

SOLUTIONS TO PROBLEM SET #1

#1 a) $\tilde{A} \times (\tilde{B} \times \tilde{C}) = \tilde{A} \times (\epsilon_{ijk} B_j C_k)$

$$= \epsilon_{mni} A_n (\epsilon_{ijk} B_j C_k)$$

$$= \epsilon_{mni} \epsilon_{ijk} A_n B_j C_k$$

$$= -\epsilon_{min} \epsilon_{ijk} A_n B_j C_k$$

$$= \epsilon_{imn} \epsilon_{ijk} A_n B_j C_k$$

$$= \epsilon_{ijk} \epsilon_{imn} A_n B_j C_k$$

$$= (\delta_{jm} \delta_{kn} - \delta_{jn} \delta_{km}) A_n B_j C_k$$

$$= \delta_{jm} \delta_{kn} A_n B_j C_k - \delta_{jn} \delta_{km} A_n B_j C_k$$

$$= A_n B_m C_n - A_n B_n C_m \quad (*)$$

$$= (A_n C_n) B_m - (A_n B_n) C_m$$

$\tilde{A} \times (\tilde{B} \times \tilde{C}) = (\tilde{A} \cdot \tilde{C}) \tilde{B} - (\tilde{A} \cdot \tilde{B}) \tilde{C}$

b) look @ terms on RHS separately and add.

$$\tilde{A} \times (\nabla \times \tilde{B}) = A_n \partial_m B_n - A_n \partial_n B_m$$

(using (*) above and $\nabla \rightarrow \partial_i$)

$$\tilde{B} \times (\nabla \times \tilde{A}) = B_n \partial_m A_n - B_n \partial_n A_m$$

$$(\tilde{A} \cdot \nabla) \tilde{B} = A_n \partial_n B_m$$

$$(\tilde{B} \cdot \nabla) \tilde{A} = B_n \partial_n A_m$$

Add these four to give

$$A_n \partial_m B_n - A_n \cancel{\partial_n B_m} + B_n \partial_m A_n - B_n \cancel{\partial_n A_m} + A_n \partial_n B_m \\ + B_n \cancel{\partial_n A_m}$$

$$= A_n \partial_m B_n + B_n \partial_m A_n$$

$$= \partial_m (A_n B_n) = \nabla (\tilde{A} \cdot \tilde{B})$$

This proves the identity

$$\text{c)} \quad \nabla \times \nabla \times \tilde{A} = \partial_n \partial_m A_n - \partial_n \partial_m A_m \\ = \partial_m \partial_n A_n - \partial_n^2 A_m$$

$$\boxed{\nabla \times \nabla \times \tilde{A} = \nabla (\nabla \cdot \tilde{A}) - \nabla^2 \tilde{A}}$$

$$\text{d)} \quad \nabla \cdot (\nabla \times \tilde{A}) = \partial_i \epsilon_{ijk} \partial_j A_k = \epsilon_{ijk} \partial_i \partial_j A_k^{(1)}$$

$$= \epsilon_{ijk} \partial_j \partial_i A_k \quad (\text{switching order of deriv.})$$

$$= -\epsilon_{ijk} \partial_i \partial_j A_k^{(2)} \quad (\text{switching index on } \epsilon)$$

\Rightarrow only way (1) & (2) can be true is if $\boxed{\nabla \cdot (\nabla \times \tilde{A}) = 0}$!

$$\#2 \quad \vec{F} = q\vec{E} + \frac{q}{c}\vec{v} \times \vec{B}$$

$$\vec{E} = 0 \quad \vec{B} = (0, 0, B) \quad B = \text{constant}$$

$$\Rightarrow m\dot{\vec{v}} = \frac{q}{c}\vec{v} \times \vec{B} \hat{z}$$

$$\Rightarrow m\dot{v}_x = \frac{qB}{c}v_y$$

$$m\dot{v}_y = -\frac{qB}{c}v_x$$

$$m\dot{v}_z = 0 \quad \Rightarrow \boxed{v_z = v_{z0} = \text{constant}}$$

Note:

$$m\dot{\vec{v}} = \frac{q}{c}\vec{v} \times \vec{B}$$

$$\Rightarrow \vec{v} \cdot (m\dot{\vec{v}}) = \frac{q}{c}\vec{v}(\vec{v} \times \vec{B}) = 0$$

$$\Rightarrow m\vec{v} \cdot \frac{d\vec{v}}{dt} = 0 \quad \Rightarrow \quad \underbrace{\frac{d}{dt}\left(\frac{1}{2}m(\vec{v} \cdot \vec{v})\right)}_{\text{Kinetic energy}} = 0$$

