angle of this long period binary has decreased by $120^{\circ}$ and the distance has closed in, but the orbit is still quite indeterminate.

| 11313 Ho 83 |  |  | $+27^{\circ} 3010$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 21^{\mathrm{m}} 5$ | $+27^{\circ} 29^{\prime}$ |  | 9.4-9.4 | A3 |
| 61.912 | $101: 2$ | 0'.80 |  |  |
| 62.553 | 98.4 | 0.79 |  |  |
| 62.686 | 102.5 | 0.85 |  |  |
| 62.710 | 104.1 | 0.77 |  |  |
| 64.736 | 100.6 | 0.98 |  |  |
| 62.92 | 101.4 | 0.84 | 5n |  |

To the slow increase in both coordinates corresponds $\mathrm{dp}=0!012$.



| $18^{\mathrm{h}} 39 \mathrm{~m}_{4}$ | $+30^{\circ} 15^{\prime}$ |  | $7.4-7.9$ | G 5 |
| :---: | :---: | :---: | :---: | :---: |
| 62.707 | 70.2 | 0.34 |  |  |
| 62.710 | 65.4 | 0.33 |  |  |
| 62.713 | 69.5 | 0.36 |  |  |
| $\frac{62.718}{62.71}$ | 66.2 | 0.37 |  |  |
| 67.8 | 0.35 | 4 n |  |  |

The residuals from Baize's orbit (1950) are $+5 \% 8$ and +0.02 .

| $\underline{11584}$ OE 363 |  |  | $+77^{\circ} 702$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\text {h }} 39{ }^{\text {m }} 8$ | $+77^{\circ} 38^{\prime}$ |  | 7.6-7.8 |  |
| 60.529 | $128: 6$ | 0!.16 |  |  |
| 60.601 | 125.6 | 0.20 |  |  |
| 60.606 | 129.3 | 0.15 |  |  |
| 60.609 | 129.7 | 0.16 |  |  |
| 60.59 | 128.3 | 0.17 | 4 n |  |
| Clear orbital motion corresponding to $\mathrm{dp}=$ |  | $\begin{aligned} & \text { Clear orbital motion corresponding to } \mathrm{dp}= \\ & 0.0056 \text {. } \end{aligned}$ |  |  |
| $\underline{11593}$ B 2546 |  |  | $+34^{\circ} 3$ | 285 |
| $18^{\mathrm{h}} 40 \mathrm{~m}_{3}$ | +34 ${ }^{\circ} \mathbf{2}^{\prime}$ |  | 6.5-7.5 | B5 |
| 59.656 | 255.7 | 0.'12 |  |  |
| 59.660 | 249.1 | 0.14 |  |  |
| 59.687 | 258.2 | 0.15 |  |  |
| 60.519 | $\underline{255.6}$ | 0.14 |  |  |
| 59.88 | 254.6 | 0.14 | 4 n |  |
| $\underline{11617} \sum 2369$ |  |  | $+2^{\circ} 36$ |  |
| $18^{\mathrm{h}_{4}} \mathrm{l}^{\mathrm{m}} \mathrm{m}_{4}$ | $+2^{\circ} 34^{\prime}$ |  | 8.2-8.7 | G0 |
| 59.550 | $86 \% 2$ | 0'.59 |  |  |
| 59.553 | 84.4 | 0.67 |  |  |
| 59.556 | 86.4 | 0.56 |  |  |
| 59.632 | 88.4 | 0.64 |  |  |
| 59.640 | 84.3 | 0.63 |  |  |
| 62.543 | 87.3 | 0.63 |  |  |
| 62.689 | 85.0 | 0.69 |  |  |
| 62.702 | 85.0 | 0.64 |  |  |
| 62.718 | 83.2 | 0.64 |  |  |
| 64.736 | 85.9 | $\underline{0.65}$ |  |  |
| 61.33 | 86.0 | 0.63 | 10 n |  |

The longer arc makes $d p=0!010$.

| $\underline{11640}$ ¢ 2375 |  |  | $+5^{\circ} 3941$ |
| :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 43^{\mathrm{m}} \cdot 0$ | + $5^{\circ} 27^{\prime}$ |  |  |
| $\mathrm{AB}=\mathrm{Fin} 332$ |  |  | 6.5-6.7 A0 |
| 60.523 | 140:6 | 0'.14 |  |
| 60.526 | 135.5 | 0.12 |  |
| 60.529 | 138.2 | 0.14 |  |
| 60.598 | 139.3 | 0.15 |  |
| 60.601 | 138.0 | 0.14 |  |
| $\underline{60.606}$ | 137.7 | 0.14 |  |
| 60.56 | 138.2 | 0.14 | 6 n |
| $C D=F i n 332$ |  |  | 7.5-7.5 A0 |
| 60.523 | $130 \% 8$ | 0 O 13 |  |
| 60.526 | 135.8 | 0.12 |  |
| 60.529 | 133.9 | 0.12 |  |
| 60.598 | 133.4 | 0.14 |  |
| 60.601 | 130.4 | 0.14 |  |
| $\underline{60.606}$ | 130.4 | 0.14 |  |
| 60.56 | 132.4 | 0.13 | 6 n |
| 11680 Hu 1191 |  |  | $+38^{\circ} 3292$ |
| $18^{\text {h }} 44{ }^{\text {m }}$. 9 | $+38^{\circ} 18^{\prime}$ |  | 8.6-9.1 G5 |
| 59.660 | $316: 5$ | 0'.27 |  |
| 60.519 | 318.8 | 0.25 |  |
| 60.523 | 318.5 | 0.28 |  |
| 60.526 60.529 | 317.7 320.3 | 0.28 0.26 |  |
| 60.529 60.601 | 320.3 312.7 | 0.28 |  |
| 60.621 | 315.4 | 0.28 |  |



Comparison with my orbit (1960) gives the residuals $-0: 4$ and +0.03 .

\[

\]

No certain change.
11687 Hu $252 \quad+9^{\circ} 3873$
$18^{\mathrm{h}_{4} 5^{\mathrm{m}} .4}+9^{\circ} 12^{\mathrm{I}} \quad$ 9.2-9.7 $\quad$ B9

| 60.598 | 159.3 | 0.14 |
| :--- | :--- | :--- |
| 60.601 | 152.7 | 0.12 |
| $\underline{60.606}$ | $\underline{158.1}$ | $\underline{0.12}$ |
| 60.60 | 156.7 | 0.13 |

The longer arc makes $\mathrm{dp}=0.0057$.
11717 Hu 936
$+33^{\circ} 3212$

| $18^{\mathrm{h}}{ }_{46} \mathrm{~m}_{9}$ | $+33^{\circ} 58^{\prime}$ |  | $9.4-9.7$ |
| :---: | :---: | :---: | :---: |
|  |  | 101.2 | 1.70 |
| 60.592 | 100.2 | 1.84 |  |
| 60.595 | 101.3 | 1.79 |  |
| 60.598 | 102.1 | 1.82 |  |
| 60.601 | 100.8 | 1.86 |  |
| $\frac{60.606}{60.60}$ | 101.1 | 1.80 | $5 n$ |
| 60 in 56 years. |  |  |  |

11769 Hu $199+11^{\circ} 3642$

| $18^{\mathrm{h}} 499^{\mathrm{m}} 8$ | $+11^{\circ} 44 \mathrm{r}$ | $9.1-9$ |  |
| :---: | :--- | :--- | :--- |
| 59.613 | 357.0 | 0.70 |  |
| 59.679 | 355.3 | 0.666 |  |
| 62.543 | 354.4 | 0.77 |  |
| 62.553 | 350.8 | 0.65 |  |
| 62.689 | 349.9 | 0.74 |  |
| 64.730 | 356.8 | 0.80 |  |
| 64.736 | $\underline{352.0}$ | $\underline{0.80}$ |  |
| 62.36 | 353.7 | 0.73 | $7 n$ |

To the increase in distance corresponds dp $=0!.0094$.

| 11791 A 93 |  |  | $-5^{\circ} 4798$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}_{51}}{ }^{\mathrm{m}} 0$ | $-5^{\circ} 36^{\prime}$ |  | 9.4-10.0 | G5 |
| 61.571 | $127: 6$ | 0.65 |  |  |
| 61.580 | 130.1 | 0.55 |  |  |
| 61.583 | 129.0 | 0.62 |  |  |
| 61.71 .0 | 124.5 | 0.68 |  |  |
| 61.61 | 127.8 | 0.62 | $4 n$ |  |

Wilson's orbit (1954) gives the deviations $+8: 7$ and +0.103 .

11818 J 1275

| $18^{\mathrm{h}_{52} \mathrm{~m}_{5}}$ | $+7^{\circ} 11^{\prime}$ | $9.5-10.5$ |  |
| :---: | :--- | :--- | :--- |
| 61.551 | 200.1 | 2.112 |  |
| 61.567 | 202.1 | 2.19 |  |
| 61.580 | 201.0 | 2.08 |  |
| $\frac{61.588}{61.57}$ | $\frac{200.3}{200.9}$ | $\frac{2.18}{2.14}$ | 4 n |

No certain change.
11842 A 2192
$+3^{\circ} 3836$

| $18^{h_{53}}{ }^{m_{3}} 3$ | $+3^{\circ} 23^{\prime}$ | $7.7-7.7$ | A2 |
| :---: | :--- | :--- | :--- |
| 59.663 | 141.3 | 0.28 |  |
| 59.687 | 145.7 | 0.27 |  |
| 60.500 | 143.9 | 0.27 |  |
| 60.53 | 142.8 | 0.29 |  |
| 60.523 | 145.7 | 0.29 |  |
| $\underline{60.526}$ | $\underline{142.9}$ | $\underline{0.27}$ |  |


longer arc reduces dp co 0.0029
11897

## $\sum 2438$

$+58^{\circ} 1849$

| $18^{\mathrm{h}_{56} \mathrm{~m}_{6}}$ | $+58^{\circ} 9$ |
| ---: | ---: |
| 59.613 | 10.9 |
| 59.616 | 12.3 |
| 59.620 | 10. |
| 59.640 | 12. |
| 59.676 | 13. |
| $\frac{60.592}{59.78}$ | 12.8 |
| two orbits by J1 |  |

6.8-7.4 A2

| 59.613 | 10.9 | $0!.92$ |  |
| :--- | :--- | :--- | :--- |
| 59.616 | 12.3 | 1.01 |  |
| 59.620 | 10.2 | 1.06 |  |
| 59.640 | 12.1 | 0.90 |  |
| 59.676 | 13.1 | 0.90 |  |
| $\frac{60.592}{59.78}$ | $\frac{12.8}{11.9}$ | $\frac{0.91}{0.95}$ |  |
|  | 6n |  |  |

The two orbits by Jastrzebski (1958 and 1959) give the same residuals $+5: 8$ and +0.14 .

| 11923 Mlb 414 |  |  | $+67^{\circ} 1105$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 58.1$ | $+67^{\circ} 11^{\prime}$ |  | 10.1-11.3 | G5 |
| 59.616 | $148: 2$ | 1.'93 |  |  |
| 59.638 | 147.8 | 1.94 |  |  |
| 59.641 | 148.9 | $\underline{2.15}$ |  |  |
| 59.63 | 148.3 | 2.01 | $3 n$ |  |
| Change question | ble. |  |  |  |

11939 J 1280

| $18^{\mathrm{h}_{5}} \mathrm{~m}_{\mathrm{m}} .9$ | $+22^{\circ} 1^{\prime}$ |  | $9.7-9.7$ |
| ---: | :--- | :--- | :--- |
| 60.598 | 108.6 | $3^{.177}$ |  |
| $\frac{60.609}{60.60}$ | $\frac{108.9}{108.8}$ | $\frac{3.84}{3.80}$ | 2 n |

Unchanged.
The star was identified in the Paris astrographic catalogue as $+22^{\circ}$, $18^{\mathrm{h}} 56^{\mathrm{m}}$, No. 437 and $+21^{\circ}, 19 \mathrm{~h}^{\mathrm{m}}$, No. 18.

