

No. 79 LABORATORY SPECTRA FOR TESTING THE PRESENCE OF  
MINOR CONSTITUENTS IN PLANETARY ATMOSPHERES, II:  
 $C_2H_2$ ,  $C_2H_4$ ,  $C_2H_6$ ,  $CH_3SH$ ,  $CH_3NH_2$ ,  $H_2S$

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This is the second of two papers presenting laboratory observations of small amounts of various gases that may occur in the atmospheres of the planets. In this paper are spectra of several hydrocarbons and  $H_2S$  in the spectral region of the lead sulfide detector (1.0–2.6 microns). The observations were made with the same scanning spectrometer (described by Kuiper, *et al.* 1962) that was used for the spectra in Paper I of this series (Kuiper and Cruikshank 1964). The information on the laboratory apparatus given here also applies to Paper I.

The absorption cells used in the laboratory were all single-pass tubes made of glass or, for non-corrosive gases, galvanized iron. Windows of glass or Suprasil (Englehard Industries, Inc.) were used, as well as glass lenses. An incandescent laboratory lamp was used for illumination. The spectrum was scanned at three different rates. For an initial reconnaissance, a scan was made in two increments (1.0–1.9 microns and 1.9–2.6 microns) at a rate of 5/1, which corresponds to 7.5 Å per sec and 495 Å per in. on the original records. The two spectral regions were then traced at a scan rate of 12.5/1 corresponding to 3 Å per sec and 198 Å per in. on the original records. For spectral regions containing bands of particular interest, higher resolution tracings were made with a scan rate of 25/1, corresponding to 1.5 Å per sec and 99 Å per in. on the original records. The time

constant was one second in all cases. The lead sulfide detector was  $0.1 \times 2.5$  mm in dimension. The spectrometer slit was 0.18 mm wide. A grating with 600 lines per millimeter and blazed for 1.6 microns first order was used. In the spectral range 1.0–1.9 microns a Corning 2540 filter was used, and in the range 1.9–2.6 microns, a Bausch and Lomb interference filter.

Assignments were recorded only for  $C_2H_2$  (Herzberg 1945, p. 290). Extensive references listed by Herzberg give additional information on band assignments.

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#### REFERENCES

- Kuiper, G. P., and Cruikshank, D. P. 1964, "Laboratory Spectra for Testing the Presence of Minor Constituents in Planetary Atmospheres, I:  $CH_4$ ,  $NH_3$ ,  $N_2O$ ,  $CO$ ,  $COS$ , Region 1–2.5  $\mu$ ," *Comm. LPL*, 2, 141–165.
- Kuiper, G. P., Goranson, R., Binder, A., and Johnson, H. L. 1962, "An Infrared Stellar Spectrometer," *Comm. LPL*, 1, 119–127.
- Herzberg, G. 1945, *Molecular Spectra and Molecular Structure, II, Infrared and Raman Spectra of Polyatomic Molecules* (Princeton: D. Van Nostrand Co.).