$$\Rightarrow \boxed{\text{kinetic energy} = \text{constant}}$$

define $\Omega = \frac{qB}{mc}$, and we have

$$\dot{v}_x = \Omega v_y \quad (1)$$

$$\dot{v}_y = -\Omega v_x \quad (2)$$

take $\frac{d}{dt}$ of (1) and subs. into (2)

$$\ddot{v}_x = \Omega \dot{v}_y = \Omega(-\Omega v_x) = -\Omega^2 v_x$$

$$\Rightarrow \ddot{v}_x + \Omega^2 v_x = 0$$

has solution $v_x = A \cos \Omega t + B \sin \Omega t$

$$@ t=0 \quad v_x = v_{x0} \Rightarrow A = v_{x0}$$

$$\Rightarrow v_x = v_{x0} \cos \Omega t + B \sin \Omega t$$

insert into (1) to give

$$-\Omega v_{x0} \sin \Omega t + B \Omega \cos \Omega t = \Omega v_y$$

$$\Rightarrow v_y = B \cos \Omega t - v_{x0} \sin \Omega t$$

$$@ t=0 \quad v_y = v_{y0} \Rightarrow B = v_{y0}$$

$$\therefore v_x = v_{x0} \cos \Omega t + v_{y0} \sin \Omega t$$

$$v_y = v_{y0} \cos \Omega t - v_{x0} \sin \Omega t$$

$$\text{define } \phi_0 = \begin{matrix} \text{initial} \\ \text{phase} \end{matrix} = \tan^{-1} \left(\frac{v_{y0}}{v_{x0}} \right)$$

$$V_{10} = \begin{matrix} \text{initial} \\ \text{speed} \\ \text{normal} \\ \text{to } B \end{matrix} = (v_{x0}^2 + v_{y0}^2)^{1/2}$$

we have

$$v_x = v_{\perp 0} \cos(\Omega t - \phi_0)$$

$$v_y = -v_{\perp 0} \sin(\Omega t - \phi_0)$$

$$v_z = v_{\parallel 0} \quad (v_{\parallel 0} = (v^2 - v_{\perp 0}^2)^{1/2})$$

note that $\dot{\tilde{x}} = \tilde{v}$

$$\Rightarrow \dot{x} = v_x = v_{\perp 0} \cos(\Omega t - \phi_0)$$

$$\Rightarrow x = \frac{v_{\perp 0}}{\Omega} \sin(\Omega t - \phi_0) + C$$

$$@ t=0 \quad x = x_0 \Rightarrow C = x_0 + \frac{v_{\perp 0}}{\Omega} \sin \phi_0$$

$$\therefore x = x_0 + \frac{v_{\perp 0}}{\Omega} [\sin(\Omega t - \phi_0) + \sin \phi_0]$$

$$x = x_0 + r_{g_0} [\sin(\Omega t - \phi_0) + \sin \phi_0]$$

where $r_{g_0} = \text{gyroadius} = \frac{v_{\perp 0}}{\Omega}$

Similarly $j = v_y = -v_{\perp 0} \sin(\Omega t - \phi_0)$

$$\Rightarrow y = \frac{v_{\perp 0}}{\Omega} \cos(\Omega t - \phi_0) + D$$

$$@ t=0 \quad y = y_0 \Rightarrow D = -\frac{v_{10}}{\omega} \cos \phi_0 + y_0$$

$$\therefore \boxed{y = y_0 + r_g [\cos(\omega t - \phi_0) - \cos \phi_0]}$$

and, lastly

$$\dot{z} = v_z = v_{z0} = v_{10} = \text{constant}$$

$$\Rightarrow \boxed{z = z_0 + v_{10} t}$$

This completes the complete set of solutions for the orbit of the particle.

#3 Start w/ conservation of magnetic moment

$$\frac{w_L}{B} = \text{const}$$

$$\Rightarrow \left(\frac{w_L}{B} \right)_{\text{release}} = \left(\frac{w_L}{B} \right)_{\text{mirror pt.}}$$

$$\Rightarrow \left(\frac{W \sin^2 \alpha}{B} \right)_{\text{release}} = \left(\frac{W \sin^2 \alpha}{B} \right)_{\text{mirror pt.}}$$

at release $\alpha = \alpha_0$

$$B = B(r_0, \lambda_0) \quad \lambda = \text{latitude}$$

at mirror pt. $\alpha = \pi/2$

$$B = B(r_m, \lambda_m)$$

$W = \text{constant.}$

$$\Rightarrow \frac{\sin^2 \alpha_0}{B(r_0, \lambda_0)} = \frac{1}{B(r_m, \lambda_m)}$$

$$B(r, \lambda) = \frac{M}{r^3} (1 + 3 \sin^2 \lambda)$$

$$\frac{r_0^3 \sin^2 \alpha_0}{1 + 3 \sin^2 \alpha_0} = \frac{r_m^3}{1 + 3 \sin^2 \lambda_m}$$