11947 及 973
$+8^{\circ} 3945$

| $18^{\mathrm{h}} 59{ }^{\text {m }} 3$ | $+8^{\circ} 40^{\prime}$ |  | 9.9-12.8 |
| :---: | :---: | :---: | :---: |
| 58.582 | $350: 4$ | 1:48 |  |
| 60.601 | 351.7 | 1.52 |  |
| 64.730 | 352.9 | 1.54 |  |
| 61.30 | 351.7 | 1.51 | 3 n |
| $C D$ | Howe 45 |  | 12.3-12.8 |
| 60.601 | $260: 2$ | 3'.09 |  |
| 64.730 | 260.2 | 3.23 |  |
| $\underline{64.736}$ | 259.0 | 3.24 |  |
| 63.36 | 259.8 | 3.19 | $3 n$ |

No definite change in either pair.

| $\underline{11956}$ ¢ 2437 |  |  | $+18^{\circ} 3920$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $18^{\mathrm{h}} 59 \mathrm{~m}_{7}$ | $+19^{\circ} 6^{\prime}$ |  | 8.2-8.4 | G5 |
| 59.553 | $43: 5$ | 0'.70 |  |  |
| 59.556 | 39.1 | 0.70 |  |  |
| 59.613 | 39.1 | 0.80 |  |  |
| 59.627 | 41.1 | 0.73 |  |  |
| 59.640 | 43.5 | 0.70 |  |  |
| 59.60 | 41.3 | 0.73 | 5 n |  |

The longer arc makes $\mathrm{dp}=0.1008$.
11960 A 1388
$+52^{\circ} 2321$

| $18^{\mathrm{h}_{59} \mathrm{~m}_{9} 9}$ | $+53^{\circ} 7^{\prime}$ |  | $9.1-13.6$ |
| ---: | :--- | ---: | :--- |
|  |  |  |  |
| 60.609 | 261.4 | 5 ".09 |  |
| $\frac{60.618}{60.61}$ | $\frac{261.1}{261.2}$ | $\frac{5.19}{5.14}$ | 2 n |

Unchanged. The 0'. 058 yearly proper motion proves the physical connection.

$0!011$

| 11979 Hu 757 |  |  | +52 ${ }^{\circ} 2326$ |
| :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 0 \mathrm{~m}_{.9}$ | $+52^{\circ} 11^{\prime}$ |  | 9.2-9.8 |
| BC |  |  |  |
| 59.660 | $346: 3$ | 0.'11 |  |
| 60.529 | 346.0 | 0.11 |  |
| 60.601 | 343.8 | 0.11 |  |
| 60.606 | 341.7 | 0.12 |  |
| $\underline{60.618}$ | 342.9 | 0.11 |  |
| 60.40 | 344.1 | 0.11 |  |

To the large change in both coordinates corresponds $\mathrm{dp}=0!.0053$.

| 11998 A 2992 |  |  | +26 ${ }^{\circ} 3437$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 1^{\mathrm{m}} \mathrm{B}$ | $+26^{\circ} 37^{\prime}$ |  | 9.8-9.8 | F8 |
| 60.529 | 118.3 | 0.19 |  |  |
| 60.606 | 115.3 | 0.24 |  |  |
| 60.609 | 118.9 | 0.19 |  |  |
| 60.58 | 117.5 | 0.21 | 3 n |  |

Orbital motion of indeterminate nature so far.
12013 J 1209

| $19^{\mathrm{h}} \mathrm{m}^{\mathrm{m}} 6$ | $+34^{\circ} 2^{\prime}$ |  | 9.5-10.0 |
| :---: | :---: | :---: | :---: |
| 60.598 | 155:4 | 4.76 |  |
| 60.615 | 155.2 | 4.88 |  |
| 60.61 | 155.3 | 4.82 | 2n |
| ed. The two components are in the |  |  |  |
|  |  |  |  |

astrographic catalogue $33^{\circ} 57440$ and 57491.


To the large increase in distance would correspond a $\mathrm{dp}=0 . \mathbf{D}^{\prime} 025$. Probably optical.

12123
A 150
$+20^{\circ} 4067$
$19^{\mathrm{h}} \mathrm{B}_{.} \mathrm{m}_{2}$
$+20^{\circ} 25^{\prime}$
9.5-9.6 A0

| 59.640 | 113.3 | $0!.46$ |
| :--- | :--- | :--- |
| 59.687 | 112.9 | 0.52 |
| 60.529 | 110.5 | 0.53 |

$$
\frac{60.606}{60.12} \quad \frac{110.9}{111.9} \quad \frac{0.651}{0.50} 4 \mathrm{n}
$$

To the increase in both coordinates corresponds $\mathrm{dp}=0$ 0.004


Motion in both coordinates making $\mathrm{dp}=0$ 0'0018. 12145 Secchi 2
$+38^{\circ} 3466$

| $19^{h_{9} \mathrm{~m}_{5}}$ |  | $+38^{\circ} 42^{\prime}$ |  | $8.3-9.3$ |
| :---: | :---: | :---: | :---: | :---: |
|  | BC |  | K |  |
| 59.656 | 243.8 | 0.13 |  |  |
| 59.660 | 238.6 | 0.14 |  |  |
| $\frac{60.523}{59.95}$ | $\frac{241.1}{241.2}$ | $\frac{0.12}{0.13}$ | $3 n$ |  |

Baize's orbit (1961) gives the residuals +21:7 and +0.02 showing that the orbit has to be improved.
$12147 \beta 1204$
$+2^{\circ} 3815$
$19^{\mathrm{h}} \mathrm{g}_{\mathrm{m}} .5$

| 59.687 | 189.7 | 0.30 |  |
| :--- | :--- | :--- | :--- |
| 60.615 | 187.5 | 0.28 |  |
| 60.774 | 193.5 | 0.31 |  |
| 62.689 | 193.7 | 0.32 |  |
| 62.701 | 192.0 | 0.34 |  |
| $\frac{62.704}{61.53}$ | $\underline{193.1}$ | $\underline{0.33}$ |  |
| 191.6 | 0.31 | $6 n$ |  |

I never felt doubt about the quadrant and estimated $\Delta m=0.3$. The change is very slow.

12166
A 264
$+24^{\circ} 3677$

| $19^{\mathrm{h}} 10^{\mathrm{m}_{6}}$ |  | $+24^{\circ} 29$ ' |  | $8.0-13.5$ |
| ---: | ---: | ---: | ---: | ---: |
|  | F5 |  |  |  |
| 58.655 | 289.4 | 3.45 |  |  |
| 60.529 | 290.9 | 3.43 |  |  |
| 60.615 | 288.7 | 3.49 |  |  |
| $\frac{62.710}{60.63}$ | $\frac{289.3}{289.6}$ | $\frac{3.60}{3.49}$ | $4 n$ |  |

Marked increase in distance without change in angle.

| AC |  | $8.0-12.5$ |  | F5 |
| :---: | :--- | :--- | :--- | :--- |
| 58.655 | 6.3 | $2!.28$ |  |  |
| $\frac{60.615}{59.63}$ | $\frac{4.8}{5.6}$ | $\frac{2.43}{2.36}$ | $2 n$ |  |

Probably optical system.
12187 Hu 335
$+19^{\circ} 3949$

$$
\begin{array}{rlrr}
19^{\mathrm{h}} 11^{\mathrm{m}} .4 & +20^{\circ} 7 & & 8.0 . \\
58.655 & 222: 4 & 0.60 & \\
\frac{62.710}{60.68} & \frac{219.9}{221.2} & \frac{0.53}{0.56} & 2 \mathrm{n}
\end{array}
$$

$$
8.0-11.7
$$

Hardly changed since 1901.
$12201 \sum 2484$
$+18^{\circ} 3998$
$19^{\mathrm{h}} 12^{\mathrm{m}} .1$
$+18^{\circ} 59^{\prime}$
7.9-9.4 F8

| 59.553 | 231.9 | $2!39$ |  |
| :--- | :--- | :--- | :--- |
| 59.556 | 231.3 | 2.27 |  |
| 59.627 | 234.9 | 2.38 |  |
| $\frac{59.640}{59.59}$ | $\underline{233.8}$ | $\underline{233.38}$ |  |
|  | 23.0 | 2.36 | 4 n |

The longer arc makes $d p=0!013$.
12246 इ 2491
$+28^{\circ} 3268$

| $19^{\mathrm{h}} 14^{\mathrm{m}} .2$ | $+28^{\circ} 11^{\prime}$ |  | $8.4-9.7$ |
| ---: | :--- | :--- | :--- |
|  | A2 |  |  |
| 59.556 | $224^{\circ}: 2$ | $1^{\prime \prime} \cdot 28$ |  |
| 59.638 | 226.4 | 1.24 |  |
| $\frac{59.687}{59.63}$ | $\frac{226.6}{225.7}$ | $\frac{1.23}{1.25}$ |  |
| longer arc confirms dp $=0!.0061$. |  |  |  |

The longer arc confirms $\mathrm{dp}=0,0061$.

| 12258 A 363 |  | $+7^{\circ} 4011$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{h} 14^{m} .6$ | $+7^{\circ} 19^{\prime}$ | $8.6-13.5$ |  |  |$\quad$ B9

No definite change.

| 12261 A 1392 |  | $+54^{\circ} 2113$ |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 14^{\mathrm{m}} \mathrm{m}_{8}$ | $+54^{\circ} 52^{\prime}$ |  | $8.8-9.0$ | KO |
| 59.656 | 83.9 | 0.43 |  |  |
| 59.660 | 79.4 | 0.42 |  |  |
| 59.687 | 83.8 | 0.43 |  |  |
| 60.523 | 80.3 | 0.40 |  |  |
| 60.526 | 83.1 | 0.44 |  |  |
| 60.609 | 78.3 | 0.43 |  |  |
| 60.612 | 78.3 | $\underline{0.41}$ |  |  |
| 60.18 | 81.0 | 0.42 | 7 n |  |

To the slow change in both coordinates corresponds $\mathrm{dp}=0.1004$.

| 12274 A 1176 |  |  | $+9^{\circ} 4051$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 15^{\mathrm{m}} 3$ | $+10^{\circ} 9^{\prime}$ |  | 9.8-10.8 | A0 |
| 60.609 | 106:0 | 1!.02 |  |  |
| $\underline{60.615}$ | 105.0 | 1.05 |  |  |
| 60.61 | 105.5 | 1.04 | $2 \pi$ |  |
| Unchanged in | years. |  |  |  |

$12296 \Sigma 2509$
$+62^{\circ} 1702$

| $19^{\mathrm{h}} 16^{\mathrm{m}} 4$ | $+63^{\circ} 7^{\prime}$ |  | 7.2-8.3 |
| :---: | :---: | :---: | :---: |
| 59.613 | $331: 8$ | 1.'56 |  |
| 59.616 | 333.0 | 1.66 |  |
| 59.627 | 334.2 | 1.77 |  |
| 59.640 | 332.7 | 1.66 |  |
| 59.62 | 332.9 | 1.66 | 4n |

The longer arc makes $\mathrm{dp}=\dot{0}!015$.

| 12366 ק 1129 |  |  | $+52^{\circ} 2400$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 20^{\mathrm{m}} 4$ | +52 ${ }^{\circ} 17{ }^{\prime}$ |  | 7.7-7.7 | A5 |
| 59.660 | 322.4 | 0'.23 |  |  |
| 60.526 | 323.7 | 0.19 |  |  |
| 60.601 | 321.8 | 0.24 |  |  |
| 60.609 | 320.5 | 0.21 |  |  |
| 60.615 | 320.4 | 0.24 |  |  |
| 60.621 | 322.7 | 0.23 |  |  |
| 60.44 | 321.9 | 0.22 | 6n |  |

