Craters and Resurfacing of the Martian North Polar Cap K. Galla¹, S. Byrne¹, B. Murray², A. M^CEwen¹ and The HiRISE team¹ ¹Lunar and Planetary Lab, University of Arizona. ²GPS Division, Caltech. Correspondence: shane@lpl.arizona.edu

30 Second Summary...

- Mass Balance of the north polar cap is uncertain.
- We need to connect current behavior to current climate to infer past climate from the polar layered deposits.
- Searched for craters in the CTX dataset to constrain resurfacing timescales, ~70 craters found.
- Followed up with HiRISE on ~40 of these.
- Craters are sites of preferential ice accumulation.
- Ablation and eolian erosion contribute to crater removal.
- **Population statistics consistent with** an equilibrium population crater lifetime proportional to D^{0.88}

CTX crater search

- PDS products projected in ISIS and searched using ArcMAP
- >100 craters found on the residual cap and NPLD combined (only 4 previously known)
- 70 craters on residual cap observed with CTX
- 37 followed up with HiRISE (several in stereo)

v = 1.0081x - 16.379

- Craters easy to see in CTX due to wind streaks
- •HiRISE comparison shows that our CTX measurements overestimate diameters.



Crater Removal Some craters are VERY fresh d/D ~ 0.25





PSP_009689_2645

- **PSP 009862 2645**
- Most craters are being infilled
- Accumulation here in preference to surroundings
- Shadowing kickstarts accumulation
- Starts positive feedback, fresh ice (smallgrained) is bright - stays cooler.





PSP 009792 2790

PSP 009223 2640

- Ablation pits (sun cups) are common and degrade the rims.
- Eolian action also erodes the rims.
- Some craters are barely recognizable.



PSP_009773_2675 PSP_009770_2655 Full range of morphologies allows "space-

for-time" substitution

Population Statistics

- Size-frequency plots can either give:
- An age for a 'production surface'
- A resurfacing rate for an 'equilibrium surface' Surface is very young so all these craters are probably primary impacts

| 100 | |
|----------|------------------------------|
| | Uncorrected D Corrected D |
| ה | |
| | |
| | |
| 1 | 100 |

•We can correct the CTX-only diameters.

Slopes of -2.1 or -2.2 for uncorrected or corrected diameters.

Diameter (m) Production function (from Hartmann) expected to have slopes of -3 here (Primaries + Secondaries).

- So either:
- Primary production has a lower size-freq slope than primary+secondary production.
- This is an equilibrium population and many craters have been erased.
- We favor the latter based on the range of morphologies
- An equilibrium surface with crater lifetime proportional to $D^{0.8}$ to $D^{0.9}$.

Modeling Polar Cap Resurfacing Resurfacing of cap not directly tied to crater removal rate as craters are preferred sites for new deposition. • We're starting landscape evolution modeling of accumulation, ablation and eolian redistribution of ice to modify craters. Craters make perfect control features and constrain rates of processes. Current population accumulated in $\sim 10^4$ years. **Reversed Perihel** Montme et al. 20(2 250 North Pole Time (in 1000 Years)