$C_2H_2$ , (a) 11 CM AT P = 3 CM, (b) 11 CM AT 1 ATM, 1.05 - 1.85  $\mu$

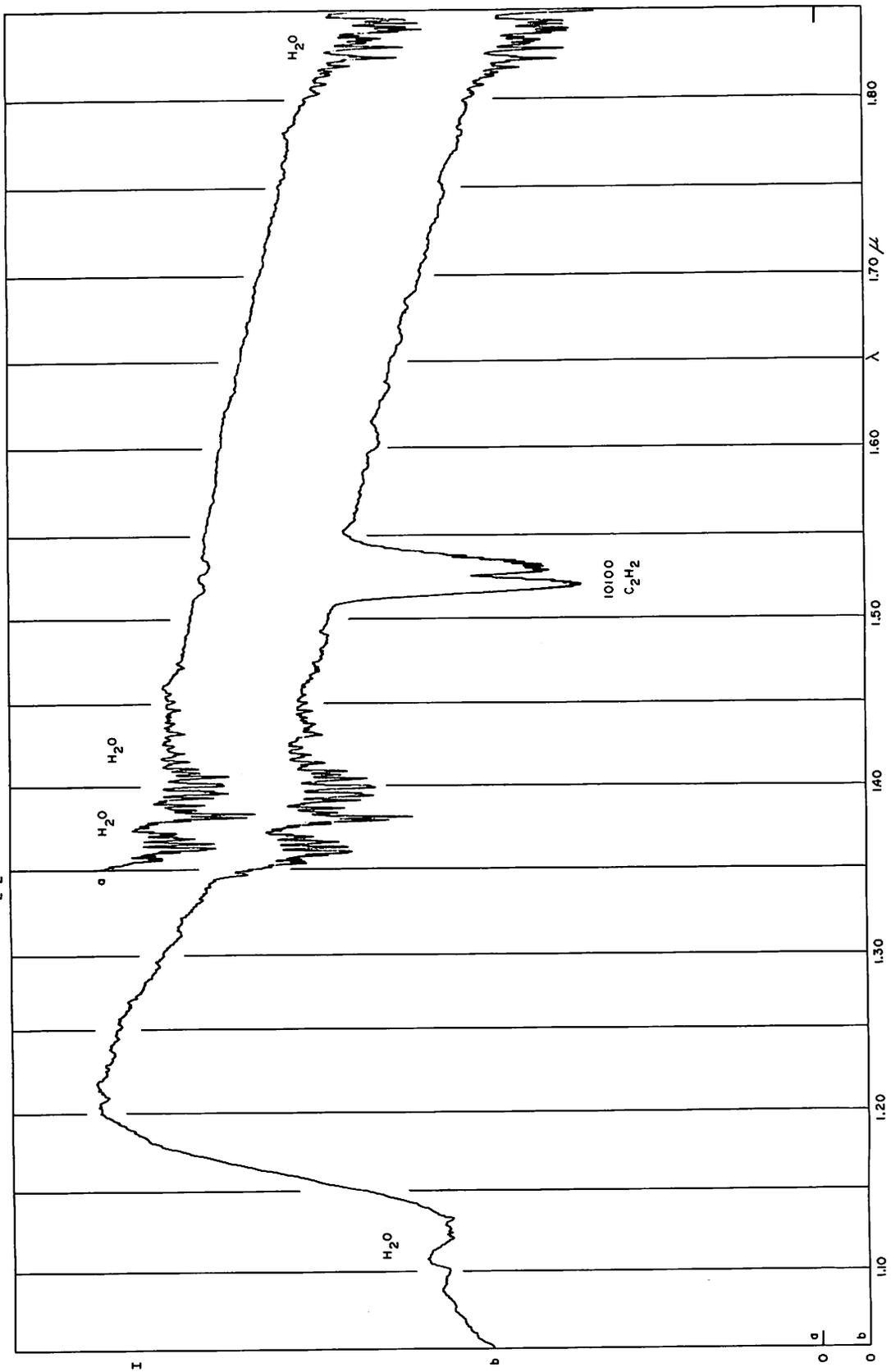


Fig. 1

$C_2H_2$ , 11 CM AT 1 ATM, 1.48-1.58  $\mu$

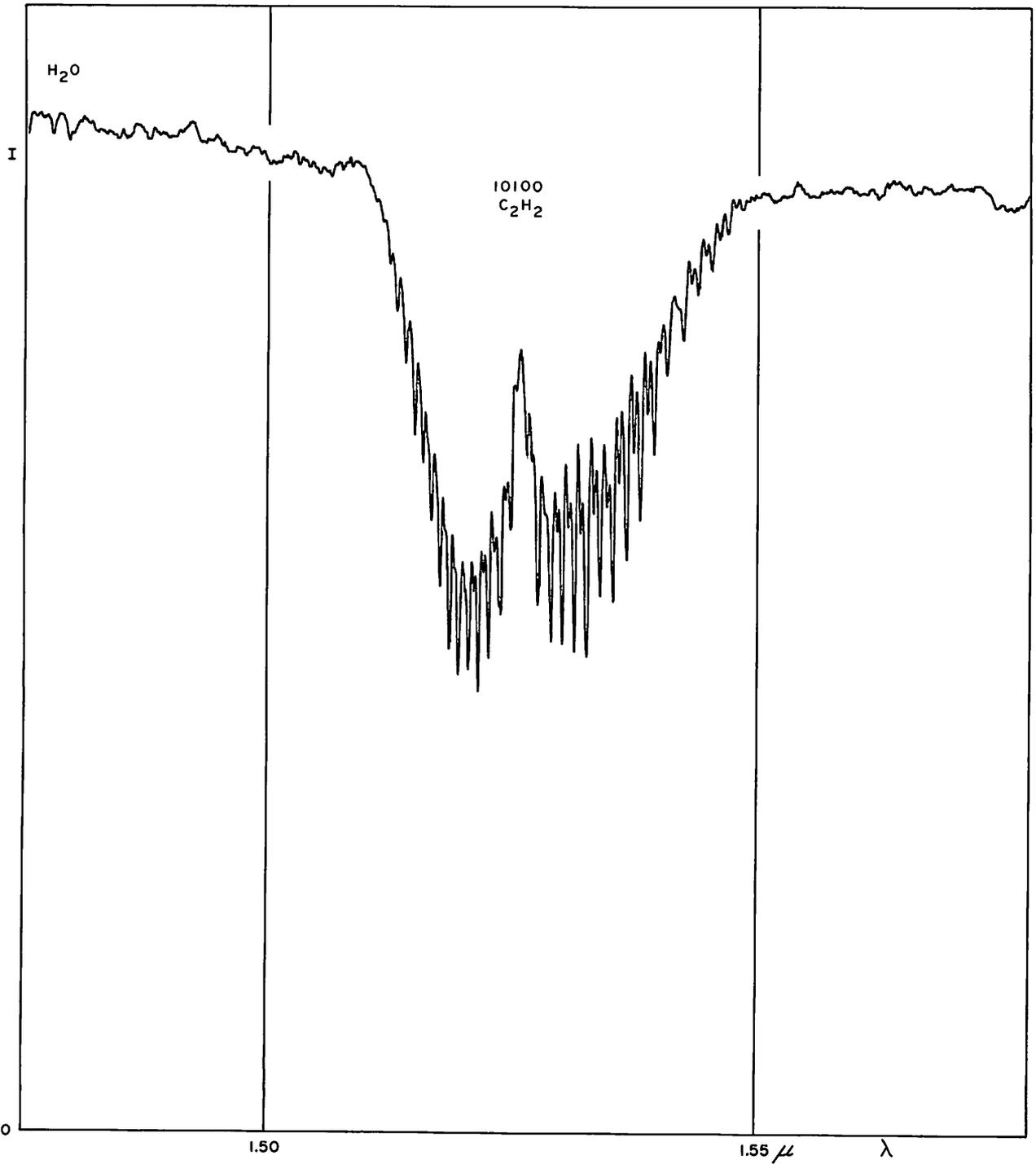


Fig. 2



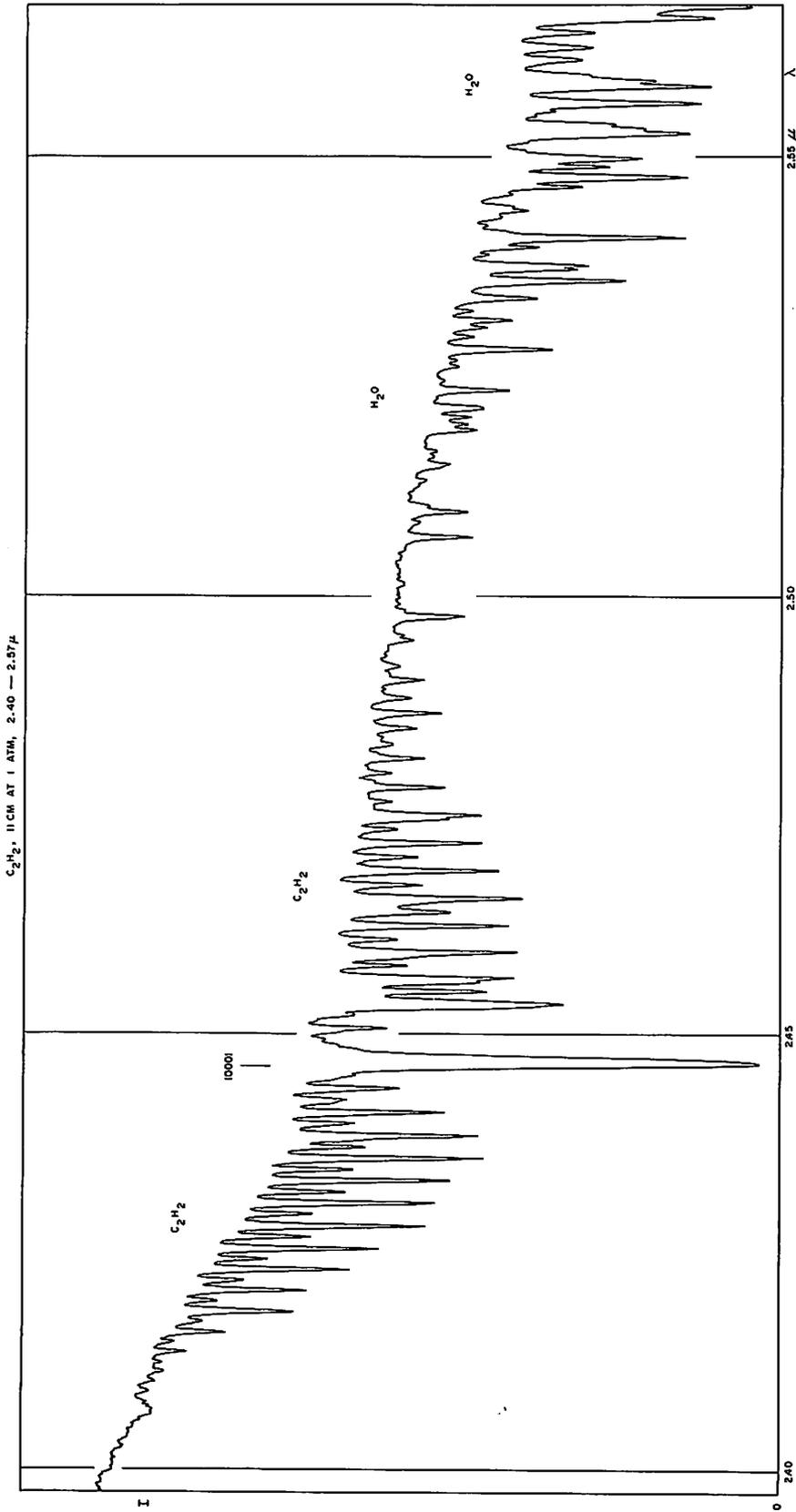


Fig. 4

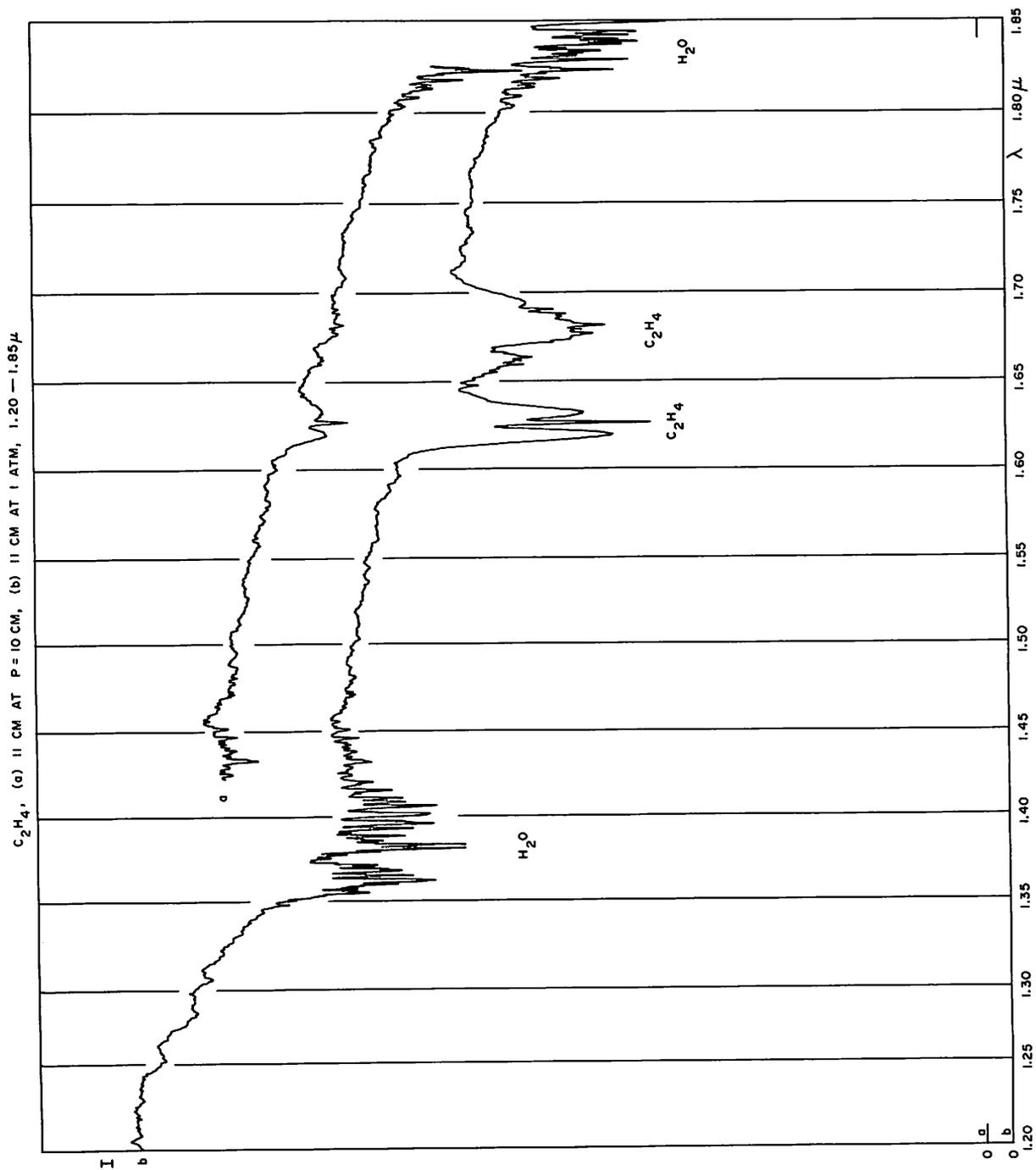


Fig. 5

$C_2H_4$ , 11 CM AT 1 ATM, 1.69-1.73  $\mu$

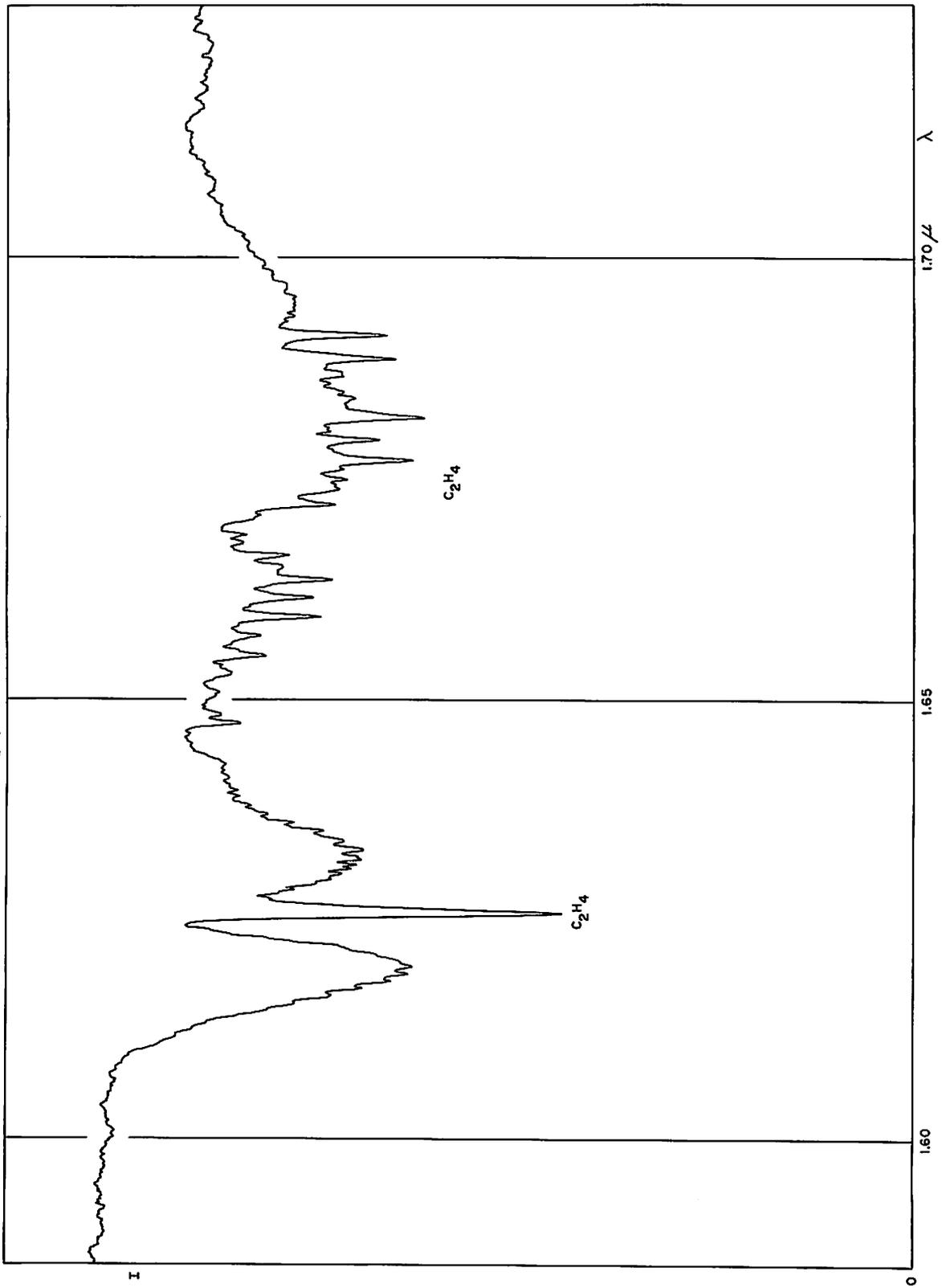


Fig. 6

C<sub>2</sub>H<sub>4</sub>, (a) 11 CM AT P=3 CM, (b) 11 CM AT P=10 CM, 1.90—2.55 μ