To the slow decrease in both coordinates corresponds dp $=0!004$.

| $\underline{12412}$ O§ 373 |  |  | $+46^{\circ} 2681$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 22 \mathrm{~m}_{6}$ | +46 ${ }^{\circ} 20^{\prime}$ |  | 7.4-10.8 | A0 |
| 59.616 | $234: 4$ | 2':11 |  |  |
| 59.627 | 235.2 | 2.12 |  |  |
| 59.638 | 235.4 | 2.14 |  |  |
| 61.220 | 235.8 | 2.23 |  |  |


| 61.567 | 235:3 | $\underline{2.109}$ |  |
| :---: | :---: | :---: | :---: |
| 60.33 | 235.2 | 2.14 | 5 n |
| Hardly changed in 103 |  |  |  |
| 12452 A 1181 |  |  | +11 ${ }^{\circ} 3842$ |
| $19^{\text {h }} 24 \mathrm{~m}^{\mathrm{m}}$ | $+11^{\circ} 58{ }^{\prime}$ |  | 6.9-9.1 A0 |
| 61.710 | 199:0 | 0 0.72 |  |
| 61.771 62.543 | 198.7 | 0.75 |  |
| 62.01 | 198.4 | 0.72 | 3 n |

To the increase in distance corresponds dp = 0'.0027.

| A 1650 |  |  | $+15^{\circ} 382$ |
| :---: | :---: | :---: | :---: |
| $19^{\text {h }} 26^{\text {m }}$. 1 | $+16^{\circ} 11^{\prime}$ |  | 9.5-9.5 |
| $\begin{aligned} & 60.529 \\ & 60.606 \\ & \hline \end{aligned}$ | $\begin{array}{r} 56: 7 \\ 55.2 \\ \hline \end{array}$ | $\begin{aligned} & 0!17 \\ & 0.17 \\ & \hline \end{aligned}$ |  |
| 60.57 | 56.0 | 0.17 | 2n |

Change immaterial after 53 years. 12515 A 1653
$+12^{\circ} 3929$
$19^{\mathrm{h}} \mathrm{h}^{\mathrm{m}} \mathrm{m}_{3}$

$$
+12^{\circ} 18^{\prime}
$$

8.0-9.2 A3

| 60.609 | 226.0 | $0!22$ |  |
| :--- | :--- | :--- | :--- |
| $\frac{60.618}{60.61}$ | $\frac{222.2}{224.1}$ | $\frac{0.25}{0.24}$ | 2 n |

Without change in distance the angle has decreased by $78^{\circ}$ since 1907 making $\mathrm{dp}=$ 0'"003.

12552 A 712
$+56^{\circ} 2250$
$19^{\mathrm{h}} 29^{\mathrm{m}} \mathrm{m}_{2}$
$+56^{\circ} 32^{\prime}$
7.3-7.8 AO

| 59.656 | 108.9 | 0.13 |  |
| :--- | :--- | :--- | :--- |
| $\frac{59.660}{59.66}$ | $\underline{109.4}$ | $\underline{0.14}$ |  |
| 109.2 | 0.14 | 2 n |  |

It is difficult to ascertain the quadrant in such a close pair. The period is probably of the order of a century.

12557

## $\Sigma 2536$

$+17^{\circ} 3992$
$19^{\mathrm{h}} 29^{\mathrm{m}} \mathrm{m}_{4}$
$+17^{\circ} 41^{\prime}$
8.4-11.4 G0

| 59.638 | 98.6 | $2!04$ |
| :--- | :--- | :--- |
| 59.640 | 98.1 | 2.02 |
| 59.646 | 95.6 | 1.88 |
| 60.539 | 96.6 | 1.99 |
| $\frac{60.609}{60.01}$ | 96.3 | $\underline{1.89}$ |
| 97.0 | 1.96 |  |

Without change in distance the angle has in creased by $60^{\circ}$ since 1831 making $d p=0 \% 019$.

| 12567 A 713 |  |  | $+47^{\circ} 2854$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 29 \mathrm{~m}_{8}$ | $+47^{\circ} 22^{\prime}$ |  | 7.7-8.2 | A3 |
| 59.687 | $254: 4$ | 0'.48 |  |  |
| 60.687 | 252.4 | 0.44 |  |  |
| 60.526 | 256.6 | 0.45 |  |  |
| 60.609 | 255.5 | 0.46 |  |  |
| 60.38 | 254.7 | 0.46 | $4 n$ |  |

To the increase in both distance and angle corresponds $\mathrm{dp}=0$ "'009.

| 12577 Hu 951 |  |  | $+63^{\circ} 1530$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}_{3}} \mathrm{~m}_{3}$ | $+63^{\circ} 31^{\prime}$ |  | 9.4-9.6 | F |
| 59.656 | 59.1 | 0.16 |  |  |
| 59.660 | 60.4 | 0.16 |  |  |
| 60.526 | 59.1 | 0.14 |  |  |
| 60.609 | 61.7 | 0.14 |  |  |
| 60.11 | 60.1 | 0.15 | 4 n |  |

Fairly rapid orbital motion. The longer arc reduces dp to $0!$ ', 006.

12600 Ho 108
$+33^{\circ} 3499$
$19^{\mathrm{h}_{3} 1^{\mathrm{m}} 4} \quad+33^{\circ} 22^{\prime} \quad 9.0-9.0 \quad \mathrm{~F} 8$
$\begin{array}{llll}60.609 & 46.2 & 0.19 & \\ \frac{60.618}{60.61} & \frac{44.0}{45.1} & \frac{0.17}{0.18} & 2 \mathrm{n}\end{array}$
The measures are too scarce to bring out the nature of the motion.

12618 A 597
$+42^{\circ} 3382$
$\begin{array}{cllll}19^{h_{32}}{ }^{\mathrm{m}_{2}}{ }^{2} & +42^{\circ} 15^{\prime} & & 8.4-10.9 & \text { GO } \\ 62.689 & 108.3 & 1: 82 & & \\ 62.701 & 111.1 & 1.80 & & \\ 62.704 & 109.1 & 1.91 & & \\ \frac{62.710}{62.70} & \frac{108.4}{109.2} & \frac{1.92}{1.86} & 4 \mathrm{n} & \end{array}$
To the change in both coordinates corresponds $\mathrm{dp}=0!.035$. The 0'!075 proper motion establishes the physical connection.

| 12631 A 162 |  |  | $+23^{\circ} 3699$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 32.9$ | $+23^{\circ} 22^{\prime}$ |  | 8.9-8.9 | AO |
| 58.601 | 231:7 | 0.19 |  |  |
| 60.606 | 235.7 | 0.22 |  |  |
| 60.609 | 233.6 | 0.20 |  |  |
| 60.615 | 231.9 | 0.23 |  |  |
| 60.11 | 233.2 | 0.21 | 4n |  |

The angle has increased by a quadrant in 60 years. The longer arc makes $d p=0!003$.
12648 A 163
$+22^{\circ} 3738$

| $19^{\mathrm{h}} 33^{\mathrm{m}} .^{6}$ | $+23^{\circ} 9^{\prime}$ |  | $9.8-9.9$ | $\mathrm{F5}$ |
| ---: | :--- | :--- | :--- | :--- |
|  |  |  |  |  |
| 58.601 | 158.3 | 0.16 |  |  |
| 60.609 | 154.1 | 0.12 |  |  |
| $\frac{60.618}{59.94}$ | $\frac{155.9}{156.1}$ | $\frac{0.14}{0.14}$ | $3 n$ |  |

The marked decrease in both coordinates makes $\mathrm{dp}=0!005$.


| 62.543 | 204:1 | 0:77 |  |
| :---: | :---: | :---: | :---: |
| 62.633 | 206.6 | 0.58 |  |
| 62.701 | 206.4 | 0.66 |  |
| 62.704 | 203.3 | 0.67 |  |
| $\underline{62.710}$ | 204.4 | 0.64 |  |
| 62.66 | 205.0 | 0.66 | 5n |
| The slow change $\mathrm{dp}=0.012$. | in both | coordi | inates makes |
| $\underline{12752} \sum 2556$ |  |  | +21 ${ }^{\circ} 3862$ |
| $19^{\text {h }} 37{ }^{\text {m }} 3$ | $+22^{\circ} 8^{\prime}$ |  | 7.7-8.2 F2 |
| 62.543 | 67.5 | $0 \cdot 41$ |  |
| 62.689 | 66.0 | 0.40 |  |
| 62.701 | 68.8 | 0.38 |  |
| 62.704 | \% 5.5 | 0.42 |  |
| $\underline{62.710}$ | 67.2 | 0.35 |  |
| 62.67 | 67.0 | 0.39 | 5n |
| Gunzel-Lingner's residuals $+13: 7$ | s orbit and $+0!3$ | (1956) | gives the |
| $\underline{12803} \Sigma 2574$ |  |  | $+62^{\circ} 1747$ |
| $19^{\text {h }} 40^{\text {ma }} 0$ | $+62^{\circ} 33^{\prime}$ |  | 8.1-8.1 F5 |
| 58.650 | $223: 9$ | 0'28 |  |
| 59.656 | 225.3 | 0.30 |  |
| 59.660 | 227.7 | 0.31 |  |
| 59.687 | 225.2 | 0.26 |  |
| 60.601 | 229.4 | 0.30 |  |
| 60.609 | 224.0 | 0.29 |  |
| 61.567 | 227.7 | 0.30 |  |
| 61.574 | 233.8 | 0.30 |  |
| 61.580 | 230.5 | 0.34 |  |
| $\underline{61.588}$ | $\underline{231.1}$ | 0.35 |  |
| 59.81 | 225.9 | 0.29 | 6 n |
| 61.58 | 230.8 | 0.32 |  |
| The change since | e 1832 m | makes dp | $\mathrm{p}=0!0095$. |
| Kui 94 |  |  | $+39^{\circ} 3878$ |
| $19^{\mathrm{h}} 40 \mathrm{~m}_{2}$ | $+40^{\circ} 8^{\prime}$ |  | 6.5-7.8 A3 |
| 59.656 | $134: 8$ | 0'.25 |  |
| $\underline{60.618}$ | 135.4 | 0.28 |  |
| 60.14 | 135.1 | 0.26 | 2n |

The change since 1937 makes $\mathrm{dp}=00.011$.
Kui $95=$ Ross 165

| $19^{\mathrm{h}} 43^{\mathrm{m}} \mathrm{m}_{8}$ | $+27^{\circ} 2^{\prime}$ | $12.7-13.6 \mathrm{Ma}$ |  |
| ---: | :--- | :--- | :--- |
| 58.660 | 242.6 | 0.94 |  |
| 59.656 | 246.1 | 0.80 |  |
| $\frac{60.618}{59.64}$ | $\underline{247.4}$ | $\frac{1.06}{}$ |  |
| 245.4 | 0.93 | 3 n |  |

The angle has increased by $100^{\circ}$ since 1934. The period is probably less than a century. The corresponding dp is $0!057$.

| 12898 A 600 |  |  | $+43^{\circ} 3352$ |
| :---: | :---: | :---: | :---: |
| $19^{\text {h }} 44^{\text {m }} 0$ | +43 ${ }^{\circ} 22^{\prime}$ |  | 9.5-10.0 |
| 60.526 | 30:8 | 00.34 |  |
| 61.567 | 32.9 | 0.32 |  |
| 61.710 | 34.5 | 0.32 |  |
| 62.701 | 36.5 | 0.29 |  |
| 62.710 | 36.4 | 0.28 |  |
| 61.84 | 34.2 | 0.31 | 5n |
| To the slow increase in angle corresponds |  |  |  |
| 12910 J 1335 |  |  | +19 ${ }^{\circ} 4145$ |
| $19^{\mathrm{h}} 44^{\mathrm{m}} .3$ | $+19^{\circ} 16^{\prime}$ |  | 11.2-11.3 |
| $\begin{aligned} & 60.615 \\ & 60.618 \\ & 60.621 \\ & \hline \end{aligned}$ | 176:2 | 1'14 |  |
|  | 180.5 | 1.03 |  |
|  | 180.7 | 0.91 |  |
| Unchanged since | 179.1 | 1.03 | 3 n |
|  | 1914. |  |  |