the final lens equation is

$$r = r_0 \cos^2 \lambda$$

$$\Rightarrow r_m = r_0 \cos^2 \lambda_m$$

thus, we have

$$r_0^3 \sin^2 \alpha_0 = \frac{r_0^3 \cos^6 \lambda_m}{1 + 3 \sin^2 \lambda_m}$$

$$\Rightarrow \sin^2 \alpha_0 = \frac{\cos^6 \lambda_m}{1 + 3 \sin^2 \lambda_m}$$

must solve for λ_m . This could be done on a computer, or by trial and error. But, since the release pitch angle (70°) is not far from 90° , its minor pt., it will likely not move far from the equator; so, let's use small angle formulas

for $\lambda_m \ll 1$

$$\cos \lambda_m \approx 1 - \frac{1}{2} \lambda_m^2$$

$$\sin \lambda_m \approx \lambda_m$$

Correct to second order in λ_m

We have

$$\sin^2 \alpha_0 = \frac{1 - 3\lambda_m^2}{1 + 3\lambda_m^2} \quad \leftarrow \text{using binomial approx.}$$

$$\Rightarrow (1 + 3\lambda_m^2) \sin^2 \alpha_0 = 1 - 3\lambda_m^2$$

$$\Rightarrow \lambda_m^2 (3 \sin^2 \alpha_0 + 3) = 1 - \sin^2 \alpha_0$$

$$\Rightarrow \lambda_m \approx \left[\frac{1}{3} \left(\frac{1 - \sin^2 \alpha_0}{1 + \sin^2 \alpha_0} \right) \right]^{1/2}$$

$$\alpha_0 = 70^\circ$$

leads to $\lambda_m \approx 0.14$ radians

$$\Rightarrow \boxed{\lambda_m \approx 8^\circ}$$

Note that this is correct to order $\lambda_m^2 \approx 0.02$
or, about 2%.

#4 (a) 1 MeV proton ; $E = 10^6 \text{ eV}$

$$= 10^6 \times (1.6022 \times 10^{-12} \text{ erg})$$

$$= 1.6022 \times 10^{-6} \text{ erg}$$

non relativistic

$$\therefore v = \left(\frac{2E}{m} \right)^{1/2} = \left[\frac{(2)(1.6022 \times 10^{-6})}{1.6726 \times 10^{-24}} \right]^{1/2} \frac{\text{cm}}{\text{s}}$$

$$v = 1.384 \times 10^9 \text{ cm/s}$$

$$(b) r_g = v/\Omega \quad ; \quad \Omega = \frac{qB}{mc} = \frac{(4.8032 \times 10^{-10})(5 \times 10^{-4})}{(1.6726 \times 10^{-24})(2.9979 \times 10^9)} \text{ s}^{-1}$$

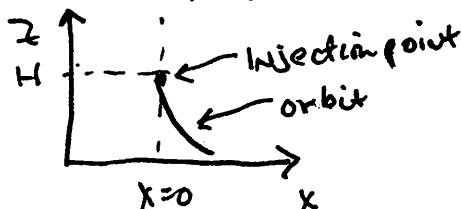
$$= 4.79 \text{ sec}^{-1}$$

$$\therefore r_g = \frac{1.384 \times 10^9}{4.79} \text{ cm} = 2.89 \times 10^8 \text{ cm}$$

$$r_g = 2890 \text{ km}$$

$$\therefore \frac{r_g}{H} = \frac{2890}{50} = 57.8$$

(c) consider a side view, y is into paper, \vec{B} points out of the paper, x is to the right, z is up



Because the gyroradius is so large, the particle only completes a very small part of its orbit before striking the surface. Thus, it is reasonable to estimate that its acceleration in the x direction is constant.

Thus,

$$m \frac{dv_x}{dt} = -\frac{q}{c} v_z B_y = \frac{q}{c} v_z B \approx \frac{q}{c} v B$$

$= \text{constant}$

$$\therefore \frac{dv_x}{dt} = v \Omega = \text{const.}$$

$$\Rightarrow v_x = v_{x_0} + v \Omega t = v \Omega t$$

$$\therefore \frac{dx}{dt} = v_x = v \Omega t$$

$$\Rightarrow x = x_0 + \frac{1}{2} v \Omega t^2 = \frac{1}{2} v \Omega t^2$$

t is just H/v , the time it takes to hit the surface from a height $Z=H$. Note that, initially, all its speed is in the downward direction

$$\therefore \Delta x = \frac{1}{2} v \Omega t^2 = \frac{1}{2} v \Omega \left(\frac{H}{v}\right)^2 = \frac{1}{2} \frac{H^2 \Omega}{v}$$

