

PTYS 558: Wednesday, April 29, 2020

- Finish derivation relating the spatial diffusion coefficient (along the magnetic field) to the pitch-angle diffusion coefficient.
- Discuss energy change
- Particle acceleration at shocks

The Parker Transport Equation:

$$\frac{\partial f}{\partial t} = -V_{w,i} \frac{\partial f}{\partial x_i} + \frac{\partial}{\partial x_i} \kappa_{ij}^{(S)} \frac{\partial f}{\partial x_j} - V_{D,i} \frac{\partial f}{\partial x_i} - \frac{1}{3} \frac{\partial V_{w,i}}{\partial x_i} \frac{\partial f}{\partial \ln p} + Q$$

advection
diffusion
drift
energy change
source

Where the drift velocity due to the large scale curvature and gradient of the average magnetic field is:

$$\mathbf{V}_d = \frac{pcw}{3q} \nabla \times \left[\frac{\mathbf{B}}{B^2} \right]$$

And the symmetric part of the diffusion tensor is:

$$\kappa_{ij}^{(S)} = \kappa_{\perp} \delta_{ij} + (\kappa_{\parallel} - \kappa_{\perp}) \frac{B_i B_j}{B^2}$$

**The focus
of today's
lecture**

The Parker Transport Equation is valid whenever the anisotropy is small (as observed for GCRs). It is widely used and remarkably general.

A review of Monday's discussion

- We showed that magnetic fields with fluctuations on the scale of the gyroradius of the particles, cause the particles' pitch angles to vary in a manner that is consistent with "diffusion" – this is called pitch-angle diffusion.
- The pitch-angle diffusion coefficient is determined by the power spectrum of the turbulent magnetic field through quasi-linear theory
- We also showed that particles undergo a random walk in space, so the distribution can be described as a diffusion in space as well. This is the term in the Parker equation.
- Since the Parker equation is an average over pitch-angle – there is no pitch angle in this equation, we must relate the spatial diffusion to the pitch-angle diffusion separately in order to relate the spatial diffusion to the magnetic field power spectrum. This derivation started towards the end of the lecture on Monday

Particle Transport (cont.)

Particle angle diffusion & its relation to spatial diffusion
recall, we had.

$$\frac{\partial f}{\partial t} + v\mu \frac{\partial f}{\partial z} = \frac{\partial}{\partial \mu} \left(D_{\mu\mu} \frac{\partial f}{\partial \mu} \right)$$

assume

$$f = \frac{1}{2} f_0 + f_1(\mu) \quad f_0 \rightarrow \text{isotropic part}$$

$$\langle f \rangle = \int_{-1}^1 f d\mu = f_0$$

$$\int_{-1}^1 f_1 d\mu = 0$$

substitute in the p.a. diff. eq.

$$\frac{1}{2} \frac{\partial f_0}{\partial t} + \frac{\partial f_1}{\partial t} + \frac{1}{2} v\mu \frac{\partial f_0}{\partial z} + v\mu \frac{\partial f_1}{\partial z} = \frac{\partial}{\partial \mu} \left(D_{\mu\mu} \frac{\partial f_1}{\partial \mu} \right) \quad (1)$$

Integrate over μ from -1 to +1

$$\frac{\partial f_0}{\partial t} + v \int_{-1}^1 \mu \frac{\partial f_1}{\partial z} d\mu = \left[D_{\mu\mu} \frac{\partial f_1}{\partial \mu} \right]_{-1}^{+1} = 0$$

$$\boxed{\frac{\partial f_0}{\partial t} = -v \int_{-1}^1 \mu \frac{\partial f_1}{\partial z} d\mu} \quad (2)$$

multiply (2) $\times \frac{1}{2}$, subtract from (1)

$$\frac{\partial f_1}{\partial t} + \frac{1}{2} n \mu \frac{\partial f_0}{\partial z} + n \mu \frac{\partial f_1}{\partial z} = \frac{\partial}{\partial \mu} \left(D_{\mu\mu} \frac{\partial f_1}{\partial \mu} \right) + n \int_{-1}^1 \mu \frac{\partial f_1}{\partial z} d\mu$$

because T , characteristic ^{time} scale & L characteristic length scale are large compared to scattering time, τ , and mean-free path λ

$$T \gg \tau$$

$$L \gg \lambda$$

$$\text{and } f_1 \ll f_0$$

only terms 2 & 4 are important in this eq

$$\therefore \frac{1}{2} n \mu \frac{\partial f_0}{\partial z} \approx \frac{\partial}{\partial \mu} \left(D_{\mu\mu} \frac{\partial f_1}{\partial \mu} \right)$$

$$\Rightarrow \int_{-1}^{\mu} \frac{1}{2} n \mu' \frac{\partial f_0}{\partial z} d\mu' = D_{\mu\mu} \frac{\partial f_1}{\partial \mu}$$

$$\Rightarrow \frac{\nu}{4} (\mu^2 - 1) \frac{\partial f_0}{\partial z} = D_{\mu\mu} \frac{\partial f_1}{\partial \mu}$$

$$\Rightarrow \frac{\partial f_1}{\partial \mu} = \frac{v}{4} (\mu^2 - 1) \frac{1}{D_{\mu\mu}} \frac{\partial f_0}{\partial z}$$

$$f_1 = \int_{-1}^{\mu} \frac{v}{4} (\mu'^2 - 1) \frac{1}{D_{\mu\mu}(\mu')} \frac{\partial f_0}{\partial z} \mu' d\mu'$$