Van de Kamp
$19^{\mathrm{h}} 44_{\mathrm{m}}^{\mathrm{m}} 4$ $+31^{\circ} 54^{\prime}$ 10.0-11.0 MO

| 61.560 | 133.7 | 3.43 |  |
| :--- | :--- | :--- | :--- |
| 61.567 | 131.0 | 3.60 |  |
| 61.574 | 131.6 | 3.54 |  |
| 61.577 | 132.8 | 3.52 |  |
| $\frac{61.580}{61.57}$ | $\frac{134.0}{132.6}$ | $\frac{3.59}{3.54}$ |  |
|  | 5n |  |  |

The change is nearly all in increased distance and corresponds to $\mathrm{dp}=0!!083$. The period will be long.

| 12927 A 717 |  |  | $-2^{\circ} 5116$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $19^{\mathrm{h}} 5^{\text {m }}$. 1 | $-2^{\circ} 2^{\prime}$ |  | 9.4-10.9 | F8 |
| 58.669 | 98.2 | 0.46 |  |  |
| 61.710 | 99.6 | 0.50 |  |  |
| 62.701 | 99.2 | 0.40 |  |  |
| 61.03 | 99.0 | 0.45 | 3 n |  |
| To the slow increase in angle corresponds $\mathrm{dp}=$ 0'.0046. |  |  |  |  |
| 12937 Hu 347 |  |  | $+18^{\circ} 4242$ |  |
| $19^{\mathrm{h}} 45^{\mathrm{m}} 4$ | $+19^{\circ} 9^{\prime}$ |  | 8.7-11.7 |  |
| 58.669 | $335: 4$ | 1.18 |  |  |
| 61.710 | 333.6 | 1.37 |  |  |
| 62.689 | 334.0 | 1.35 |  |  |
| 62.701 | 332.1 | 1.27 |  |  |
| 62.707 | 328.1 | 1.20 |  |  |
| 64.730 | 331.3 | 1.31 |  |  |
| 62.20 | 332.4 | 1.28 | $6 n$ |  |
| The slow chang 0 01009. | in both | coord | inates ma | S |



| 59.933 | 48.5 | 0.31 |  |
| :--- | :--- | :--- | :--- |
| 60.526 | 48.2 | 0.28 |  |
| 61.574 | 46.4 | 0.29 |  |
| $\underline{62.701}$ | $\underline{48.1}$ | $\underline{0.27}$ |  |
| 60.88 | 47.7 | 0.29 | $5 n$ |

Hardly changed in 57 years.

| 12993 Hu 349 |  |  | $+16^{\circ} 4023$ |
| :---: | :---: | :---: | :---: |
| $19^{\text {h }} 47 \mathrm{~m}$ \% 7 | $+16^{\circ} 54^{\prime}$ |  | 8.3-12.7 B9 |
| 59.640 | $234: 4$ | 2'.57 |  |
| 59.643 | 236.3 | 2.49 |  |
| 59.682 | 236.6 | 2.60 |  |
| 60.529 | 236.7 | $\underline{2.60}$ |  |
| Unchanged in 58 | $\begin{aligned} & 236.0 \\ & \text { years. } \end{aligned}$ | 2.56 | $4 n$ |
| Djurkovic |  |  | 13 Vul |
| $19^{\mathrm{h}_{51} \mathrm{~m}_{3}}$ | $+23^{\circ} 57^{\prime}$ |  | 4.6-7.8 A0 |
| 61.560 | 242.2 | 0.76 |  |
| 61.567 | 239.3 | 0.72 |  |
| 61.571 | 239.5 | 0.68 |  |
| 61.57 | 240.3 | 0.72 | 3 n |
| Unchanged so far |  |  |  |
| $13104 \sum 2597$ |  |  | $-7^{\circ} 5102$ |
| $19^{\text {h }}$ 52 ${ }^{\text {m }}$ 6 | - $6^{\circ} 52{ }^{\prime}$ |  | 6.7-7.8 F2 |
| 58.669 | $86: 3$ | 0.48 |  |
| 60.774 | 90.5 | 0.41 |  |
| 62.543 62.689 | 85.5 | 0.43 |  |
| 62.701 | 86.0 | 0.43 |  |
| 62.707 | 90.4 | 0.43 |  |
| 61.68 | 88.3 | 0.44 | 6n |

To the marked decrease in distance corresponds $\mathrm{dp}=0$ ".018.
 0':006.

13258 J 784

$$
19^{\mathrm{h}} 59 \mathrm{~m}_{0} \quad+19^{\circ} 9^{\mathrm{l}} \quad 9.7-9.9
$$

| 60.615 | 159.9 | 2.35 |  |
| :--- | :--- | :--- | :--- |
| 60.618 | $\underline{159.6}$ | $\frac{2.29}{}$ |  |
| 60.62 | 159.8 | 2.32 | 2 n |

Unchanged.
Identified in the astrographic catalogue as $+19^{\circ}$, $19 \mathrm{~h} 56^{\mathrm{m}}$, No 183 and $+20^{\circ}, 20^{\mathrm{h}} 0 \mathrm{~m}$, No 1176 which gives the corrected position.
$13262 \beta 1289$

$$
+37^{\circ} 3723
$$



| 13289 Hu 353 |  |  | $+19^{\circ} 4258$ |
| :---: | :---: | :---: | :---: |
| $20^{\mathrm{h}} 0.4$ | $+19^{\circ} 57^{\prime}$ |  | 9.3-10.9 |
| 59.640 | $354 \% 6$ | 0.47 |  |
| 61.574 | 352.7 | 0.54 |  |
| 61.710 | 353.0 | 0.48 |  |
| 60.97 | 353.4 | 0.50 | 3 n |

The longer arc makes $d p=0.0004$.

| 13304 A 1666 |  |  | $+14^{\circ} 4155$ |
| :---: | :---: | :---: | :---: |

Probably an optical system. To the increase of $l^{\prime \prime}$ in the distance since 1923 would correspond an improbably large value of $d p=0!.075$.

13401 A 867
$+72^{\circ} 933$

| $20^{\text {h }} 5 . \mathrm{m}_{3}$ | +72 ${ }^{\circ} 51^{\prime}$ |  | 8.1-13.1 G |
| :---: | :---: | :---: | :---: |
| 59.640 | $149: 8$ | 2".28 |  |
| 61.574 | 146.2 | 2.37 |  |
| 61.588 | 147.2 | $\underline{2.23}$ |  |
| 60.93 | 147.7 | 2.29 | 3 n |

Change doubtful.
13418 A $383 \quad+41^{\circ} 3617$

| $20^{\mathrm{h}_{6} \mathrm{~m}_{4}}$ | $+41^{\circ} 53^{\prime}$ | $10.1-10.1$ |  |
| :---: | :--- | :--- | :--- |
| 59.660 | 216.3 | 0.33 |  |
| 61.567 | 212.0 | 0.33 |  |
| $\frac{61.574}{60.93}$ | $\frac{213.5}{213.9}$ | $\frac{0.36}{0.34}$ | 3 n |

To the slow decrease in angle corresponds $\mathrm{dp}=$ 0':0034.
$13449 \quad \sum 2652$
$+61^{\circ} 1975$

| $20^{\text {h }} 8 \mathrm{~m} \cdot 2$ | $+61^{\circ} 56{ }^{\prime}$ |  | 7.2-7.5 A0 |
| :---: | :---: | :---: | :---: |
| 60.601 | $239: 3$ | $0!31$ |  |
| 61.580 | 238.9 | 0.33 |  |
| 61.710 | 240.6 | 0.31 |  |
| 61.754 | 236.2 | 0.32 |  |
| 61.41 | 238.7 | 0.32 | 4 n |

To the slow retrograde motion corresponds $d p=$

0:,0012.

| 13461 O§ 400 |  |  | $+43^{\circ} 3513$ |
| :---: | :---: | :---: | :---: |
| $20^{\text {h }}$ 8. ${ }^{\text {m }} 6$ | +43 ${ }^{\circ} 48^{\prime}$ |  | 7.5-8.7 G5 |
| 60.519 | $286: 2$ | 0! 24 |  |
| 60.526 | 286.1 | 0.27 |  |
| 60.615 | 285.2 | 0.27 |  |
| 62.701 | 283.4 | 0.24 |  |
| 61.09 | 285.2 | 0.26 | 4 n |

$+7: 1$ and $+0^{\prime \prime} .04$

| $20^{\mathrm{h}} 14^{\mathrm{m}} 5$ | +33 ${ }^{\circ} 56^{\prime}$ |  | 9.5-9.8 |
| :---: | :---: | :---: | :---: |
| AB |  |  |  |
| 59.613 | 26:8 | 2.90 |  |
| 59.638 | 25.0 | 2.93 |  |
| 59.643 | 27.0 | 2.71 |  |
| 60.588 | 28.1 | 2.92 |  |
| 62.707 | 25.7 | 2.72 |  |
| 64.710 | 24.5 | 2.65 |  |
| $\underline{64.730}$ | 25.3 | $\underline{2.88}$ |  |
| 61.38 | 26.1 | 2.82 | 7 n |
| AC |  |  | 9.5-9.9 |
| 59.613 | 14.4 | 10.43 |  |
| 59.638 | 14.0 | 10.36 |  |
| 60.588 | 13.7 | 10.27 |  |
| 62.707 | 13.8 | 10.40 |  |
| 62.710 | 13.8 | 10.49 |  |
| $\underline{64.730}$ | 14.2 | 10.20 |  |
| 61.66 | 14.0 | 10.36 | 6n |

Change questionable.

| 13641 A 1423 |  | $+36^{\circ} 3987$ |
| :--- | :--- | :--- | :--- | :--- |



13650 J 1340

| $20^{\mathrm{h}} 15^{\mathrm{m}} .5$ | $+17^{\circ} 48 \mathrm{l}$ | $9.4-9.4$ |  |
| :---: | :--- | :--- | :--- |
| 62.702 | 132.6 | $1: .93$ |  |
| 62.707 | 133.8 | 2.00 |  |
| 62.718 | 134.7 | 1.86 |  |
| 62.71 | 133.7 | 1.93 | $3 n$ |

Unchanged in 48 years. The star was identified in the astrographic catalogue as $+18^{\circ}, 20 \mathrm{~h} 16^{\mathrm{m}}$ No 25, prob. double" and $+17^{\circ}$ $20^{\mathrm{h}} 12 \mathrm{~m}$ No 272.
The star is erroneously called J 1440 in Aitken's G. C. and J 1370 in the Lick Index Catalogue.