$$= \frac{1}{2} \frac{H}{r_g} H = \frac{1}{2} \frac{H}{r_g} H$$

From part (b), we have $H/r_g = \frac{1}{57.8}$, thus

$$\Delta x = \frac{1}{2} \left(\frac{1}{57.8} \right) 50 \text{ km}$$

$$= \frac{50}{2(57.8)} \text{ km}$$

$$\boxed{\Delta x = 0.432 \text{ km}}$$

#5 we note that since B is in \hat{r} direction only and there is no electric field, the equations of motion imply

$$v_r = \text{constant} \quad (\text{i.e. } m \ddot{v}_r = 0 \Rightarrow v_r = \text{const})$$

The equations of motion are written

$$m \ddot{\tilde{r}} = m \ddot{r} = \frac{q}{c} \frac{K}{r^2} \tilde{r} \times \hat{r}$$

$$= \frac{q K}{c r^2} \dot{\tilde{r}} \times \hat{r}$$

$$\Rightarrow \ddot{\tilde{r}} = \frac{q K}{mc} \frac{\dot{\tilde{r}} \times \hat{r}}{r^2}$$

$$\ddot{\tilde{r}} = \frac{q K}{mc} \frac{\dot{\tilde{r}} \times \hat{r}}{r^3}$$

Consider the angular momentum vector (divided by m)

$$\tilde{L} = \tilde{r} \times \dot{\tilde{r}} \quad \leftarrow \text{angular momentum divided by } m$$

note:

$$\frac{d\tilde{L}}{dt} = \dot{\tilde{r}} \times \dot{\tilde{r}} + \tilde{r} \times \ddot{\tilde{r}}$$

$$= \tilde{r} \times \left[\frac{q K}{mc} \frac{\dot{\tilde{r}} \times \hat{r}}{r^3} \right]$$

$$\begin{aligned}
 \frac{d\mathbf{L}}{dt} &= \frac{qK}{mc} \frac{1}{r^3} \mathbf{r} \times (\dot{\mathbf{r}} \times \dot{\mathbf{r}}) \\
 &= \frac{qK}{mc} \frac{1}{r^3} \left[(\mathbf{r} \cdot \dot{\mathbf{r}}) \dot{\mathbf{r}} - (\dot{\mathbf{r}} \cdot \dot{\mathbf{r}}) \mathbf{r} \right] \\
 &= \frac{qK}{mc} \frac{1}{r^3} (r^2 \dot{\mathbf{r}} - r \dot{r} \mathbf{r}) \\
 &= \frac{qK}{mc} \left(\frac{\dot{\mathbf{r}}}{r} - \frac{\dot{r} \mathbf{r}}{r^2} \right) \\
 &= \frac{qK}{mc} \frac{d}{dt} \left(\frac{\mathbf{r}}{r} \right)
 \end{aligned}$$

vector identity

$$\Rightarrow \boxed{\mathbf{L} - \frac{qK}{mc} \frac{\mathbf{r}}{r} = \text{constant}}$$

Thus,

$$\mathbf{r} \times \dot{\mathbf{r}} - \lambda \frac{\mathbf{r}}{r} = \text{const.}$$

$$\lambda = \frac{qK}{mc}$$

$$\text{or } \dot{\mathbf{r}} \times \mathbf{r} + \frac{\lambda \mathbf{r}}{r} = \text{const.} = \mathbf{l}' = \mathbf{l}$$

the last term is along \mathbf{r} , and the first term normal to \mathbf{r} . the angle between \mathbf{l}' & \mathbf{r} is a constant. $\cos \theta = \frac{\mathbf{l}' \cdot \mathbf{r}}{\mathbf{l}' \cdot \mathbf{r}} = \frac{\lambda r}{\mathbf{l}' r} = \frac{\lambda}{\mathbf{l}'} = \text{const.}$

the path is along the surface of a cone and is a geodesic

Consider components of \dot{r}

$$\dot{\underline{r}} \times \dot{\underline{r}} = \begin{pmatrix} \dot{r} & \dot{\theta} & \dot{\phi} \\ v_r & v_\theta & v_\phi \\ r & 0 & 0 \end{pmatrix} = rv_\phi \hat{\theta} - rv_\theta \hat{\phi}$$

$$\therefore \lambda \hat{r} + rv_\phi \hat{\theta} - rv_\theta \hat{\phi} = \text{const.}$$

so, the radial velocity is constant and the other two components of the speed depend inversely on $r \rightarrow$ this is a cone!

take $v_\theta = 0$ as the initial speed. Because

$$rv_\theta = \text{const.} \Rightarrow v_\theta \text{ is always zero.} \Rightarrow \theta = \text{const.} = \theta_0$$