(Insert this into (2) (boxed eq.))

$$\frac{\partial f_0}{\partial z} = \frac{v^2}{4} \int_{-1}^1 \mu d\mu \frac{\partial}{\partial z} \left[\int_{-1}^{\mu} \frac{1 - \mu'^2}{D_{\mu\mu}} \frac{\partial f_0}{\partial z} d\mu' \right]$$

$$= \frac{\partial}{\partial z} \left(\frac{v^2}{4} \int_{-1}^1 \mu d\mu \int_{-1}^{\mu} \frac{1 - \mu'^2}{D_{\mu\mu}} d\mu' \frac{\partial f_0}{\partial z} \right)$$

$$= \frac{\partial}{\partial z} K \frac{\partial f_0}{\partial z}$$

where

$$K = \frac{v^2}{4} \int_{-1}^1 \mu d\mu \int_{-1}^{\mu} \frac{1 - \mu'^2}{D_{\mu\mu}(\mu')} d\mu'$$

Change order of integration to give

$$K = \frac{v^1}{4} \int_{-1}^1 \frac{1-\mu'^2}{D_{\mu\mu}(\mu')} d\mu' \int_{\mu}^1 \mu d\mu$$

:

$$K = \frac{v^2}{4} \int_0^1 \frac{(1-\mu^2)^2 d\mu}{D_{\mu\mu}}$$

Ead, 1974

Whinnery, 1976

recall

$$D_{\mu\mu} = \frac{\pi}{4} \Omega_0 (1-\mu^2) \frac{k_r P(k_r)}{B_0^2}$$

quasi-linear theory

Tokipiti, 1966

$$k_r = \frac{\Omega}{v_{\mu}}$$

resonant wave #

P → power spectrum

$$K = \frac{1}{3} \lambda v \leftarrow \text{particle speed}$$

↖ mean free path

Particle Transport: Energy change

recall eq. of motion of a charged particle

$$\frac{d\vec{p}}{dt} = q\vec{E} + \frac{q}{c}\vec{w} \times \vec{B}$$

$$\vec{w} \cdot \frac{d\vec{p}}{dt} = q\vec{w} \cdot \vec{E}$$

$$\frac{dT}{dt} = q\vec{w} \cdot \vec{E} \quad T = \text{Kinetic energy}$$

$$\vec{E} = -\frac{1}{c}\vec{U} \times \vec{B} \rightarrow \text{ideal MHD}$$

$$\Rightarrow \frac{dT}{dt} = -\frac{q}{c}\vec{w} \cdot (\vec{U} \times \vec{B})$$

T is in the lab frame - inertial frame.

Consider instead the frame moving w/ fluid

$$T' = T - \vec{p} \cdot \vec{U} + \frac{1}{2}mU^2$$

$$\Rightarrow \frac{dT'}{dt} = \frac{dT}{dt} - \frac{d}{dt}(\vec{p} \cdot \vec{U}) + \frac{d}{dt}\left(\frac{1}{2}mU^2\right)$$

$$\frac{dT'}{dt} = -\frac{q}{c} \underline{w} \cdot (\underline{U} \times \underline{B}) - \underline{p} \cdot \frac{d\underline{U}}{dt} - \underline{U} \cdot \frac{d\underline{p}}{dt} + m \underline{U} \cdot \frac{d\underline{U}}{dt}$$

$\underline{w} \cdot (\underline{E} + \underline{v} \times \underline{B})$

and $q \underline{E} \cdot \underline{U} = -\frac{q}{c} \underline{U} \times \underline{B} \cdot \underline{U} = 0$

$$= -\frac{q}{c} \underline{w} \cdot (\underline{U} \times \underline{B}) - \underline{p} \cdot \frac{d\underline{U}}{dt} - \frac{q}{c} \underline{U} \cdot (\underline{w} \times \underline{B}) + m \underline{U} \cdot \frac{d\underline{U}}{dt}$$



these cancel after implementing a vector identity

$$= -(\underline{p} - m \underline{U}) \cdot \frac{d\underline{U}}{dt}$$

$$\frac{dT'}{dt} = -\underline{p}' \cdot \frac{d\underline{U}}{dt}$$

$\frac{d}{dt} \rightarrow$ along particle trajectory

$$= -\underline{p}' \cdot \left(\frac{\partial \underline{U}}{\partial t} + \underline{w} \cdot \nabla \underline{U} \right) \leftarrow \text{use advective derivative}$$