| $20^{\mathrm{h}} 16^{\text {m/ }}$. | $+34^{\circ} 57^{\prime}$ |  | 9.4-9.4-11.1 |
| :---: | :---: | :---: | :---: |
| AB |  |  |  |
| 60.618 | $106 \% 4$ | 0 0.13 |  |
| $\underline{60.621}$ | 104.6 | 0.16 |  |
| 60.62 | 105.5 | 0.15 | 2n |
| The angle is decreasing. |  |  |  |
| $\mathrm{AB}-\mathrm{C}=\beta 986$ |  |  |  |
| 60.618 | 240:8 | 4!'56 | $1 n$ |
| Unchanged since 1880. |  |  |  |
| 13686 A 1425 |  |  | $+37^{\circ} 3879$ |
| $20^{\mathrm{h}} 17 \mathrm{~m}_{0}$ | $+38^{\circ} 7^{\prime}$ |  | 8.5-8.5 B1 |
| AB |  |  |  |
| 60.618 | 272:3 | 0!19 |  |
| 60.621 | 275.9 | 0.17 |  |
| 60.62 | 274.1 | 0.18 | 2n |
| Slow decrease in angle. |  |  |  |
| 13728 A 1427 |  |  | $+38^{\circ} 4021$ |
| $20^{\mathrm{h}} 18.4$ | +39 ${ }^{\circ} 15^{\prime}$ |  | 6.3-8.3 A0 |
| 60.621 | 92:6 | $0!25$ |  |
| 61.567 | 89.4 | 0.28 |  |
| 61.571 | 89.4 | 0.24 |  |
| 61.711 | 92.0 | 0.27 |  |
| 61.37 | 990.8 | 0.26 |  |
| Muller's orbit $+4: 2$ and +0.104 . | (1954) gi | ves th | residuals |
| 13744 A 725 |  |  | $+44^{\circ} 3436$ |
| $20^{\mathrm{h}} 19^{\text {m }} \cdot 3$ | $+44^{\circ} 27^{\prime}$ |  | 9.3-10.1 |
| 61.571 | 23898 | . 29 |  |
| 62.701 | 239.4 | 36 |  |
| 64.903 | $\underline{237.3}$ | 0.30 |  |
| 63.06 | (1955) gives the residuals |  |  |
| Muller's orbit $+3: 6$ and $+0!.05$. |  |  |  |  |
| 13830 - 432 |  |  | $+35^{\circ} 4102$ |
| $20^{\text {h }} 22 .{ }^{\text {m }} 9$ | $+35^{\circ} 36^{\prime}$ |  | 9.0-10.3 F8 |
| 59.635 | $200 \% 2$ |  | 42 |
| 59.638 | 199.2 | . 58 |  |
| 59.640 | 198.7 | . 54 |  |
| $\underline{59.643}$ | 200.1 | 1.60 |  |
|  |  |  |  |
|  |  |  |  |  |  |  |  |
| 13850 A 730 |  |  | $+59^{\circ} 2228$ |
| $20^{\text {h }} 24.0$ | +59 ${ }^{\circ} 26^{\prime}$ |  | 6.8-7.0 AO |
| 60.609 | 349.5 | . 22 |  |
| 60.621 | 343.2 | 0.24 |  |
| 64.905 | 337.3 | 0.20 |  |
| 64.947 | 339.1 | 0.22 |  |
| 62.77 | $(1955)^{342} \text { gives the small residuals }$ |  |  |
| Baize's orbit ( <br> $-1 \% 2$ and +0.01 . |  |  |  |  |
| 13869 A 293 |  |  | $+41^{\circ} 3752$ |
| $20^{\mathrm{h}} 25^{\mathrm{m}} .1$ | $+41^{\circ} 42^{\prime}$ |  | 9.6-9.8 |
| 59.608 | 119.6 | 1 ".61 |  |
| 59.613 | 122.5 | 1.68 |  |
| 59.627 | 121.4 | 1.54 |  |
| $\underline{59.635}$ | 120.8 | 1.74 |  |
| Distance increased. |  |  |  |
|  |  |  |  |  |  |  |  |
| 13894 A 610 |  |  | $+6^{\circ} 4543$ |
| $20^{\text {h }} 26{ }^{\text {m }}$. ${ }^{\text {a }}$ | $+6^{\circ} 59^{\prime}$ |  | 9.2-9.4 G0 |


| 60.722 | 281.8 | 0.48 |  |
| :--- | :--- | :--- | :--- |
| 60.768 | 282.2 | 0.40 |  |
| 60.777 | 283.6 | 0.42 |  |
| 61.567 | 283.0 | 0.42 |  |
| $\underline{61.571}$ | $\underline{281.0}$ | $\underline{0.39}$ |  |
| 61.08 | 282.3 | 0.42 | 5 n |

Heintz' orbit (1962) gives the residuals: +3.1 and $0!.00$.

| 13944 A 1675 |  | $+15^{\circ} 4181$ |  |
| :---: | :---: | :---: | :---: |
| $20^{\mathrm{h}} 28 . \mathrm{m} .8$ | $+15^{\circ} 38^{\prime}$ |  | 7.6-7.6 A2 |
| 59.687 | $149^{\circ} 1$ | 0..17 |  |
| 59.933 | 153.5 | 0.19 |  |
| 60.588 | 156.0 | 0.21 |  |
| 60.609 | 155.5 | 0.18 |  |
| 60.20 | 153.5 | 0.19 | $4 n$ |
| Baize's orbit (1954) gives the residuals $-13: 1$ and $-0!.01$. |  |  |  |
| 13946 Da $1=O \sum 407$ |  |  | $\div 10^{\circ} 4307$ |
| $20^{\mathrm{h}} 28 . \mathrm{m} \cdot 8$ | $\div 11^{\circ} 5^{\prime}$ |  | 7.9-8.0 A0 |
| 59.656 | $242^{\circ} \mathrm{T}$ | 0'. 24 |  |
| 59.687 | 245.4 | 0.18 |  |
| 60.609 | 243.8 | 0.22 |  |
| 59.98 | 243.8 | 0.21 | 3 n |

Long period orbital motion approaching a critical phase.

## $13964 \sum 2695$

$+25^{\circ} 4272$

| ${ }_{20} 0^{\mathrm{h}} 29^{\mathrm{m}} .8$ | $+25^{\circ} 38^{\prime}$ |  | $6.5-8.3$ |
| :---: | :---: | :---: | :---: |
|  | A. $2-\mathrm{G}$ |  |  |
| 59.566 | $86^{\circ} .3$ | 0.160 |  |
| 59.627 | 87.5 | 0.62 |  |
| 59.640 | 85.5 | 0.52 |  |
| $\frac{60.768}{59.90}$ | $\frac{84.3}{85.9}$ | $\underline{0.59}$ |  |

This pair deserves attention as it closes in. 13966 Hu 761 $+60^{\circ} 2132$

$$
\begin{array}{llll}
20^{\mathrm{h}} 29.9 & 160^{\circ} 56^{\prime} & & 9.5-9.5 \\
& & \mathrm{~F} 8 \\
\frac{60.609}{60.62} & \frac{142.0}{142.3} & 0.123 & \\
\frac{0.19}{142.2} & 0.21 & 2 \mathrm{n} &
\end{array}
$$

There may have been a change of quadrant around 1955.

## $13986 \beta 670$

| $20^{\mathrm{h}} 30^{\mathrm{m}} .6$ | $+13^{\circ} 46^{\prime}$ |  | 8.9 |
| ---: | :---: | :---: | :---: |
|  | AB |  |  |
| 62.543 | 26.1 | 0.171 |  |
| 62.689 | 22.8 | 0.62 |  |
| 62.701 | 26.7 | 0.59 |  |
| 62.704 | 23.8 | 0.71 |  |
| $\frac{62.710}{62.67}$ | $\frac{23.1}{24.5}$ | $\frac{0.61}{0.65}$ | 5 n |

8.9-9.2
2




## $59.640 \quad 183^{\circ} .4 \quad 1.23$ <br> $59.37 \quad 183.7 \quad 1.19 \quad 4 n$

To the increase in both coordinates corresponds $\mathrm{dp}=\mathrm{n}^{\prime \prime} \cdot 0034$. The $0!074$ proper motion confirms the plysical connection.

| $\underline{15426} \beta 841$ |  |  | $+53^{\circ} 2728$ |
| :---: | :---: | :---: | :---: |
| $21^{\mathrm{h}} 51 . \mathrm{m}$ | $\div 54{ }^{\circ} 2^{1}$ |  | 9.5-12.5 |
| 59.631 | $196: 8$ | 2: 40 |  |
| 61.300 | 197.5 | 2.48 |  |
| 64.905 | 196.2 | $\underline{2.45}$ |  |
| 61.95 | 196.8 | 2.44 | $3 n$ |
| Unchanged after 81 years. |  |  |  |
| 15444 Kui 62 |  |  | $+38^{\circ} 4636$ |
| $21^{\mathrm{h}_{52}} \mathrm{~m} 9$ | $+38^{\circ} 28^{\prime}$ |  | 8.9-10.0 A0 |
| 60.765 | 53.7 | $1{ }^{1 .} 76$ |  |
| 60.777 | 53.5 | 1.72 |  |
| 62.543 | 53.5 | 1.79 |  |
| 62.718 | 52.8 | 1.70 |  |
| 61.70 | 53.4 | 1.74 | $4 n$ |

No definite change in 60 years.

| $15447 \beta 75$ |  |  | $+10^{\circ} 4659$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $21^{\mathrm{h}} 53{ }^{\text {m }} 1$ | $\therefore 10^{\circ} 39^{\prime}$ |  | 8.4-8.9 | G5 |
| 59.656 | $165: 9$ | 0'.28 |  |  |
| 59.687 | 168.2 | 0.28 |  |  |
| 59.942 | 162.9 | 0.28 |  |  |
| 59.964 | 161.2 | 0.26 |  |  |
| 60.609 | 172.6 | 0.31 |  |  |
| 60.615 | 172.6 | 0.30 |  |  |
| 60.618 | 170.6 | 0.30 |  |  |
| 60.774 | 169.1. | 0.32 |  |  |
| 60.870 | 170.6 | 0.32 |  |  |
| 62.543 | 1.80 .3 | 0.34 |  |  |
| 62.701 | 184.8 | 0.35 |  |  |
| 62.704 | 180.6 | 0.35 |  |  |
| 62.704 | 135.0 | 0.32 |  |  |
| 62.940 | 180.0 | 0.34 |  |  |
| 64.730 | 188.0 | 0.34 |  |  |
| 64.736 | 187.6 | 0.38 |  |  |
| 64.740 | 191.4 | 0.33 |  |  |
| 59.81 | 164.6 | 0.28 | 4 n |  |
| 60.70 | 171.1 | 0.31 | $5 \pi$ |  |
| 62.72 | 182.1 | 0.34 | 5 n |  |
| 64.74 | 189.0 | 0.35 | 3 n |  |
| The residuals | from Heint | $z^{\prime}$ orb | bit (196 | a |
| 59.67 | $+2.1$ | +0'.02 |  |  |
| 60.72 | + 1.8 | +0.04 |  |  |
| 62.29 | + 3.6 | +0.05 |  |  |
| 64.74 | - 0.9 | +0.02 |  |  |
| 15472 Hu 382 |  |  | $+18^{\circ}$ | 892 |
| $21^{\mathrm{h}} 54{ }^{\mathrm{m}} 2$ | $+19^{\circ} 26^{\prime}$ |  | 9.9-10. |  |
| 59.656 | 193.4 | 0.139 |  |  |
| 59.942 | 196.2 | 0.40 |  |  |
| 59.964 | 189.9 | 0.38 |  |  |
| 60.722 | 196.7 | 0.38 |  |  |
| 60.744 | 190.6 | 0.34 |  |  |
| 60.870 | 195.2 | 0.36 |  |  |
| 62.701 | 193.3 | 0.34 |  |  |
| 62.959 | 192.2 | 0.38 |  |  |
| 60.94 | 193.4 | 0.37 | 8 n |  |

The angle increases without change in distance. The longer arc reduces dp to 0.003 .
$\underline{15476} \mathrm{O} 537 \quad+59^{\circ} 2435$

| $21^{\mathrm{h}_{54} \mathrm{~m}_{4}}$ | $\div 59^{\circ} 35^{\prime}$ |  | $7.9-11.0$ | $\Lambda 0$ |
| :---: | :--- | :--- | :--- | :--- |
| 59.624 | $205^{\circ} .1$ | 2.22 |  |  |
| 59.635 | 206.8 | 1.99 |  |  |
| 59.638 | 204.1 | 2.23 |  |  |
| $\frac{59.640}{59.63}$ | $\frac{204.5}{205.1}$ | $\frac{2.08}{2.13}$ | 4 n |  |