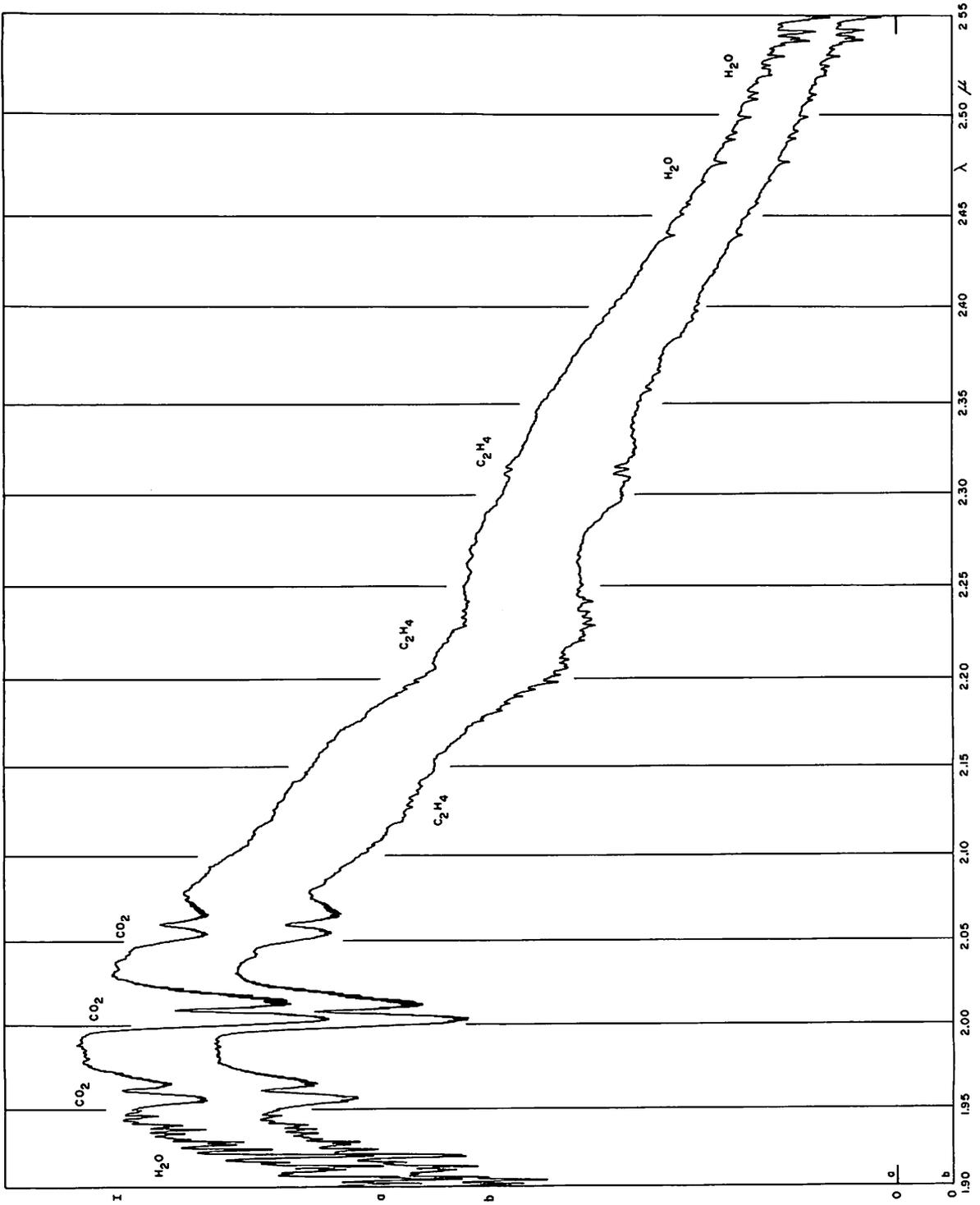


Fig. 7

C<sub>2</sub>H<sub>4</sub>, 11 CM AT 1 ATM, 1.85 — 2.57 μ

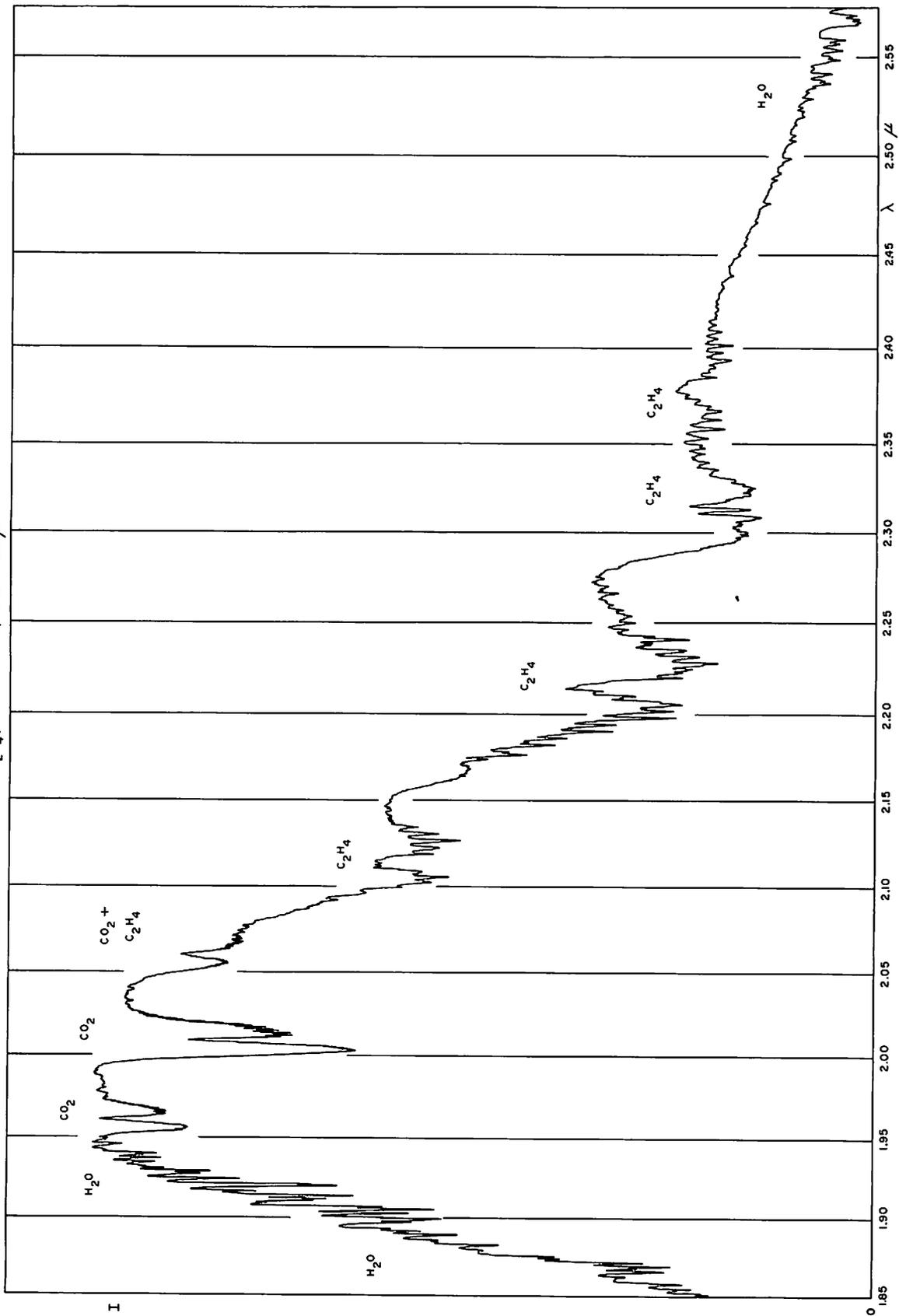


Fig. 8

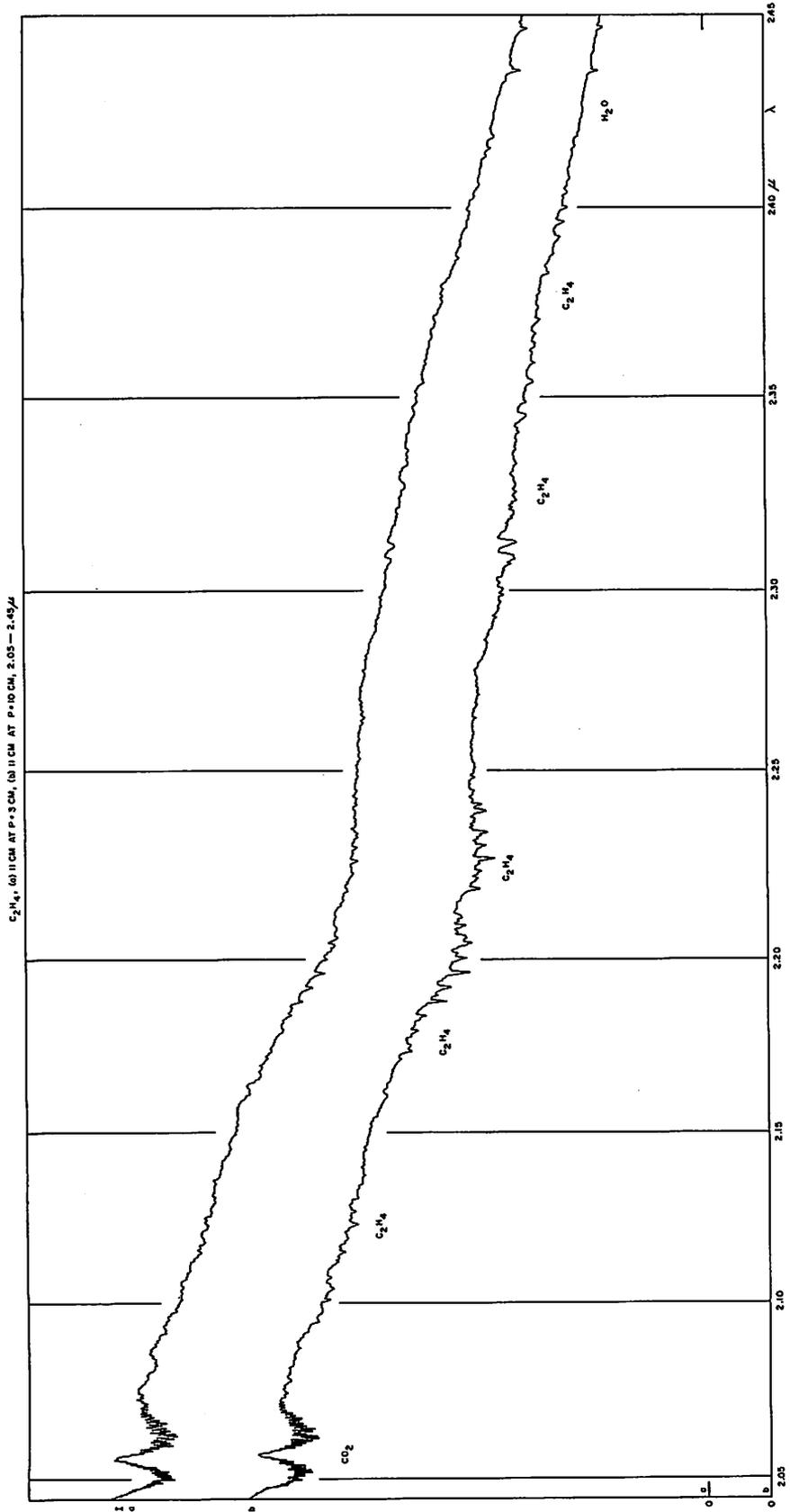


Fig. 9

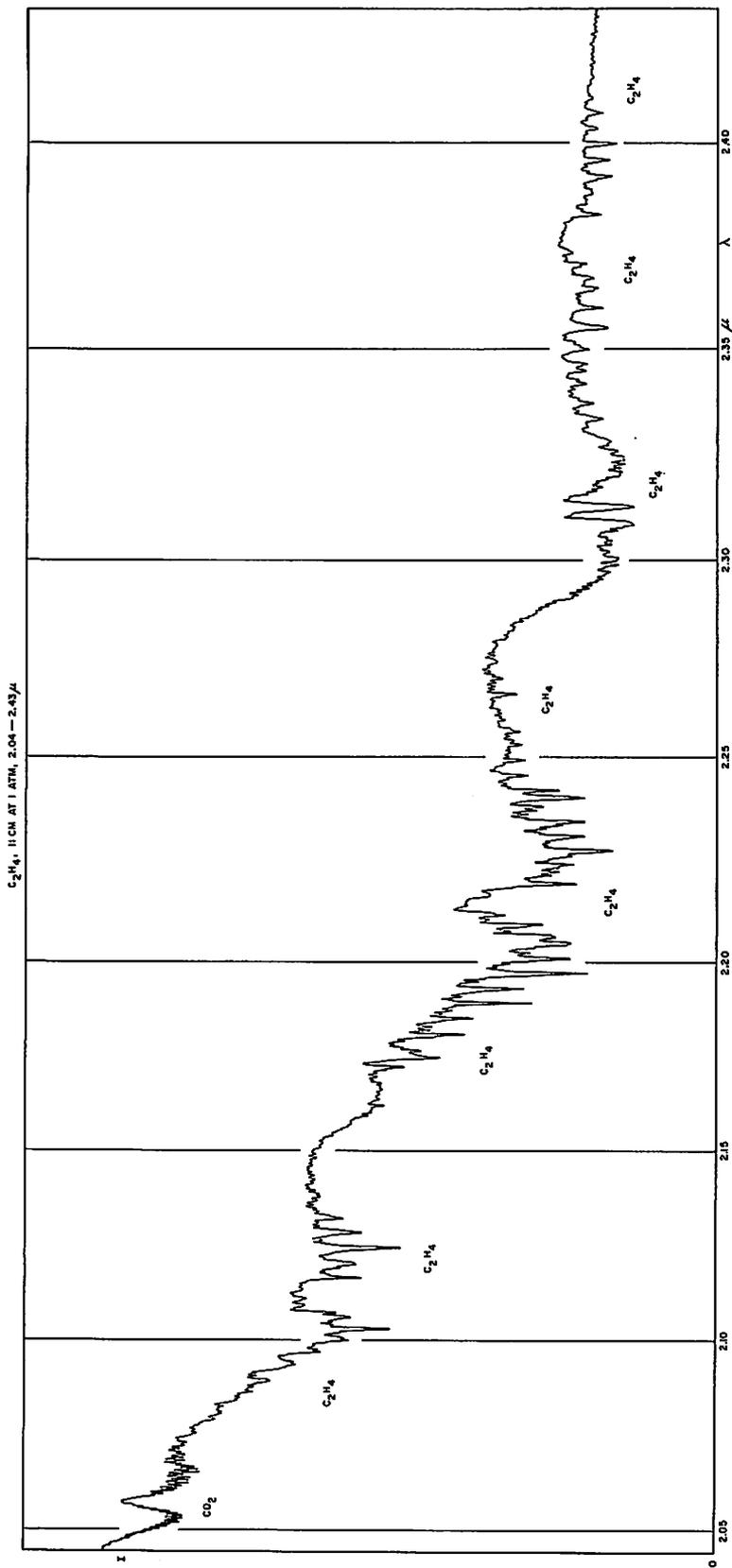