Consider a collection of particles, distributed isotropically, then

$$\frac{d\langle T' \rangle}{dt} = \left\langle - \rho' \left(\frac{\partial \underline{u}}{\partial t} + \underline{w}' \cdot \nabla \underline{u} \right) \right\rangle$$

avg. over isotropic dist.

$$= - \left\langle \rho' (\underline{w}' \cdot \nabla \underline{u}) \right\rangle$$

assume \underline{u} varies slowly in time compared to scattering time

$$= - \left\langle \rho'_i w'_j \partial_j u_i \right\rangle$$

under int.

$$= - \left\langle \rho'_i w'_j \right\rangle \partial_j u_i$$

assume \underline{u} varies in space slowly compared to λ (mean free path)

$$\underbrace{\left\langle \rho'_i w'_j \right\rangle}_{\frac{1}{3} \rho' w' \delta_{ij}}$$

Kronecker delta

$$= - \frac{1}{3} \rho' w' \nabla \cdot \underline{u}$$

recall $dT'/dt = w' dp'/dt$

$$\Rightarrow \boxed{\frac{dp'}{dt} = - \frac{1}{3} \rho' \nabla \cdot \underline{u}}$$

energy change of an isotropic dist. of particles

Cosmic-Ray Transport Equation

$$\frac{\partial f}{\partial t} + \underline{v} \cdot \nabla f = \nabla \cdot (\underline{\kappa} \cdot \nabla f) + \frac{1}{3} \nabla \cdot \underline{v} \frac{\partial f}{\partial \ln p} + S - L$$

↑
↑
↑
↑

advection
of fluid
spatial
diffusion
and drifts
energy
change
Sources
losses

where,

$$K_{ij} = K_{\perp} \delta_{ij} + (K_{\parallel} - K_{\perp}) \frac{B_i B_j}{B^2} + \epsilon_{ijk} K_A \frac{B_k}{B}$$

K_{\perp} = cross-field diffusion coeff.

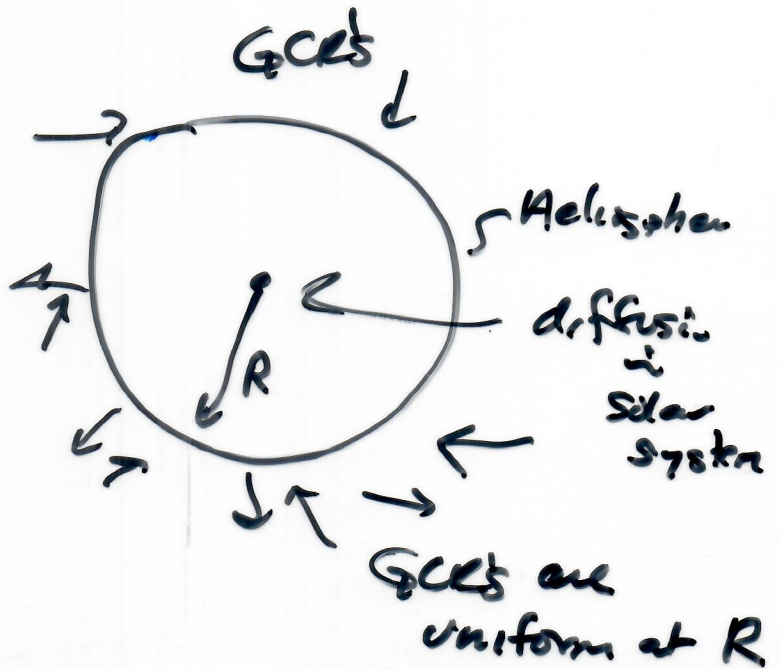
K_{\parallel} = parallel " " (previous discussion)

K_A = anti-symmetric component

→ this is where drifts are!

GCR modulation

Consider the spherical coord. rep. of the CR transport eq. (Parker equation)
in steady state



$$\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial f}{\partial r} \right) - v \frac{\partial f}{\partial r} - \frac{2U}{3r} \frac{\partial f}{\partial p} = 0$$

this can be written

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left[r^2 \left(\kappa \frac{\partial f}{\partial r} - v f \right) \right] + \frac{2U}{3r p^2} \frac{\partial}{\partial p} (p^3 f) = 0$$

if one ignores the last term, we find

$$f(r, p) = f(R, p) e^{-\int_r^R \frac{v dr'}{\kappa(r', p)}}$$

commonly used "model" (not a good one, tho)

Cosmic Rays on Earth Since 1964



In recent years, cosmic rays have been unusually strong, reaching a Space Age maximum in 2009-2010.

**Space Age Maximum
(2009-2010)**

**Peaking Again
(2018-2019)**