Change in distance is questionable but the
increase in angle makes $\mathrm{d} p=0^{\prime \prime} .006$. The 0.106 proper motion confirms the physical connection.

| 15477 - 1214 |  |  | $+33^{\circ} 4387$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $21^{\mathrm{h}} 54{ }^{\mathrm{m}} 4$ | $+34^{\circ} 4^{\prime}$ |  | 9.7-11.0 | A0 |
| 59.635 | $207: 4$ | 1:38 |  |  |
| 59.638 | 208.3 | 1.43 |  |  |
| 59.643 | 206.4 | 1.46 |  |  |
| 59.933 | 209.1 | 1.47 |  |  |
| 59.942 | 205.5 | 1.40 |  |  |
| 59.76 | 207.3 | 1.43 | 5 n |  |
| Unchanged in 70 years. |  |  |  |  |
| 15481 O§ 458 |  |  | $+59^{\circ} 2436$ |  |
| $21^{\mathrm{h}_{54}} \mathrm{~m}_{9} 9$ | $+59^{\circ} 32^{\prime}$ |  | 7.0-8.5 | A0 |
| 59.640 | 348.9 | 0".98 |  |  |
| 59.643 | 349.9 | 1.00 |  |  |
| 59.64 | 349.4 | 0.99 | 2 n |  |
| Slow increase in distance. |  |  |  |  |
| $15494 \times 2847$ |  |  | $-4^{\circ} 5585$ |  |
| $21^{\mathrm{h}} 55^{\mathrm{m}_{5}}$ | $-3^{\circ} 43^{\prime}$ |  | 8.4-8.8 F0 |  |
| 59.613 | $307: 6$ | 0.98 |  |  |
| 59.643 | 308.8 | 0.98 |  |  |
| 59.646 | 310.0 | 0.88 |  |  |
| 62.929 | 310.1 | 0.96 |  |  |
| 62.949 | 307.9 | 1.09 |  |  |
| 62.959 | 308.8 | 1.10 |  |  |
| 64.740 | 309.6 | 1.07 |  |  |
| 61.78 | 309.0 | 1.01 |  |  |
| To the very slow change corresponds $\mathrm{d} p=$ 0'. 003 . |  |  |  |  |
| 15505 A 1898 |  |  | $\div 55^{\circ} 2658$ |  |
| $21^{\mathrm{h}} 56{ }^{\text {m/6 }}$ | $+56^{\circ} 1^{\prime}$ |  | 9.2-9.8 A2 |  |
| 59.613 | 229.7 | 1:34 |  |  |
| 59.624 | 231.4 | 1.36 |  |  |
| 59.635 | 230.9 | 1.41 |  |  |
| 59.640 | 232.8 | 1.46 |  |  |
| 59.63 | 231.2 | 1.39 | $4 n$ |  |
| Unchanged in 51 years. |  |  |  |  |
| 15691 O§ 463 |  |  | $+13^{\circ} 4860$ |  |
| $22^{\mathrm{h}} \mathrm{m}_{.9}$ | $+13^{\circ} 30^{\prime}$ |  | 8.2-12.1 G0 |  |
| 61.551 | 359:8 | 4:13 |  |  |
| 61.571 | 359.4 | 3.89 |  |  |
| 61.577 | 360.5 | 3.95 |  |  |
| 61.580 | 358.9 | 4.09 |  |  |
| 61.661 | 359.1 | 3.92 |  |  |
| 61.59 | 359.5 | 4.00 | 5 n |  |

The 0'15 proper motion proves the physical connection; to the slow change in both coordinates corresponds $\mathrm{dp}=0.025$.
1.5707 O 464
$22^{\mathrm{h}} 9^{\mathrm{m}} 1$

| 59.656 | 107.7 | 0.129 |  |
| :--- | :--- | :--- | :--- |
| 59.964 | 108.5 | 0.26 |  |
| 60.595 | 110.8 | 0.29 |  |
| 60.618 | 109.8 | 0.32 |  |
| 60.701 | 111.2 | 0.32 |  |
| 62.950 | 108.7 | 0.29 |  |
| $\frac{62.959}{61.06}$ | $\underline{109.4}$ | $\underline{0.29}$ |  |
|  | 109.4 | 0.29 | $7 n$ |

This pair is closing in and should be watched in the coming years. To the large change in both coordinates corresponds $\mathrm{dp}=0!.008$.

= 0."010.

| 15738 | но 179 | $+29^{\circ} 4607$ |
| :---: | ---: | ---: |
| $22^{\mathrm{h}} 10^{\mathrm{m}_{4}}$ | $+29^{\circ} 58^{\prime}$ | $8.7-9.7$ |


| 60.768 | 272.4 | 0.170 |  |
| :--- | :--- | :--- | :--- |
| 62.543 | 275.1 | 0.74 |  |
| 62.701 | 273.8 | 0.64 |  |
| 62.704 | 275.4 | 0.78 |  |
| $\frac{62.718}{62.29}$ | $\frac{275.2}{274.4}$ | $\frac{0.84}{0.74}$ |  |
|  | 5n |  |  |

Increase in both coordinates making $d p=0 \% 006$.
$15835 \mathrm{Hu} 383 \quad+20^{\circ} 5127$


No definite change in 59 years.

| 15838 Es 1020 |  |  | $+52^{\circ} 31$ |
| :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 17^{\mathrm{m}} / 4$ | $4.52^{\circ} 54^{\prime}$ |  | 10.2-10.4 |
| 59.613 | $259: 8$ | 1'.85 |  |
| 59.624 | 260.4 | 2.13 |  |
| 59.635 | 259.8 | 1.96 |  |
| 59.950 | 259.1 | 1.91 |  |
| 60.588 | 262.4 | $\underline{2.01}$ |  |
| 59.88 | 260.3 | 1.97 | 5 n |

To the increasein angle corresponds $\mathrm{dp}=0!.015$.

| 15858 ק1217 |  |  | $+30^{\circ} 4685$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 18^{\mathrm{m}} 7$ | $\therefore 31^{\circ} 2^{\prime}$ |  | 7.7-10.6 | K0 |
| 60.774 | $221: 3$ | $0!.43$ |  |  |
| 62.701 | 219.1 | 0.42 |  |  |
| 62.954 | 223.5 | 0.44 |  |  |
| 62.14 | 221.3 | 0.43 | $3 n$ |  |

No definite change so far.

| 15859 A 628 |  |  | $+10^{\circ} 4735$ |  |
| ---: | :--- | :--- | :--- | :--- |
| $22^{\mathrm{h}} 18 \mathrm{~m}_{7}$ | $+10^{\circ} 39^{\prime}$ |  | $8.9-11.9 \quad \mathrm{KO}$ |  |
| 59.624 | 227.7 | $1!20$ |  |  |
| 59.635 | 226.9 | 1.25 |  |  |
| $\frac{60.774}{60.01}$ | $\frac{229.8}{228.1}$ | $\frac{1.29}{1.25}$ | $3 n$ |  |

No definite change but the 0.05 proper motion proves the physical connection.
15861 Kr 58
$+59^{\circ} 2508$
$22^{\mathrm{h}} 18{ }^{\mathrm{m}} \cdot 9$
$+59^{\circ} 37$

| 59.613 | $26{ }^{\circ}$ |
| :--- | :--- |
| 59.624 | 27. |
| 59.635 | 28. |
| 59.950 | 28. |
| 59.71 | 27. |

Change questionable.
15889 Ho 292
$+4^{\circ} 4849$

| 61.580 | 66.5 | $4!39$ |  |
| :--- | :--- | :--- | :--- |
| $\frac{61.661}{61.60}$ | $\frac{68.0}{66.3}$ | $\frac{4.59}{4.50}$ |  |
| $3 n$ |  |  |  |

No definite change in 74 years but the 0.006 proper motion establishes the physical connection.

| 15915 Ho 183 |  |  | $+21^{\circ} 4747$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 22.4$ | $+22^{\circ} 19^{\prime}$ |  | 9.2-12.6 | MO |
| 59.624 | 203:4 | 1 1.34 |  |  |
| 59.628 | 201.0 | 1.42 |  |  |
| 59.933 | 203.2 | 1.31 |  |  |
| 59.939 | 2.02 .2 | 1.36 |  |  |
| 60.618 | 204.1 | 1.41 |  |  |
| 59.95 | 202.8 | 1.37 | 5n |  |

The slow change in both coordinates makes dp $=0.1008$. Physical connection is confirmed by the $0!22$ proper motion.

15921 Es 536

$$
\begin{array}{rlrl}
22^{\mathrm{h}} 22^{\mathrm{m} .8} & & +27^{\circ} 9 & \\
& & 10.2-10.2 \\
59.624 & 270.8 & 3.32 & \\
59.635 & 272.3 & 3.40 & \\
59.939 & 270.1 & 3.48 & \\
\frac{60.588}{59.95} & \frac{270.1}{270.8} & \frac{3.48}{3.42} & 4 \mathrm{n}
\end{array}
$$

Change doubtful after 52 years. The star was identified in the astrographic catalogue as $+27^{\circ} 66759$, $+26^{\circ} 76093$ and $+26^{\circ} 76543$ which gives the corrected position.

| 15939 Ho 185 |  |  | $+37^{\circ} 4573$ |
| :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 24^{\mathrm{m}} 3$ | +38 ${ }^{\circ} 22^{\prime}$ |  | 9.6-12.1 |
| 61.571 | 205:0 | 3.23 |  |
| 61.580 | 204.6 | 3.26 |  |
| 61.588 | 206.3 | 3.05 |  |
| 61.664 | $\underline{206.2}$ | 3.12 |  |
| 61.60 | 205.5 | 3.16 | 4 n |

An optical system. The $0!.035$ proper motion of the bright component accounts for the change.
$\underline{15956} \beta 291$
$+3^{\circ} 4709$
$22^{\mathrm{h}} 25^{\mathrm{m}} .2$
$+4^{\circ} 16^{\prime}$
9.8-9.8 F8

| 60.774 | 205.5 | $0!.28$ |  |
| :--- | :--- | :--- | :--- |
| $\frac{60.870}{60.82}$ | $\frac{203.1}{204.3}$ | $\underline{0.30}$ |  |
|  | 0.29 | $2 n$ |  |

To the uniform angular increase corresponds a $\mathrm{dp}=0.10026$.

| 15962 | $\beta 701$ |  |  | $+11^{\circ} 4804$ |
| ---: | :--- | :--- | :--- | :--- |
| $22^{\mathrm{h}_{25} 5^{\mathrm{m}} .6}$ | $+12^{\circ} 0^{\prime}$ |  | $7.3-10.3 \mathrm{KO}$ |  |
| 59.624 | $226^{\circ} .6$ | $1^{\prime \prime} 18$ |  |  |
| 59.635 | 226.0 | 1.06 |  |  |
| $\frac{59.638}{59.63}$ | $\frac{228.2}{226.9}$ | $\frac{1.01}{1.08}$ | 3 n |  |

The distance has hardly changed but the decrease in angle makes $\mathrm{dp}=0$. 012 .
15963
$+8^{\circ} 4873$
$22^{\mathrm{h}} 25^{\mathrm{m}} .7+8^{\circ} 55^{\prime}$
$A B=$

| 59.624 | 249.4 | $1!10$ |  |
| :--- | :--- | :--- | :--- |
| 59.633 | 251.0 | 1.01 |  |
| 59.935 | 247.6 | 1.12 |  |
| 60.609 | 254.2 | 1.05 |  |
| 60.870 | 252.4 | 0.91 |  |
| 64.740 | 251.7 | 1.10 |  |
| 59.73 | 249.3 | 1.08 | $3 n$ |
| 62.07 | 252.8 | 1.02 | $3 n$ |

Change questionable so far.