Fig. 10

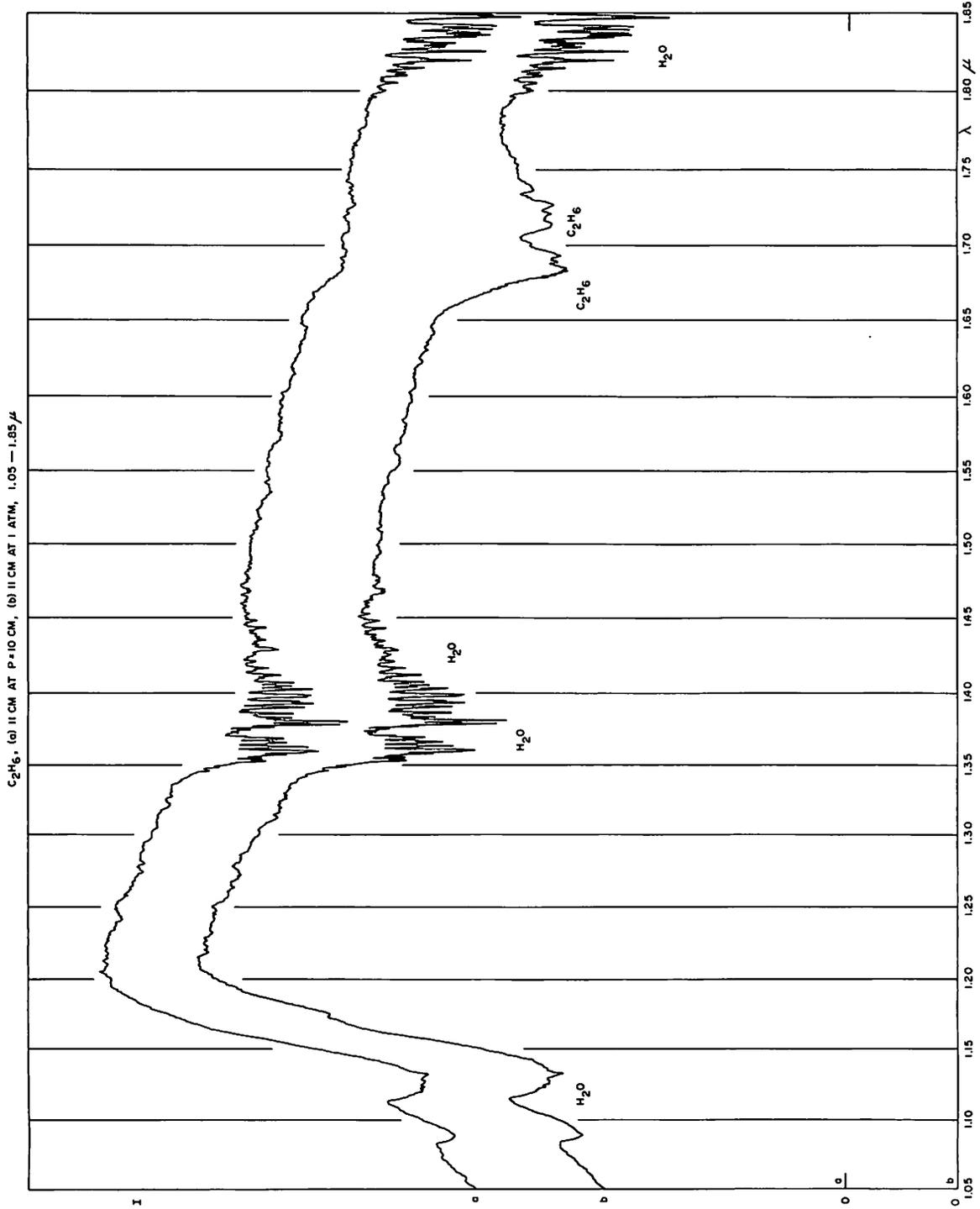


Fig. 11

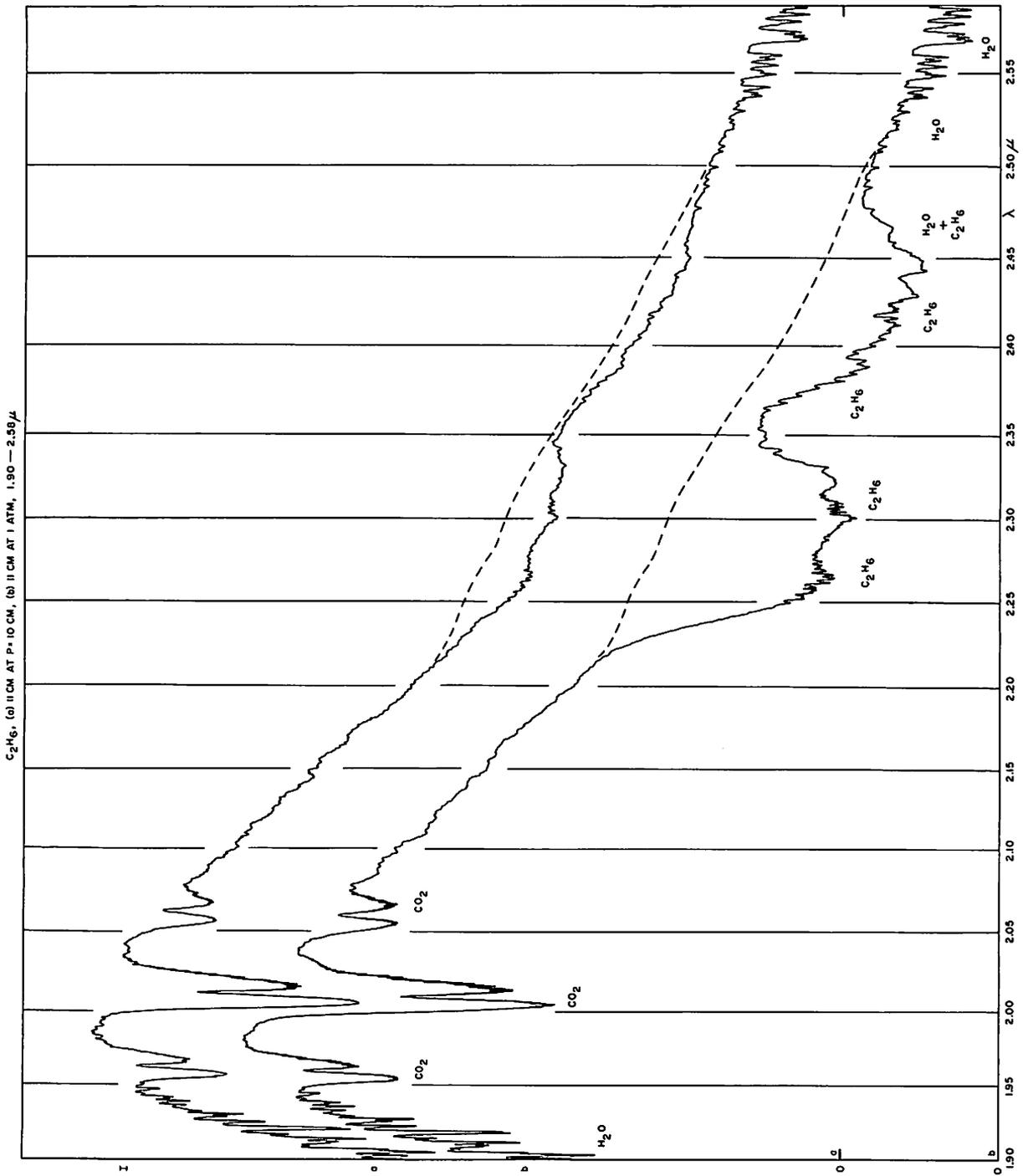


Fig. 12

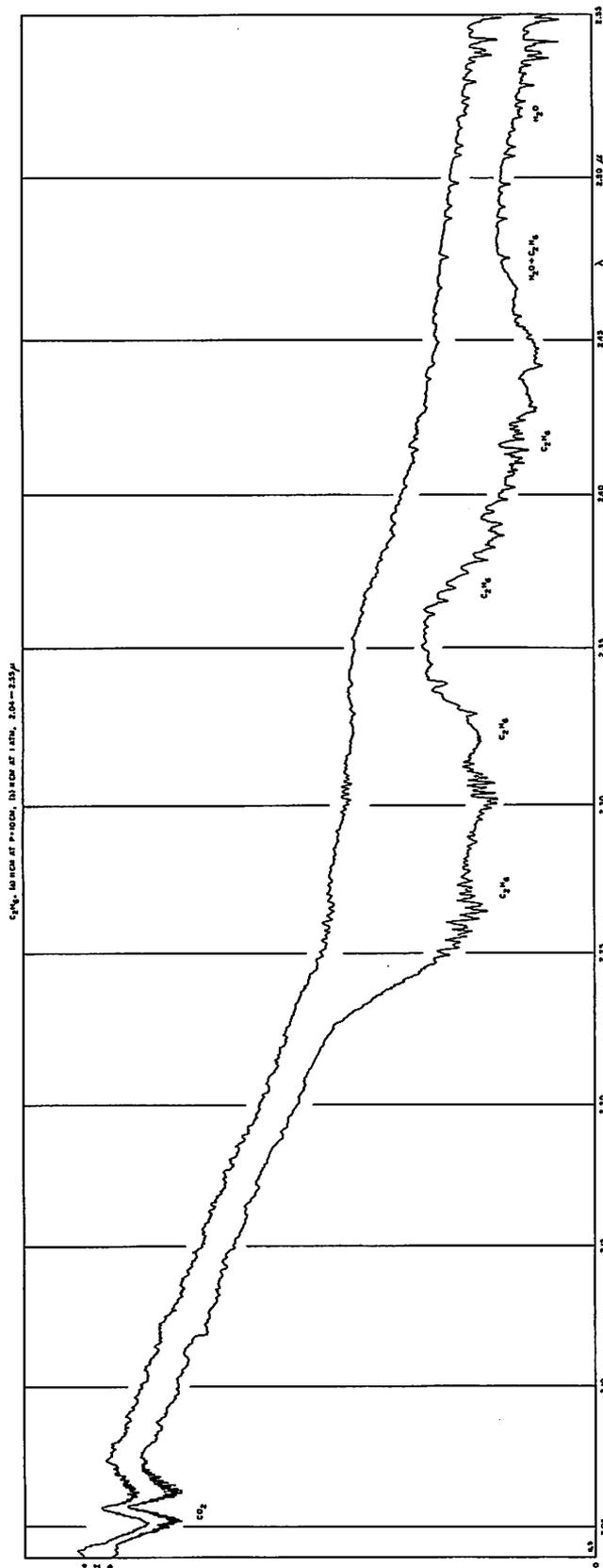


Fig. 13

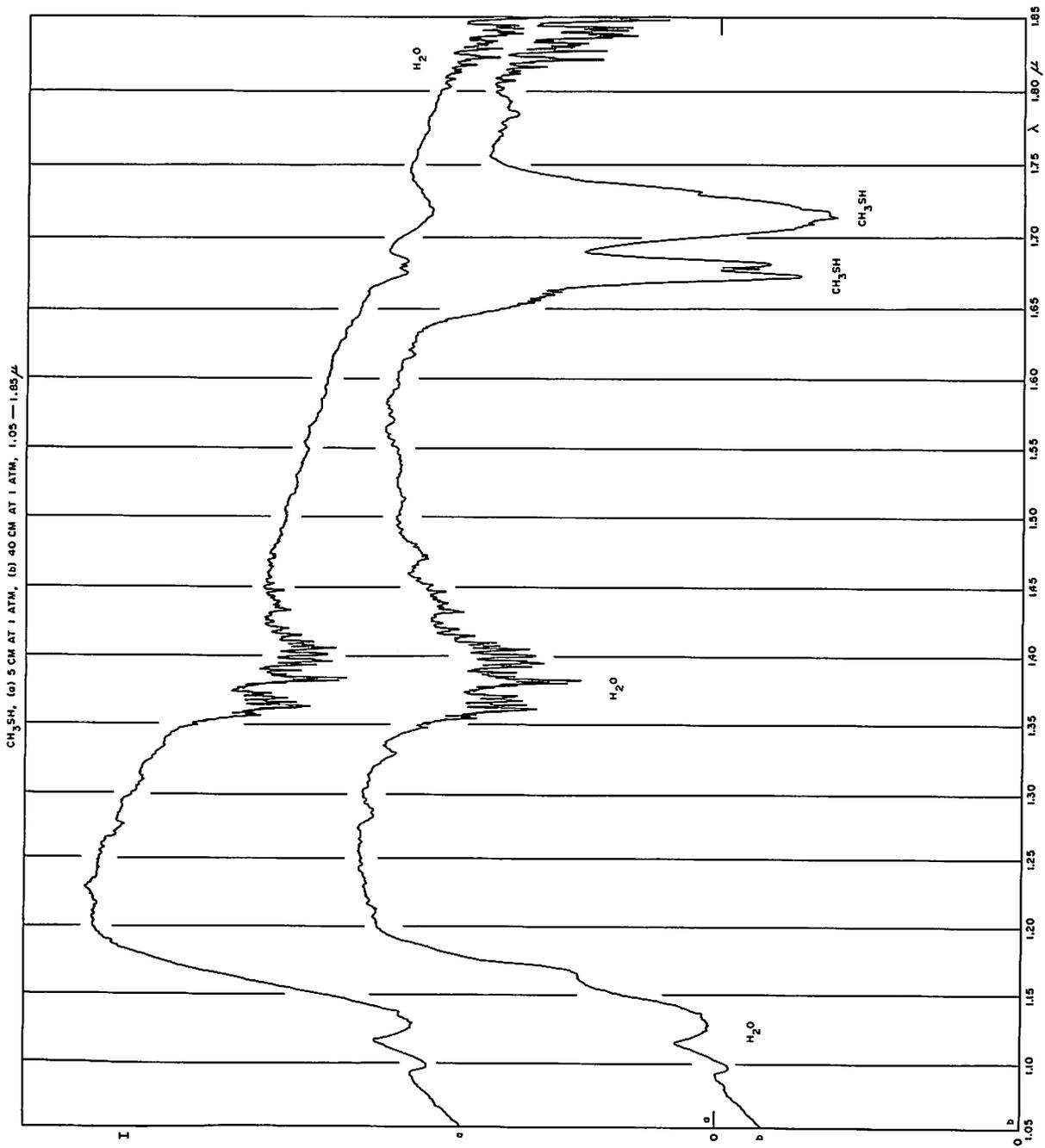


Fig. 14