| 15971 之2909 |  |  | $\zeta$ Aqr |  |
| :---: | :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 26^{\mathrm{m}} 2$ | $-0^{\circ} 17^{\prime}$ |  | 4.4-4.6 | F2 |
| 59.613 | $265: 3$ | 1'.96 |  |  |
| 59.624 | 264.2 | 1.76 |  |  |
| 59.635 | 263.2 | 1.93 |  |  |
| 59.942 | 260.3 | 1.95 |  |  |
| 59.950 | 264.1 | 1.84 |  |  |
| 62.543 | 258.6 | 1.95 |  |  |
| 62.704 | 261.2 | 1.90 |  |  |
| 62.718 | 258.5 | 1.84 |  |  |
| 62.929 | 261.2 | 1.90 |  |  |
| 64.730 | 256.4 | 2.04 |  |  |
| 64.736 | 254.0 | 2.01 |  |  |
| 64.740 | 255.4 | 1.93 |  |  |
| 59.75 | 263.4 | 1.89 | 5n |  |
| 62.72 | 259.9 | 1.90 | 4 n |  |
| 64.74 | 255.3 | 1.99 | 3 n |  |

Comparison with Rabe's orbit (1954) and the one by Franz (1958):

|  | Rabe |  | Franz |  |
| :--- | :--- | :--- | :--- | :--- |
| 59.75 | +6.5 | $-0!.07$ | +2.3 | +0.02 |
| 62.71 | +2.0 | -0.04 | +5.8 | +0.06 |
| 64.73 | +1.9 | -0.06 | +5.6 | +0.17 |

seems to favor the earlier orbit.
15980 J 856

$$
\begin{array}{clll}
22^{\mathrm{h}} 26^{\mathrm{m} .7} & +29^{\circ} 4^{\prime} & & 9.0-10.4 \\
59.939 & 218.0 & 1: 60 & \\
59.942 & 216.0 & 1.43 & \\
\frac{59.950}{59.94} & \frac{218.6}{217.5} & \frac{1.36}{1.46} & 3 \mathrm{n}
\end{array}
$$

The $\Delta \mathrm{m}$ is much greater than 0.2 given in the
Lick Index Catalogue. I estimated it as $1 \mathrm{~m}, 1 \mathrm{~m} .5$ and 1 m 6 on the three nights. Questionable change.


Knipe's orbit (1959) gives the residuals +1.3 and $\div 0.05$.

15992 Hu $388 \quad+21^{\circ} 4770$

and 0.00.

| 16011 Hu 981 |  | $+60^{\circ} 2403$ |  |
| :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 28^{\mathrm{m}}$. 8 | $+61^{\circ} 22^{\prime}$ | 7.6-7.8 | A0 |
| 62.704 | 225:1 | 0.335 |  |
| 62.718 | 225.1 | 0.40 |  |
| 62.620 | 226.1 | 0.35 |  |
| 64.740 | $\underline{224.2}$ | 0.36 |  |
| 63.20 | 225.1 | 0.36 4n |  |
| The longer arc makes $\mathrm{dp}=0^{\prime \prime} .0036$. |  |  |  |
| $\underline{16057}$ ₹ 2924 |  | $+69^{\circ} 1262$ |  |
| $22^{\mathrm{h}} 31^{\mathrm{m}} 6$ | $+69^{\circ} 39^{\prime}$ | 6.5-7.0 | F2-A5 |
| 59.656 | $73: 8$ | 0.150 |  |
| 59.687 | 72.0 | 0.46 |  |
| 59.942 | 75.7 | 0.43 |  |
| 59.950 | 75.2 | 0.47 |  |
| 59.964 | 72.4 | 0.45 |  |



No definite change in distance but increase in angle making $\mathrm{dp}=0.010$.

16142 Hu 393
$+19^{\circ} 4976$
$22^{\mathrm{h}} 36^{\mathrm{m}} \mathrm{m}_{7}$

| 59.624 | 231.1 | 0.42 |  |
| :--- | :--- | :--- | :--- |
| 59.638 | 231.8 | 0.38 |  |
| 59.939 | 230.7 | 0.40 |  |
| 60.774 | 233.8 | 0.40 |  |
| $\frac{61.820}{60.36}$ | $\underline{232.6}$ | $\underline{0.46}$ |  |
| 0.0 | 0.41 | $5 n$ |  |

Increase in angle without change in distance
The longer arc reduces dp to 0.003 .
16164 Но 188
$+36^{\circ} 4905$
$22^{\mathrm{h}} 37^{\mathrm{m}} \cdot 9$
$+37^{\circ} 16^{\prime}$
8.7-8.7 F8

| 59.964 | 160.4 | 0.115 |  |
| :--- | :--- | :--- | :--- |
| 60.595 | 162.9 | 0.19 |  |
| $\underline{60.609}$ | $\underline{165.4}$ | $\underline{0.18}$ |  |
| 60.39 | 162.9 | 0.17 | $3 n$ |

The angle has increased by $120^{\circ}$ since 1885 and the distance has been halved. The corresponding dp is 0'.006.

| 16173 Но 296 |  |  | $+43^{\circ} 4971$ |
| :---: | :---: | :---: | :---: |
| $22^{\text {h }} 38.4$ | $\because 14^{\circ} 17^{\prime}$ |  | 6.6-6.6 G5 |
| 60.774 | $1: 8$ | 0.15 |  |
| 64.740 | 105.0 | 0.19 |  |
| 64.905 | 106.8 | 0.23 |  |
| 64.947 | 107.3 | 0.24 |  |
| 60.77 | 1.8 | 0.15 | 1 n |
| 64.86 | 106.4 | 0.22 | 3 n |
| Baize's orbit | (1956) mak | s the | residuals |


| $\begin{aligned} & 60.77 \\ & 64.86 \end{aligned}$ | $\begin{aligned} & +3 \circ 1 \\ & +12.8 \end{aligned}$ | $\begin{array}{r} -0.01 \\ 0.00 \end{array}$ |  |
| :---: | :---: | :---: | :---: |
| $\underline{16185} \Sigma 2934$ |  |  | $+20^{\circ} 5208$ |
| $22^{\text {h }} 39{ }^{\text {m }}$ 4 | $+21^{\circ} 10^{\prime}$ |  | 8.7-9.7 G0 |
| 60.777 | $95: 7$ | 1 1.00 |  |
| 62.695 | 94.2 | 0.98 |  |
| 62.704 | 95.2 | 0.94 |  |
| 62.710 | 93.1 | 1.02 |  |
| $\underline{62.718}$ | 94.3 | 0.92 |  |
| 62.32 | 94.5 | 0.97 | 5 n |

Heintz' orbit (1960) leaves the residuals +3.7 and +0.07 .

16186 Hu 781
$+14^{\circ} 4851$
$22^{\mathrm{h}_{39} \mathrm{~m}_{5}}$
$+14^{\circ} 58^{\prime}$
8.9-9.1 F8

| 61.588 | 340.7 | 0.40 |
| :--- | :--- | :--- |
| 61.664 | 338.1 | 0.40 |
| 61.710 | 337.7 | 0.44 |
| $\underline{61.771}$ | $\underline{337.8}$ | $\underline{0.39}$ |
| 61.68 | 338.6 | 0.41 |

The change in both coordinates leads to $\mathrm{dp}=$ 0 0.0044.

| 16204 A 414 |  |  | $+43^{\circ} 4276$ |
| :---: | :---: | :---: | :---: |
| $22^{\text {h }} 40{ }^{\text {m }}$ 4 | +43 ${ }^{\circ} 45^{\prime}$ |  | 9.6-9.7 60 |
| 59.613 | $16: 8$ | 1 1'96 |  |
| 59.624 | 16.2 | 1.80 |  |
| 59.635 | 16.5 | 1.97 |  |
| 59.950 | 15.1 | 1.89 |  |
| 59.71 | 16.2 | 1.90 | $4 n$ |
| Unchanged in 57 years. |  |  |  |
| 16209 Hu 394 |  |  | +5 ${ }^{\circ} 5060$ |
| $22^{\mathrm{h}} 40^{\mathrm{m}} .5$ | $+6^{\circ} 21^{\prime}$ |  | 9.3-11.0 G5 |
| 59.624 | $95^{\circ} 1$ | 0.155 |  |
| 59.638 60.774 | 91.1 | 0.60 <br> 0.48 <br> 0.54 |  |
| 60.01 | 93.4 | 0.54 | 3 n |

The longer arc reduces $d p$ to $0!\cdot 008$.

| 16237 A 2295 |  | $+1{ }^{\circ} 4644$ |  |
| :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 42.6$ | +. $2^{\circ} 20^{\prime}$ |  | 10.1-12.8 K2 |
| 60.774 | 86.4 | 11.21 |  |
| 60.870 | 84.7 | 1.34 |  |
| 61.588 | 82.5 | 1.29 |  |
| 61.710 | 82.5 | 1.40 |  |
| 61.921 | 83.7 | 1.31 |  |
| 61.37 | 84.0 | 1.31 | 5 n | The slow increase in angle makes $d p=0!004$.


| 16256 Hu 784 |  |  | $+51^{\circ} 3462$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 3^{\mathrm{m}} .7$ | $+52^{\circ} 16^{\prime}$ |  | 9.3-12.2 | B9 |
| 59.624 | $273: 7$ | 2".04 |  |  |
| 59.635 | 275.4 | 2.17 |  |  |
| $\underline{59.950}$ | $\underline{275.4}$ | $\underline{2.30}$ |  |  |
| No definite change in 55 years. |  |  |  |  |
|  |  |  |  |  |
| 16314 Но 482 |  |  | $\pm 25^{\circ}$ |  |


| 16314 но 482 |  |  | $+25^{\circ} 4828$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}_{4} \mathrm{~m}^{\mathrm{m}} .0}$ | $+26^{\circ} 8^{\prime}$ |  | $7.5-7.5$ | A3 |
| 60.722 | 74.4 | $0 . .20$ |  |  |
| 60.768 | 76.8 | 0.22 |  |  |
| 60.774 | $\frac{76.0}{75.7}$ | $\underline{0.23}$ |  |  |
| 60.75 | 75.7 | 0.22 | $3 n$ |  |

Residuals from Couteau's orbit (1963) are +3.5 and 0".00.
16326 A 632
$+56^{\circ} 2890$
$22^{h_{50}}{ }^{m} 0$
$+57^{\circ} 27^{\prime}$
8.2-9.0 K0

| 60.777 | 184.7 | 1.05 |
| :--- | :--- | :--- |
| 62.695 | 180.1 | 1.02 |
| 62.701 | 184.2 | 1.11 |
| 62.710 | 182.8 | 1.10 |
| 62.718 | 185.0 | 0.99 |
| $\underline{64.947}$ | $\underline{181.1}$ | $\underline{0.99}$ |
| 62.76 | 183.0 | 1.04 |

6n
Comparison with Heintz' orbit (1962) gives the residuals +3.1 and +0.102 .

| 345 - 382 |  |  | +430\%331 |
| :---: | :---: | :---: | :---: |
| $22^{\text {h }} 51{ }^{\text {m }}$ / 4 | $+44^{\circ} 29$ |  | 5.8-7.8 A0 |
| 59.624 | $180: 5$ | 0.67 |  |
| 59.635 | 186.3 | 0.67 |  |
| 59.942 | 181.1 | 0.58 |  |
| 59.950 | 183.8 | 0.64 |  |
| 62.695 | 188.8 | 0.67 |  |
| 62.701 | 187.2 | 0.66 |  |
| 62.710 | 186.3 | 0.69 |  |
| 62.718 | $\underline{185.8}$ | 0.72 |  |
| 59.79 | 182.9 | 0.64 | 4 n |
| 62.71 | 187.0 | 0.68 | $4 n$ |