CH<sub>3</sub>SH, 40 CM AT 1 ATM, 1.57 — 1.87 μ

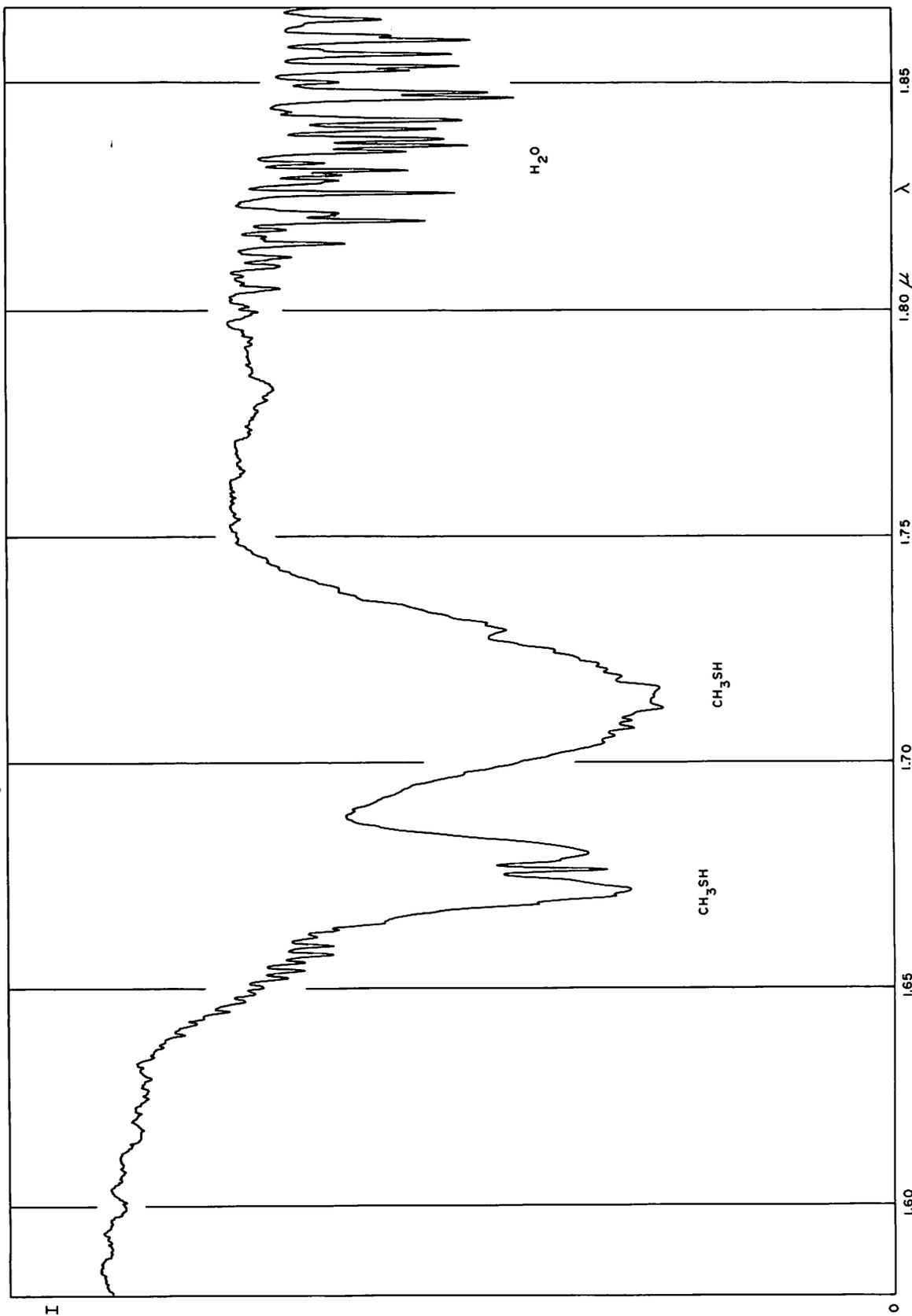


Fig. 15

CH<sub>3</sub>SH, 40 CM AT 1 ATM, 1.88 — 2.57μ

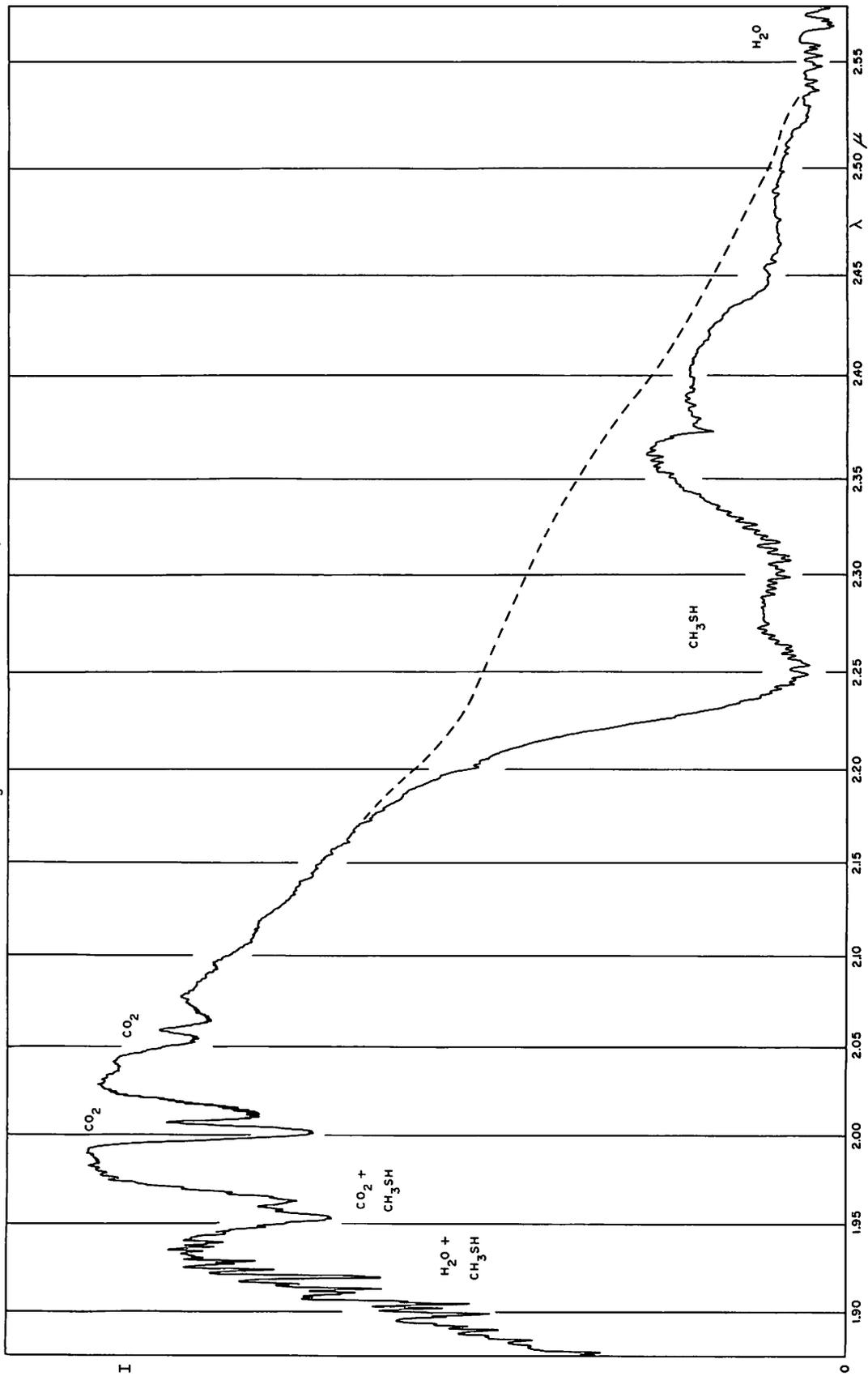


Fig. 16

CH<sub>3</sub>SH, 40 CM AT 1 ATM, 1.87 — 2.21 μ

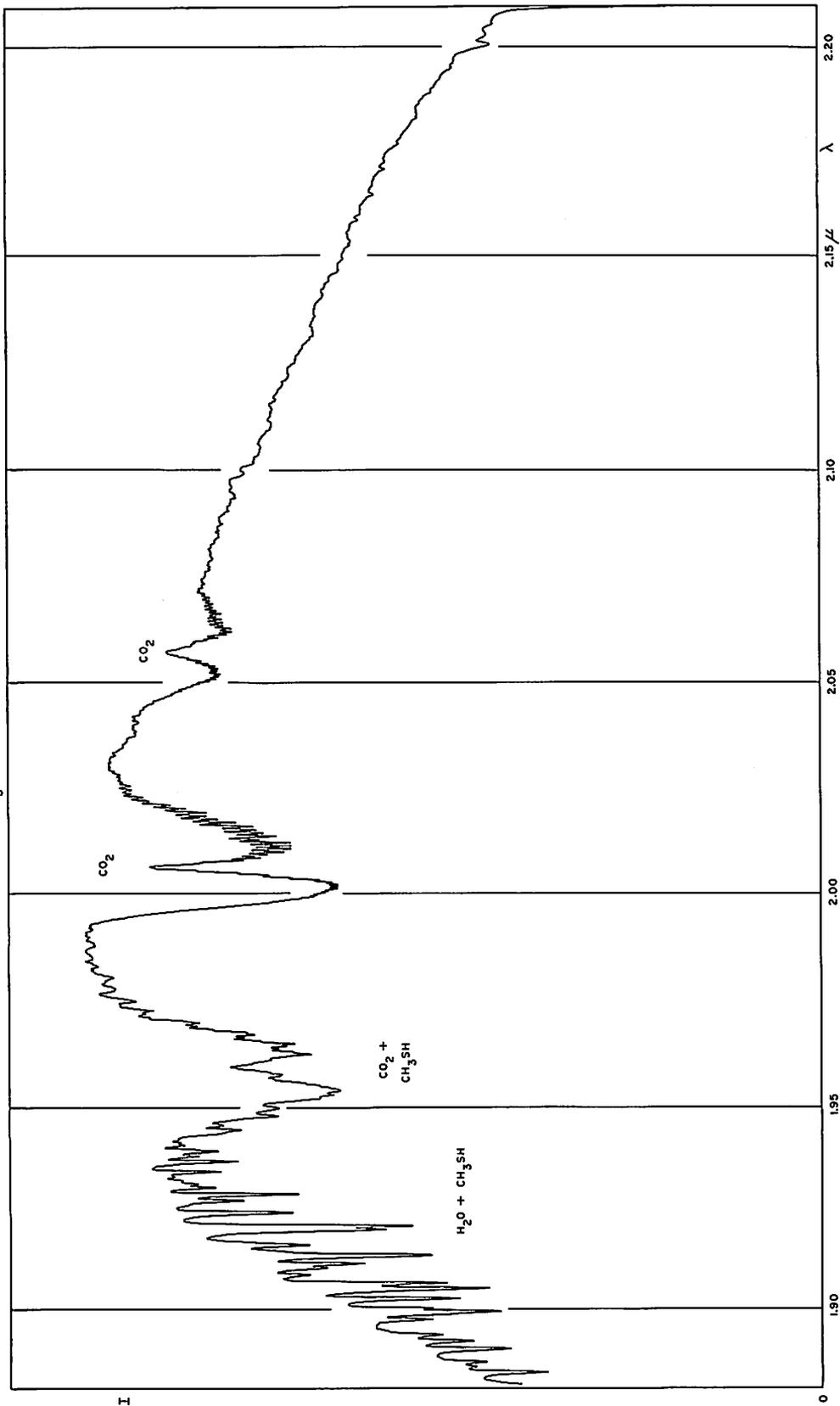


Fig. 17

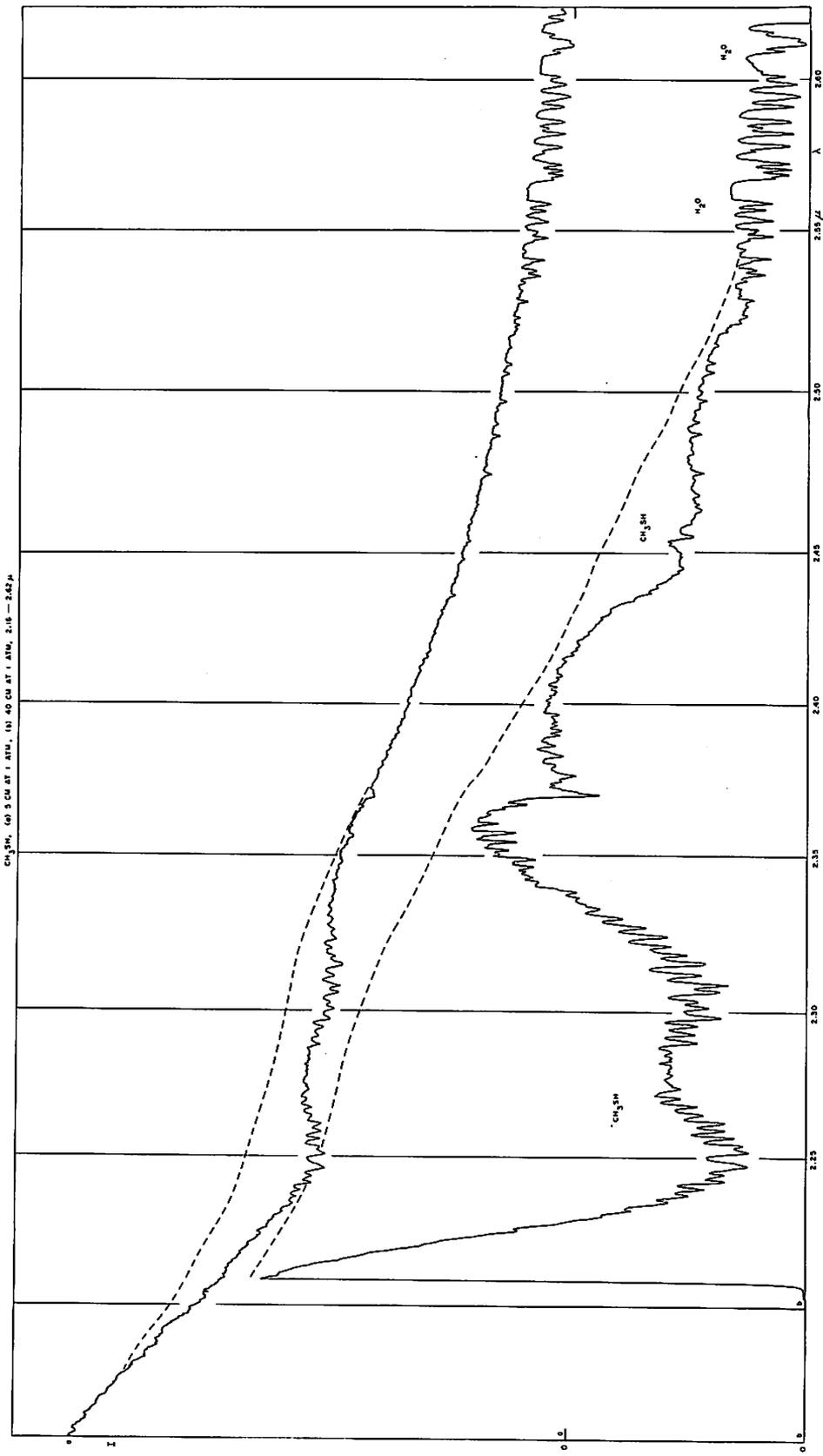


Fig. 18



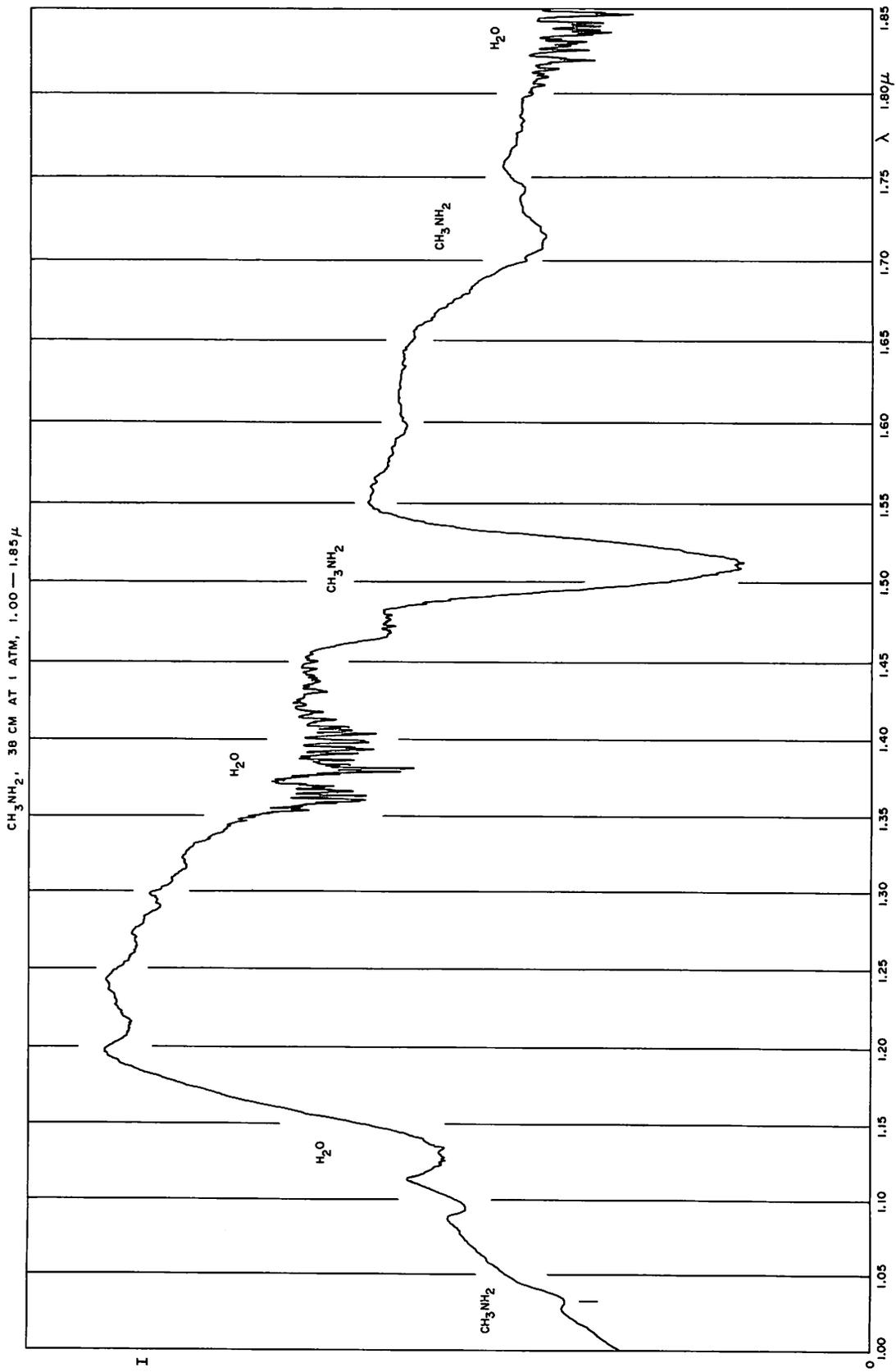


Fig. 20

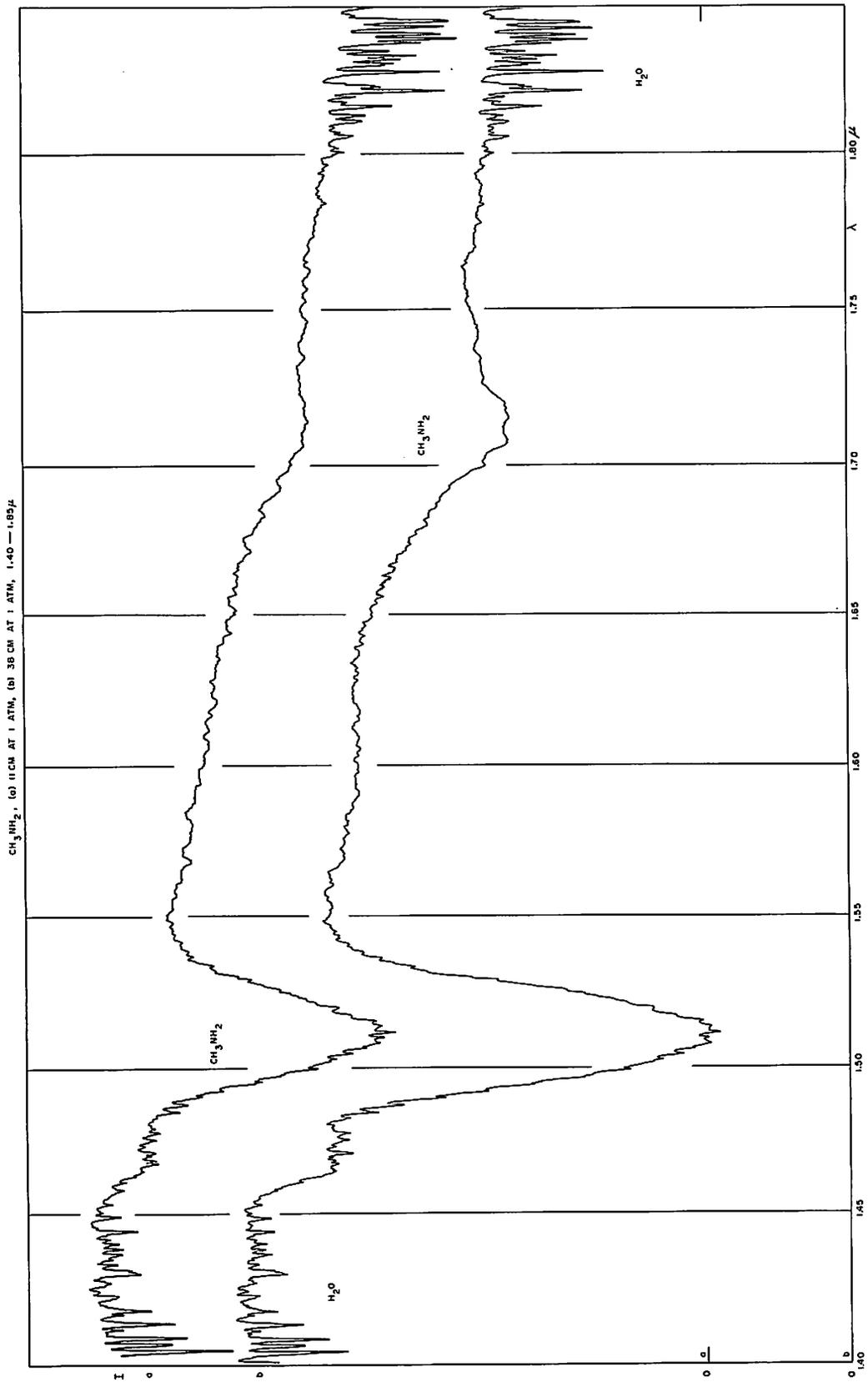


Fig. 21

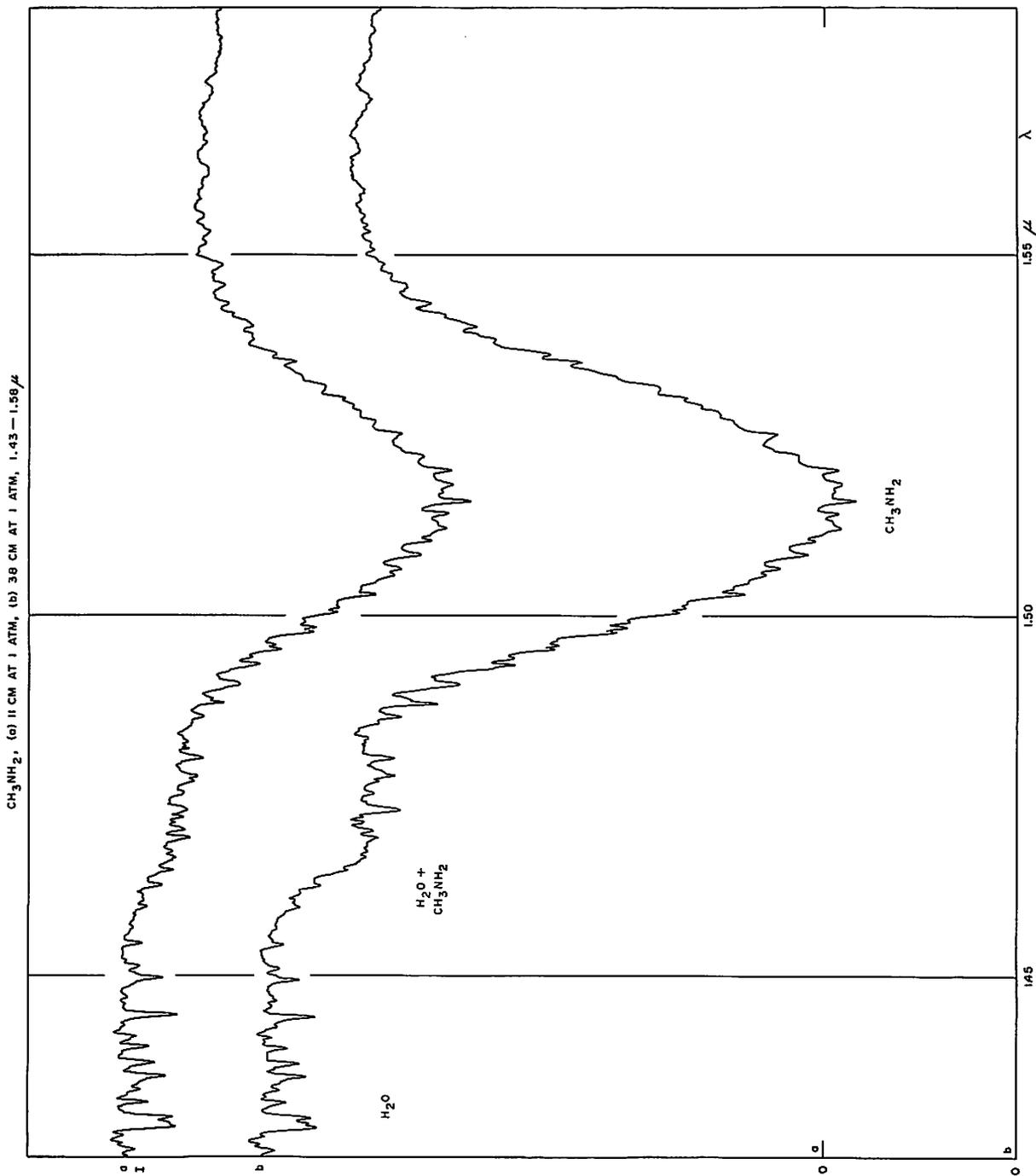