Comparison with Muller's orbit (1954) shows the residuals
$59.79+3.8-0.02$
$62.71+2.6-0.04$
16367 B848 $+57^{\circ} 2639$
$22^{\mathrm{h}_{5} 2^{\mathrm{m}} .8}+58^{\circ} 6^{\circ} \quad 9.0-13.4 \quad$ A0

| 61.588 |  | 3.4 | 2.70 |
| :--- | :--- | :--- | :--- |
| $\frac{62.701}{62.14}$ | $\frac{6.4}{4.9}$ | $\frac{2.47}{2.59}$ |  |
|  | $2 n$ |  |  |

No definite change in 81 years.

| 16373 Hu 987 | $+15^{\circ} 472$ |  |  |
| :---: | :---: | :---: | :---: |
| $22^{\mathrm{h}} 53{ }^{\text {m }} 2$ | $+15^{\circ} 31^{\prime}$ |  | 9.1 |
| 59.635 | $118: 6$ | 0.154 |  |
| 59.933 | 118.0 | 0.54 |  |
| 59.939 | 117.8 | 0.51 |  |
| 59.84 | 118.1 | 0.53 | $3 n$ |
| Residuals from | Baize's | rbit | 1958) |

16417 O§ 536
$+8^{\circ} 4973$
$22^{h_{56}}{ }^{\text {m }} 0$

$$
+9^{\circ} 6^{\prime}
$$

$$
7.0-7.5 \quad 60
$$

$62.695 \quad 164: 2 \quad 0.3$
$62.710 \quad 169.5 \quad 0.25$
$\begin{array}{llll}62.718 & \frac{168.1}{62.71} & \frac{0.25}{167.3} & \\ 0.27\end{array}$
Residuals from my orbit (1935) are
$+1: 3$ and +0.04 .
16435 Hn 56
$+41^{\circ} 4656$
$22^{\mathrm{h}_{5} 7 \mathrm{~m}_{4}}$
$+41^{\circ} 33^{\prime}$
9.3-9.4
$61.577 \quad 100.8 \quad 1.16$
$61.580 \quad 101.4 \quad 1.04$
$61.588 \quad 104.2 \quad 1.20$
$61.664 \quad 104.0 \quad 1.06$
$\frac{61.842}{61.65} \quad \frac{103.9}{1029} \quad \frac{1.17}{1.13}$
The longer arc reduces dp to 0,004 .
16517 J 212

| $23^{\mathrm{h}} 4^{\mathrm{m}} 9$ | $+19^{\circ} 53^{\prime}$ |  | $9.6-9.6$ |
| :---: | :--- | :--- | :--- |
| 61.551 | 340.8 | 4.56 |  |
| 61.695 | 343.4 | 4.51 |  |
| 64.730 | 342.2 | 4.50 |  |
| 64.947 | 342.6 | 4.45 |  |
| 63.23 | 342.2 | 4.50 | 4 n |

No evidence of change.


Muller's orbit (1955) will require improvement since the residuals are +29.2 and $0!00$.

16582 J 623

| $23^{\mathrm{h}} 9^{\mathrm{m}} .4$ | $+36^{\circ} 3^{\prime}$ | $9.6-9.8$ |  |
| :---: | :---: | :---: | :---: |
| 61.588 | 237.7 | 2.17 |  |
| 61.664 | 240.5 | 2.24 |  |
| $\frac{61.716}{61.66}$ | $\underline{236.6}$ | $\underline{238.3}$ | $\underline{2.23}$ |
| 2 n |  |  |  |

No evidence of change.
The star was identified in the Hyberdad zone as $+36^{\circ} 62666$.


Baize's orbit (1960) makes the residuals $-0: 6$ and +0 '"06


16655 Kr 64
$23^{\mathrm{h}} 15^{\mathrm{m}} \cdot 4$

| 61.580 | 241.7 | 2.10 |  |
| :--- | :--- | :--- | :--- |
| 61.588 | 241.6 | 2.05 |  |
| 61.664 | 242.2 | 1.95 |  |
| 61.842 | 239.5 | 2.00 |  |
| $\frac{61.864}{61.71}$ | $\frac{241.2}{241.2}$ | $\frac{2.06}{2.03}$ |  |
| $5 n$ |  |  |  |

Change doubtful.
$16731 O \sum 495$
$23^{h_{2}} 1^{\mathrm{m}} .8$
59.687
59.939 59.939
59.81
$+56^{\circ} 2999$
owly closing in

| 16777 - 1222 |  |  | $+2^{\circ} 4669$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $23^{\mathrm{h}} 26^{\mathrm{m}} 0$ | $+3^{\circ} 17^{\prime}$ |  | 10.0-10.1 | G5 |
| 59.942 | $45: 5$ | 1:23 |  |  |
| 60.615 | 44.4 | 1.39 |  |  |
| 60.870 | 42.0 | 1.23 |  |  |
| 60.48 | 44.0 | 1.28 | $3 n$ |  |

Little change in 70 years.
Wirtanen

The angle has increased by $40^{\circ}$ since 1905 making $\mathrm{dp}=0.005$.
Muller 4
$+45^{\circ} 4301$

| $23^{\mathrm{h}_{3}} 38^{\mathrm{m}} \mathrm{m}_{7}$ | $+45^{\circ} 56^{\prime}$ |  | $7.8-8.1$ | F 5 |
| :---: | :--- | :--- | :--- | :--- |
| 59.656 | 248.7 | 0.18 |  |  |
| 59.687 | 244.8 | 0.18 |  |  |
| 59.937 | 249.9 | 0.21 |  |  |
| 59.950 | 250.0 | 0.20 |  |  |
| 60.595 | 256.5 | 0.22 |  |  |
| 60.601 | 259.8 | 0.22 |  |  |
| 60.615 | 260.9 | 0.24 |  |  |
| 64.740 | 290.8 | 0.13 |  |  |
| 65.040 | $\underline{291.5}$ | 0.14 |  |  |
| 59.81 | 248.3 | 0.19 | $4 n$ |  |
| 60.60 | 259.1 | 0.23 | $3 n$ |  |
| 64.89 | 291.2 | 0.14 | $2 n$ |  |

Evidently a short period binary.

| 16937 O§ 503 |  |  | $+19^{\circ} 5138$ |  |
| :---: | :---: | :---: | :---: | :---: |
| $23^{\mathrm{h}} 39^{\text {m }} 5$ | $+20^{\circ} 1^{\prime}$ |  | 8.1-8.7 | F8 |
| 60.722 | $133: 3$ | 1.47 |  |  |
| 60.765 | 134.1 | 1.39 |  |  |
| 60.768 | 134.0 | 1.51 |  |  |
| 60.774 | 131.5 | 1.49 |  |  |
| 60.76 | 133.2 | 1.46 | $4 \pi$ |  |

No change in angle in 112 years but the distance slowly diminishes indicating a $\mathrm{dp}=0$ ".006.

| 16954 Hu 796 |  | $+79^{\circ} 792$ |  |
| :---: | :---: | :---: | :---: |
| $23^{\mathrm{h}} 40^{\mathrm{m}} .8$ | $+80^{\circ} 15^{\prime}$ |  | 10.3-10.6 |
| 59.624 | $308: 2$ | 0.47 |  |
| 59.646 | 309.2 | 0.48 |  |
| 59.656 | 305.7 | 0.59 |  |
| 60.601 | 307.3 | 0.59 |  |
| 59.88 | 307.6 | 0.56 | $4 n$ |
| Unchanged in | years. |  |  |


| 16957 AGC 14 |  | 78 Peg |  |  |
| :--- | :--- | :--- | :--- | :--- |
| $23^{\mathrm{h}} \mathrm{HI}^{\mathrm{m} / 4}$ | $+29^{\circ} 6^{\prime}$ |  | $5.0-8.2$ | KO |
| 60.722 | $231^{\circ} .5$ | $0^{\prime \prime} .83$ |  |  |
| 60.768 | 232.8 | 0.99 |  |  |
| $\frac{60.870}{60.79}$ | $\frac{232.2}{232.2}$ | $\frac{0.82}{0.88}$ | $3 n$ |  | Slow orbital motion. The longer arc makes dp $=0!.0135$.



| 17036 A 792 |  |  | $+46^{\circ} 4184$ |
| :---: | :---: | :---: | :---: |
| $23^{\mathrm{h}} 48^{\mathrm{m}_{0}}$ | $+46^{\circ} 46^{\prime}$ |  | 9.3-9.5 |
| 60.601 | $263: 4$ | 0.'52 |  |
| 62.695 | 261.5 | 0.64 |  |
| 62.701 | 259.2 | 0.57 |  |
| 62.704 | 258.0 | 0.61 |  |
| 62.710 | 262.5 | 0.58 |  |
| 62.28 | 260.9 | 0.58 | $5 n$ |
| Slow change $d p=0!.006$. | both coo | ina | s indicating |

17049 Hu 699
$+50^{\circ} 4171$
$23^{\mathrm{h}} 49^{\mathrm{m}} .0$
$59.624 \quad 109: 8 \quad 1: 11$
$59.646 \quad 108.4 \quad 1.04$
$59.64 \quad 109.1 \quad 1.08 \quad 2 n$
Unchanged in 57 years.
$17062 \beta 996$

$$
23^{\mathrm{h}} 49 . \mathrm{m}_{6}
$$

$+75^{\circ} 16^{\prime}$
6.6-11.5 K2
$62.959 \quad 99.1 \quad 4 \% 45$
$64.730 \quad 101.7 \quad 4.54$
$63.84 \quad 100.4 \quad 4.50 \quad 2 n$
The longer arc makes $d p=0!067$.
17122 A 799

$$
23^{h} 55^{\mathrm{m} .0} \quad+47^{\circ} 48^{\prime}
$$

| 59.613 | 14.9 | 2.100 |  |
| :--- | ---: | ---: | ---: |
| 59.624 | 12.5 | 1.80 |  |
| 59.638 | 13.5 | 1.88 |  |
| 59.646 | 14.6 | 1.99 |  |
| 59.950 | $\frac{10.4}{59.69}$ |  | $\frac{1.91}{13.2}$ |

No change in angle but the increase in distance makes $\mathrm{dp}=0.010$.
$17149 \sum 3050$

$$
+32^{\circ} 4747
$$

$$
23^{h_{56}}{ }^{m} \cdot 9
$$

$$
+33^{\circ} 27^{\prime}
$$

6.6-6.6 F8

| 62.695 | 283.7 | 11.50 |  |
| :--- | :--- | :--- | :--- |
| 62.701 | 281.9 | 1.35 |  |
| 62.710 | 281.4 | 1.32 |  |
| 62.718 | 283.1 | 1.42 |  |
| $\frac{62.942}{62.75}$ | $\frac{281.2}{282.3}$ | $\frac{1.41}{1.40}$ |  |

Franz's orbit (1953) gives the small residuals +0.7 and +0.02 .
17178 Hn 60

$$
\div 38^{\circ} 5112
$$



| 60.777 | 196.5 | $0!.91$ |  |
| :--- | :--- | :--- | :--- |
| 61.580 | 198.2 | 0.90 |  |
| 61.588 | 196.5 | 0.90 |  |
| $\frac{61.664}{61.40}$ | $\frac{197.2}{197.1}$ | $\frac{0.94}{0.91}$ |  |
| 4 n |  |  |  |

Comparison with Heintz's orbit (1963) gives the residuals $+3: 4$ and +0.106 .

ADDENDUM
$11334 \sum 2315+27^{\circ} 3016$
$18^{h_{2}} 3^{m_{0}} 0 \quad+27^{\circ} 22^{\prime} \quad 6.6-7.6 ~ A 0$
$61.426 \quad 138.8 \quad 0.57$
$61.497 \quad 137.4 \quad 0.59$
$61.46 \quad 138.1 \quad 0.58 \quad 2 n$
Residuals from Heintz' orbit (1959) are $+1: 6$ and $0!00$.


[^0]:    3799 O§ 517
    $+1^{\circ} 938$

[^1]:    Change questionable.