Fig. 22

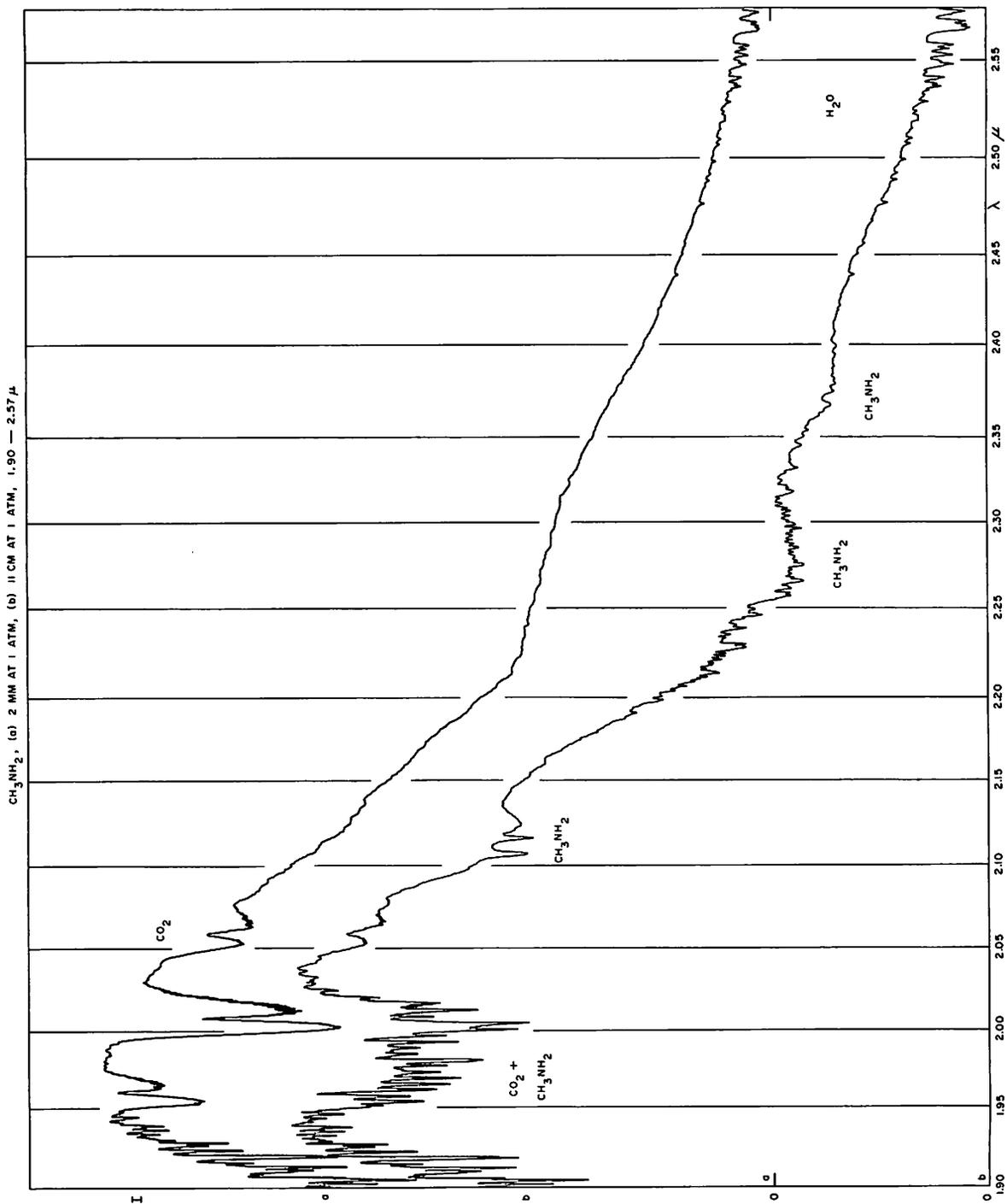


Fig. 23

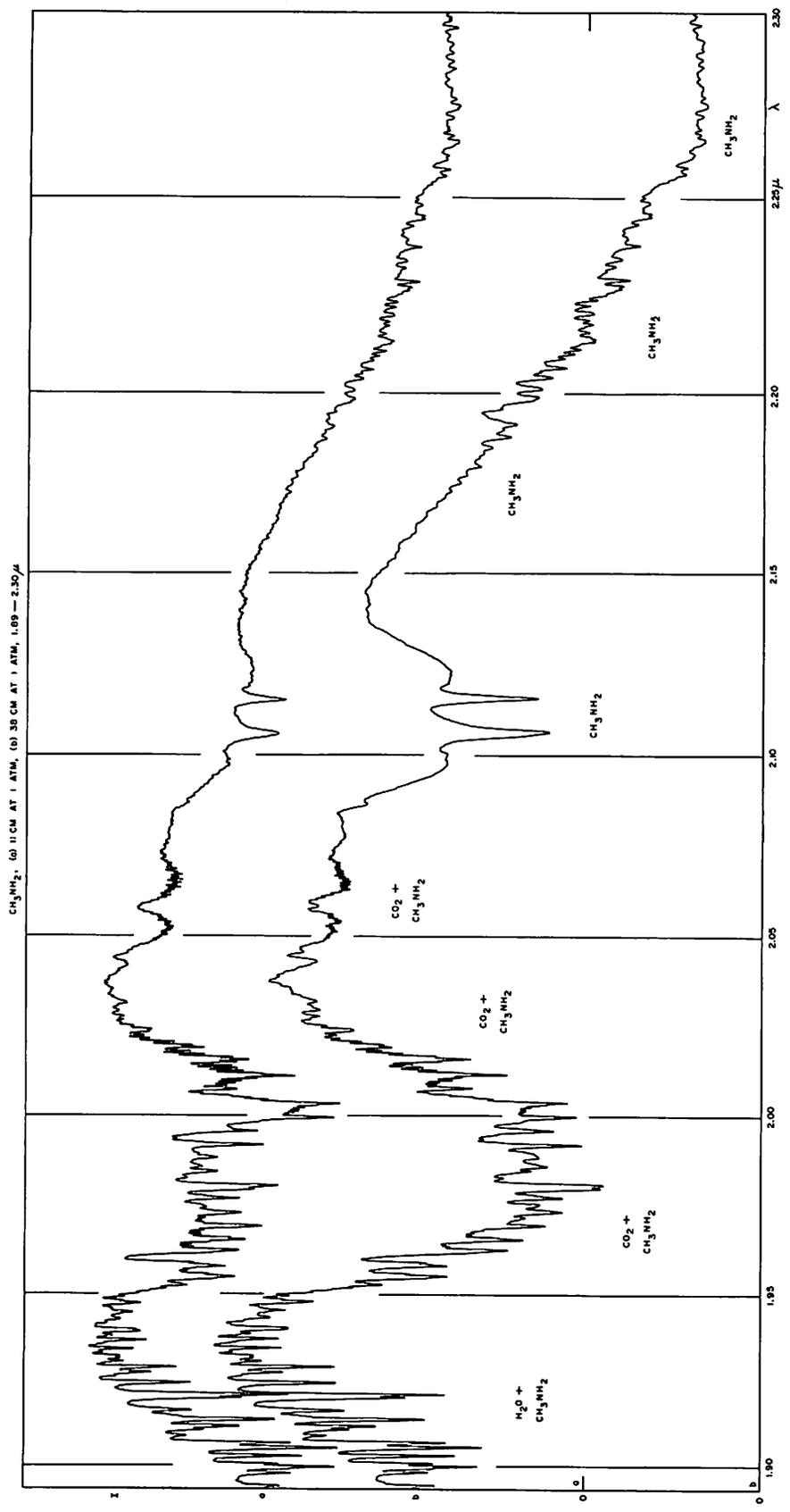


Fig. 24

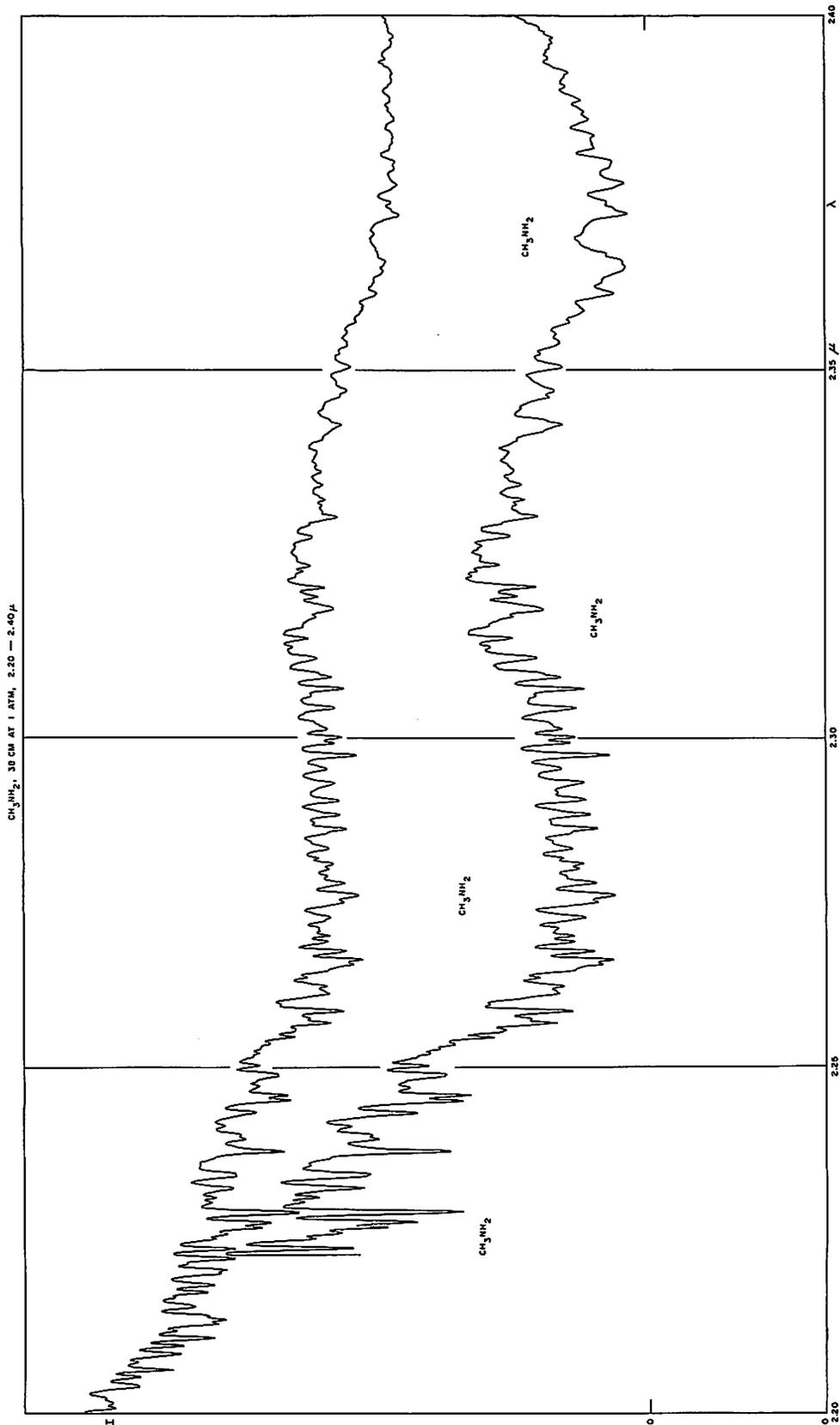


Fig. 25

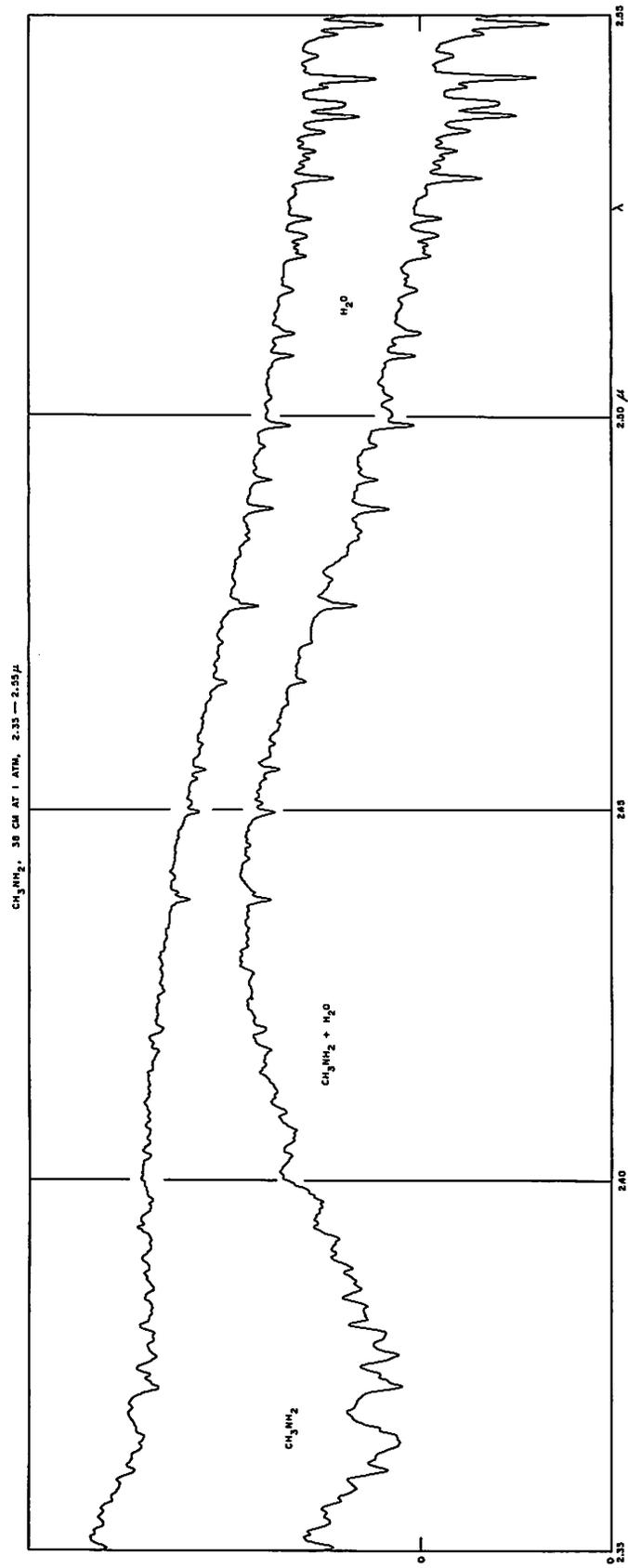


Fig. 26

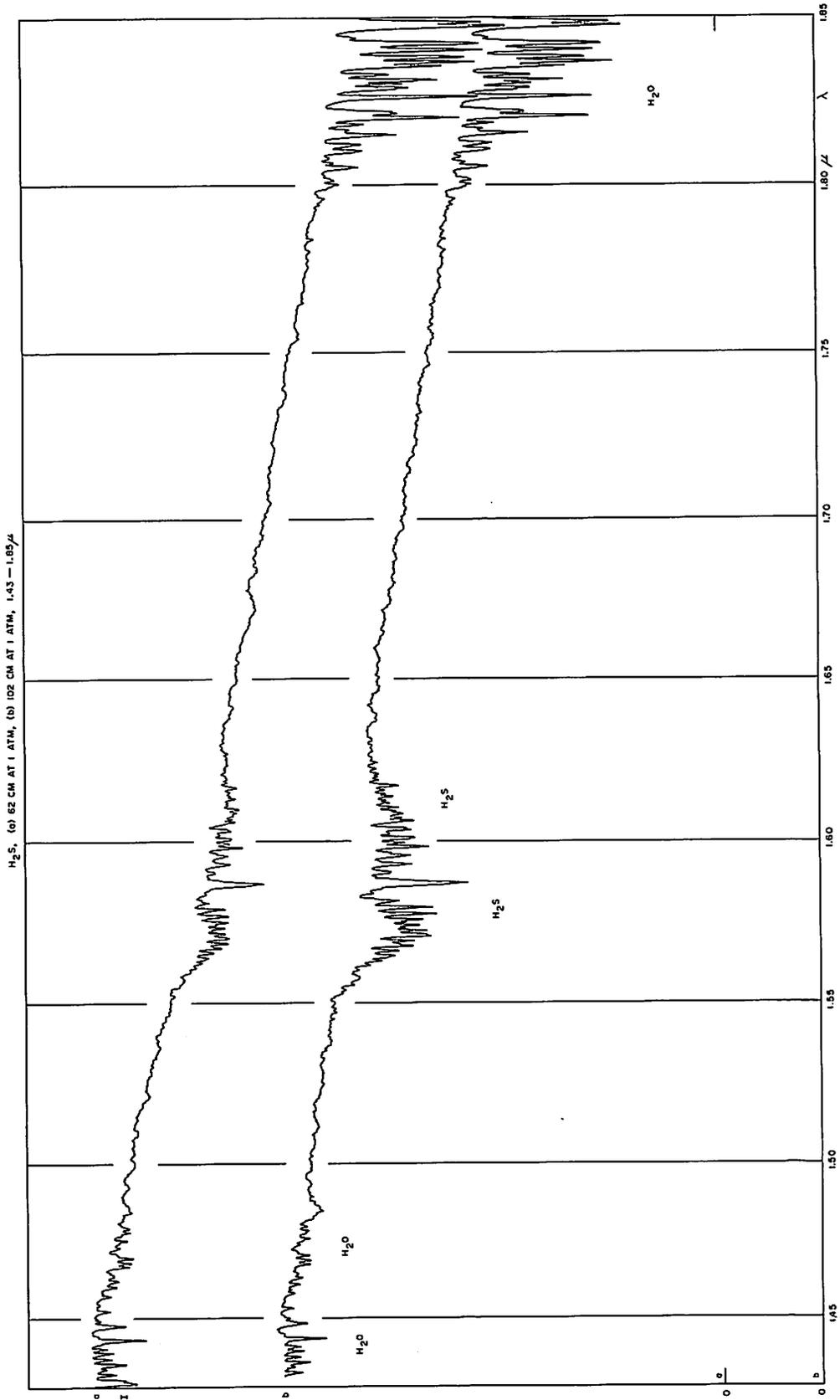


Fig. 27

H<sub>2</sub>S, (a) BLANK, (b) 11 CM AT 1 ATM, (c) 62 CM AT 1 ATM, (d) 102 CM AT 1 ATM, 1.90—2.55 μ

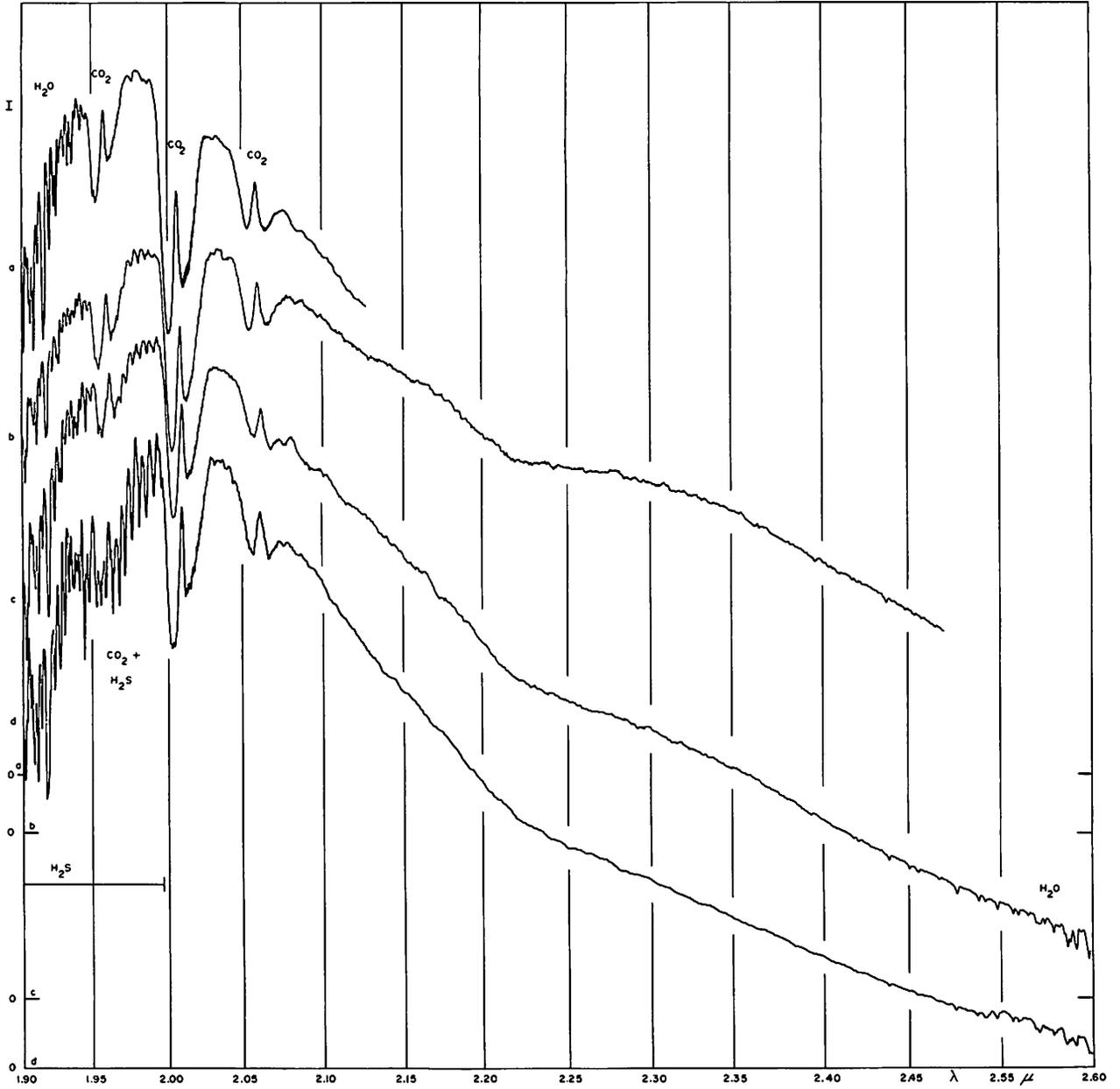


Fig. 28

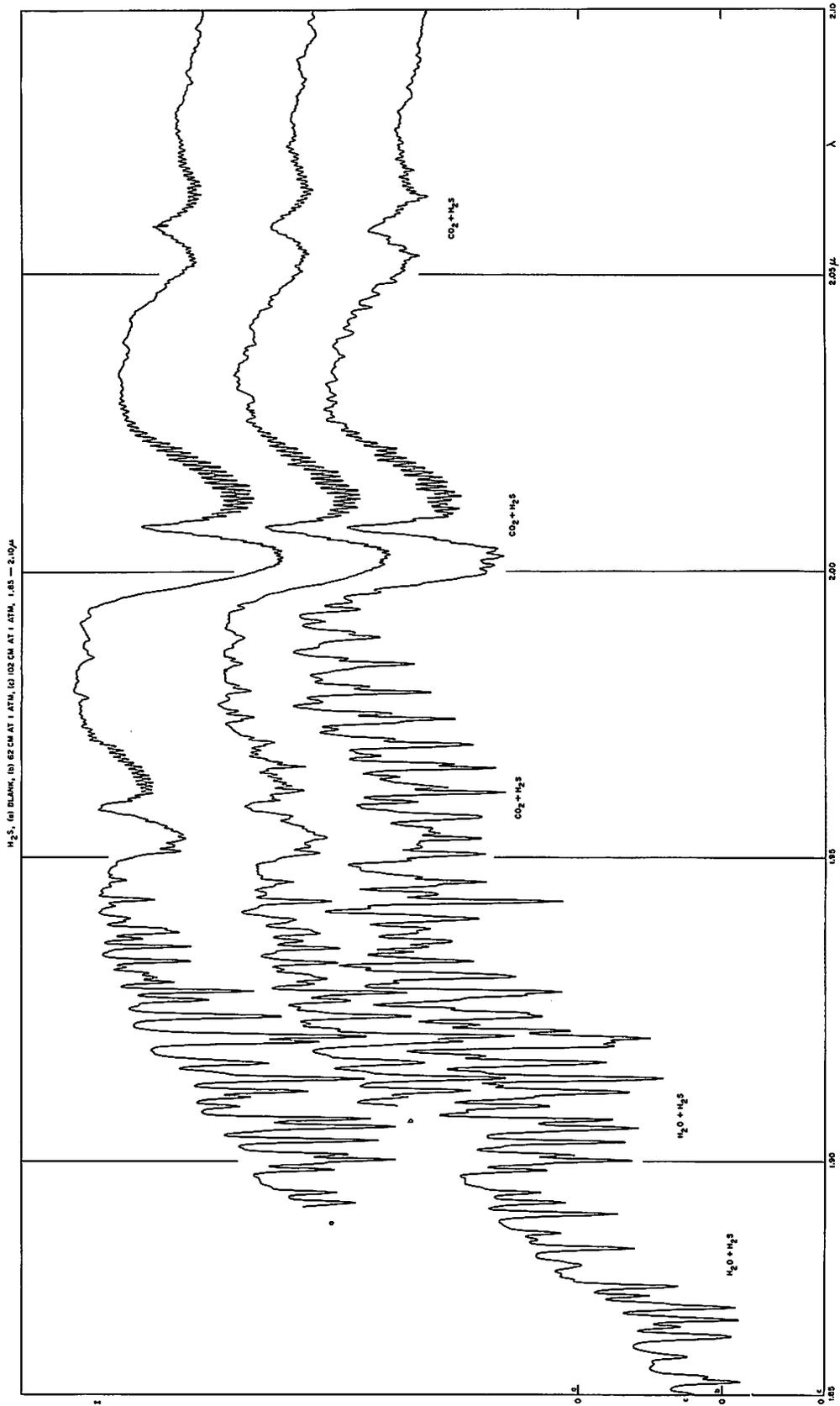


Fig. 29